

ICRC 2005
Rapporteur OG 1
 γ -ray
Science and Techniques

Rene A. Ong
University of California, Los Angeles

SESSIONS & PAPERS

Session	Topics	Oral	Poster	Total
OG 2.1	Diffuse X-rays, γ -rays	10	10	20
OG 2.2	Galactic Sources (SNRs, pulsars, etc.)	29	21	50
OG 2.3	Extragalactic Sources (AGN, clusters, etc.)	23	18	41
OG 2.4	Gamma-Ray Bursts	6	7	13
OG 2.7	New expts. & instrumentation	33	44	77
Total		101	100	201

Can choose:

- Say “nothing about everything”.
- Say “something about something”. ← try this.

Caveats & Acknowledgements

- **Large number of papers ... all received quite late ... so this summary:**
 - concentrates on highlights from this meeting,
 - is subjective and condensed.

Write-up will contain more information.

- **Relatively little from outside VHE γ -rays, and mostly observational results.**
- **Papers (results) are referenced by presenting author.**
- **Considerable thanks to:**
 - Organizers.
 - Experimental groups with results (5 or more papers):
AGRO-YBJ, CANGAROO, HESS, MAGIC, Milagro,
STACEE, TACTIC, Tibet, VERITAS.

OUTLINE

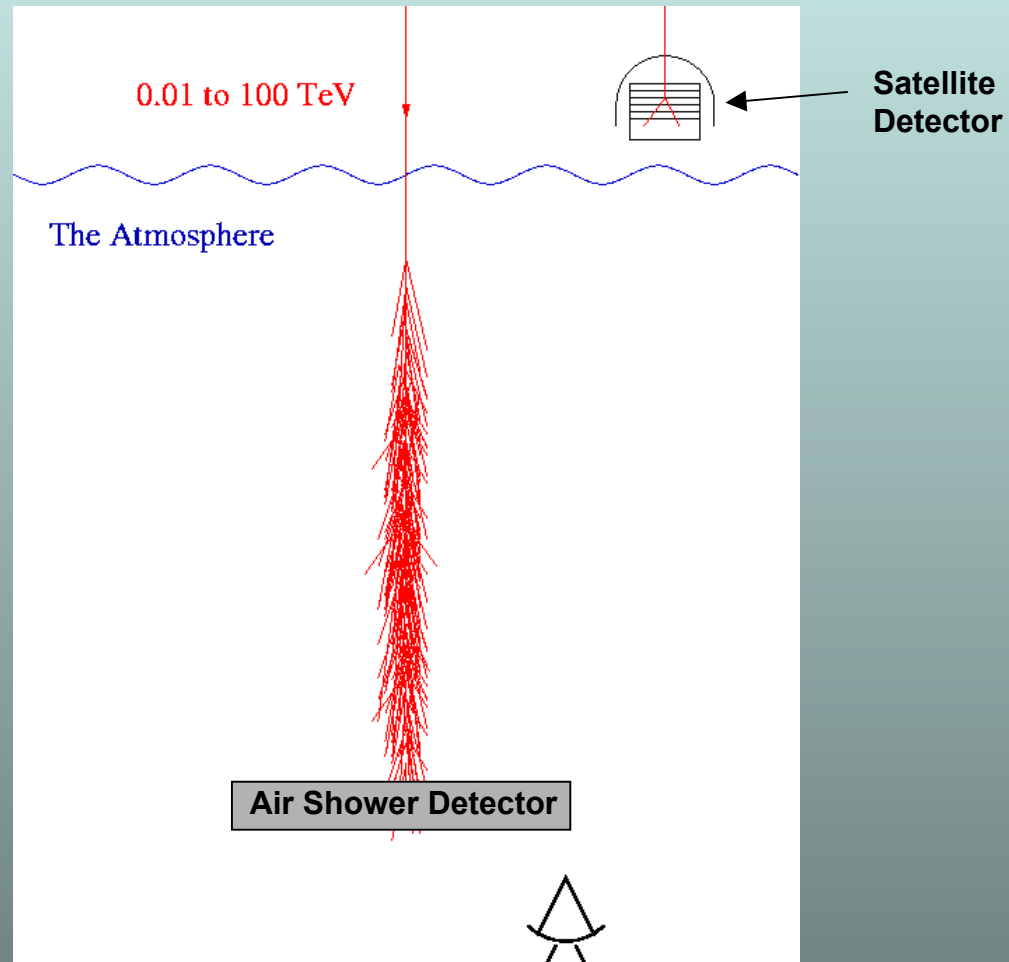
- **Instruments and Techniques**
- **Scientific Highlights**
- **Source Count & Map**
 - Where are we relative to 1995 (Hillas), 2003 (Mori)?
- **Science Results**
 - Arranged session by session.
- **New Experiments and Instrumentation**
- **Conclusions**

Instruments (Past and Present)

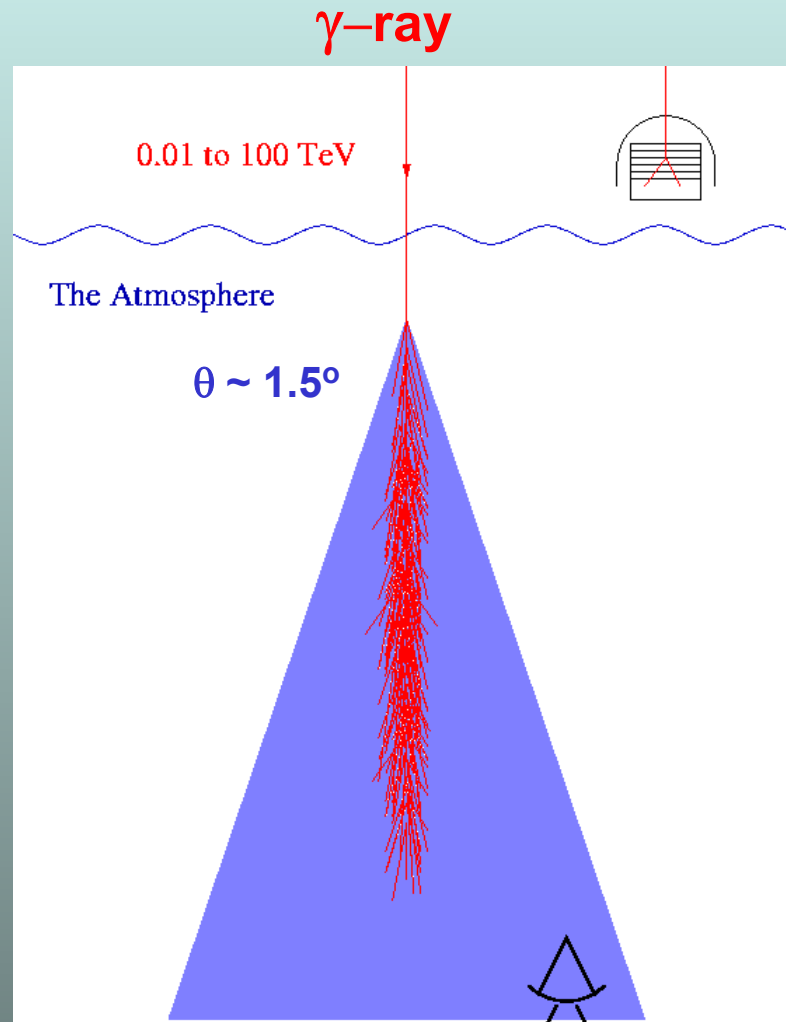
- **X-ray:**
RXTE, ASCA, Chandra, XMM-Newton
- **Gamma-ray (Satellite, 100 keV-20 GeV):**
EGRET, INTEGRAL*, Swift
- **Gamma-ray (Cherenkov, $E > 50$ GeV):**
Whipple, PACT, STACEE, TACTIC, Shalou, CANGAROO, HESS, MAGIC, VERITAS
- **Gamma-ray (Air Shower, $E > 100$ GeV):**
Tibet, Milagro, GRAPES-3, ARGO-YBJ

*Science summary by Produit (OG 2.2).

Air Shower Detectors



Cherenkov Telescopes

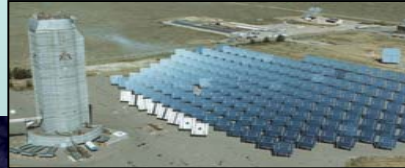


VHE Experimental World

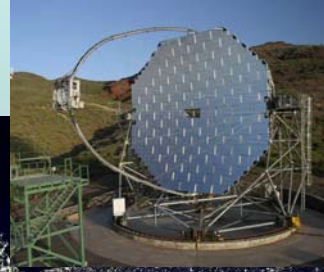
MILAGRO



STACEE



MAGIC



TIBET



MILAGRO

STACEE

MAGIC

TIBET
ARGO-YBJ

TACTIC

PACT

GRAPES

VERITAS

VERITAS

TACTIC

HESS

CANGAROO III

HESS



CANGAROO



Scientific Highlights

I. Discovery of many new Galactic sources by HESS:

- [HESS GP Survey](#) & targeted observations.

II. Detailed studies of Galactic sources:

- Precision measurements (spectra, images, etc.).
- Theoretical models and understanding.

III. Discovery of 4 new AGN by HESS and MAGIC.

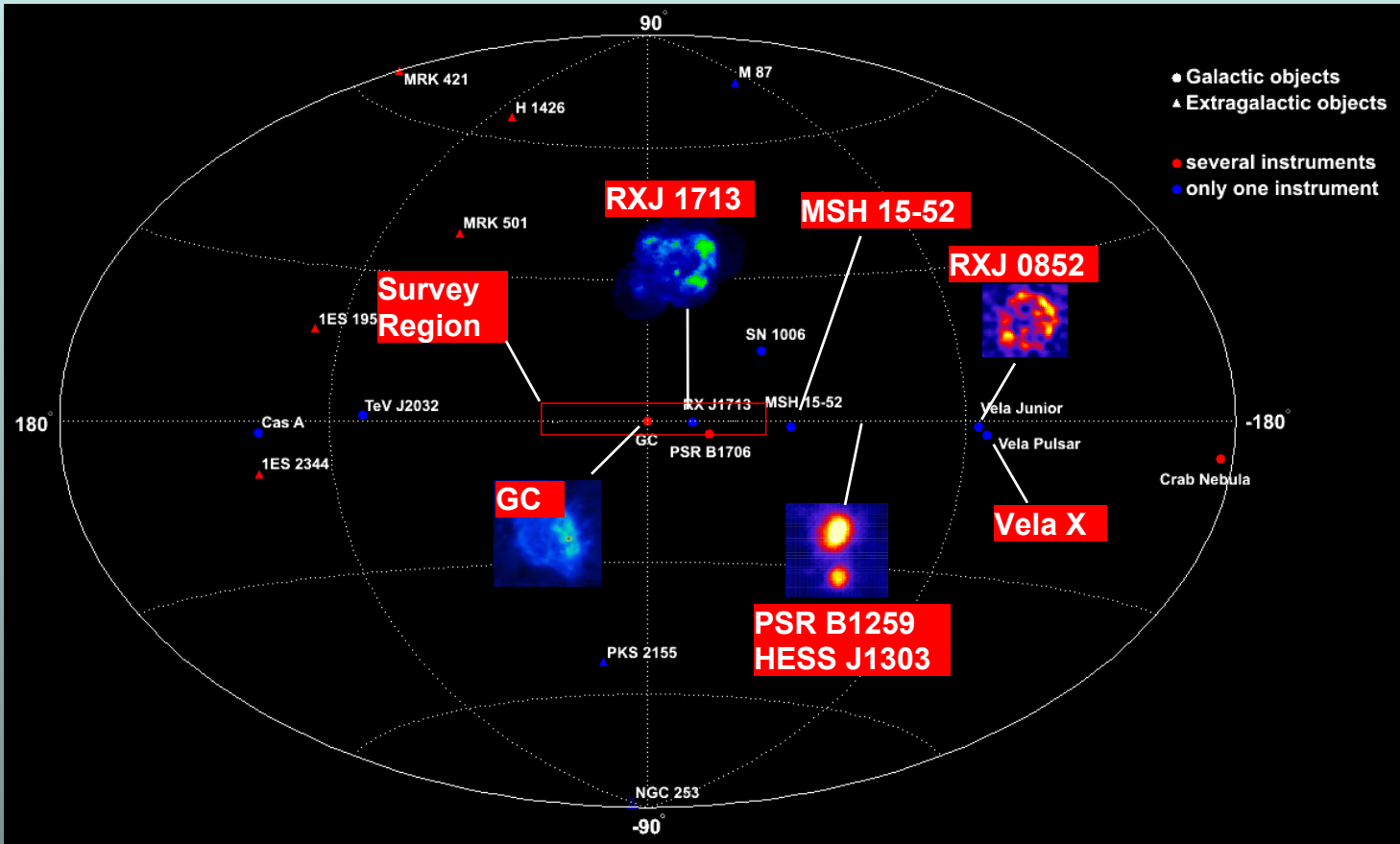
- Measurements of AGN properties and multi- λ studies.

IV. Other new observations:

- Detection of diffuse emission in Cygnus by Milagro.
- Searches for new galactic and extragalactic sources.

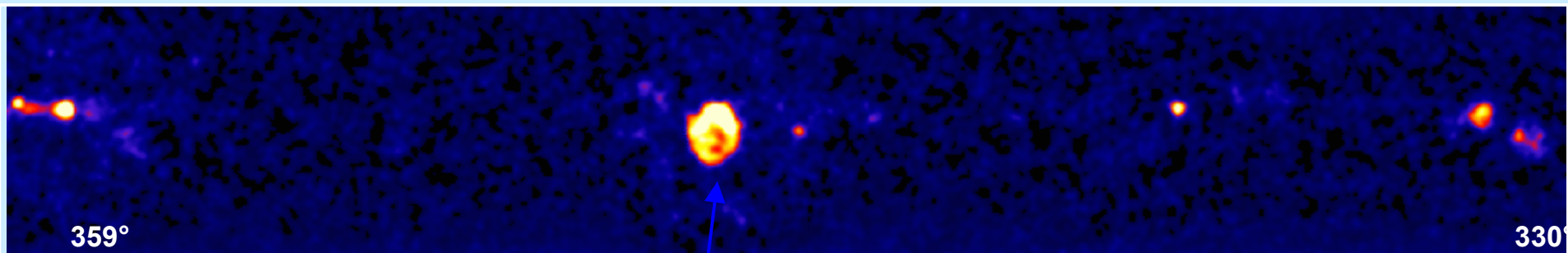
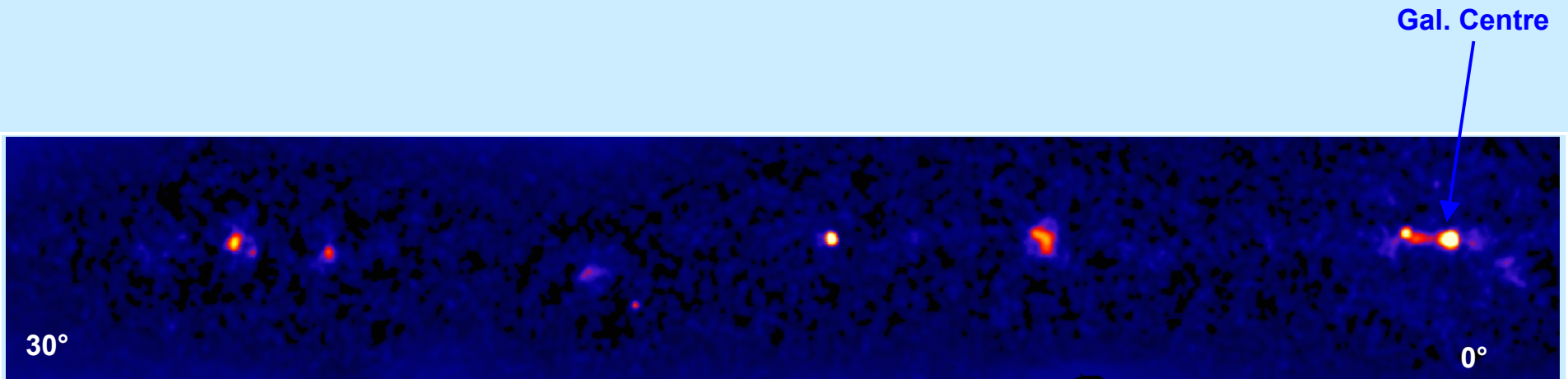
HESS Galactic Plane Survey

[Funk, Lemièrre]



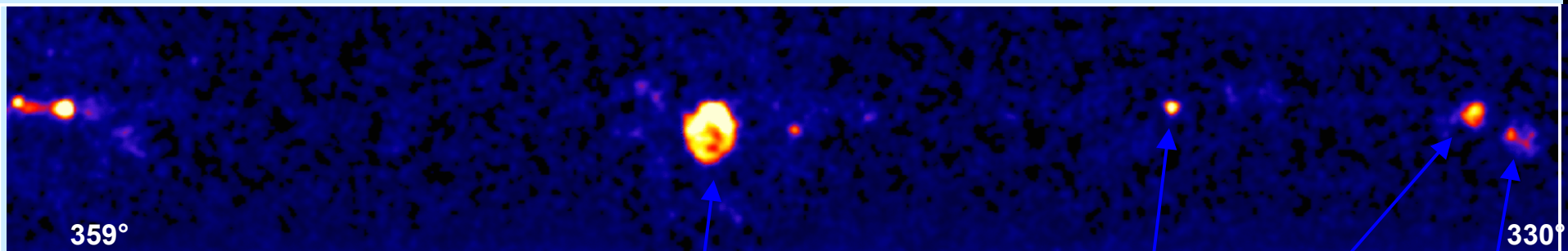
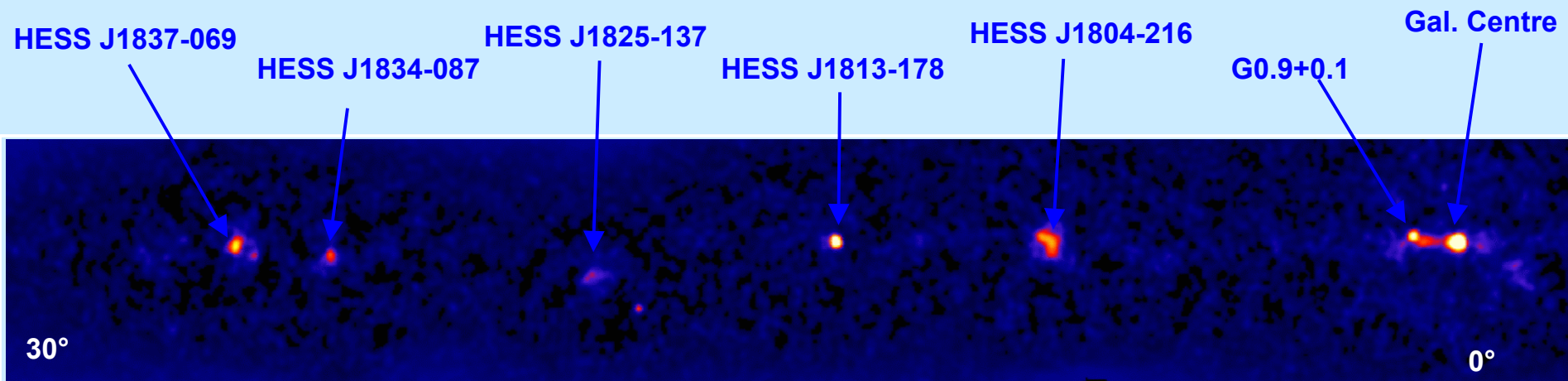
- 60° in longitude, $\pm 3^\circ$ in latitude
- 112 hrs scanning + follow-up observations

HESS GP Survey



RX J1713.7-3946

HESS Survey: New sources



Sources > 6 sigma (9 new, 11 total)

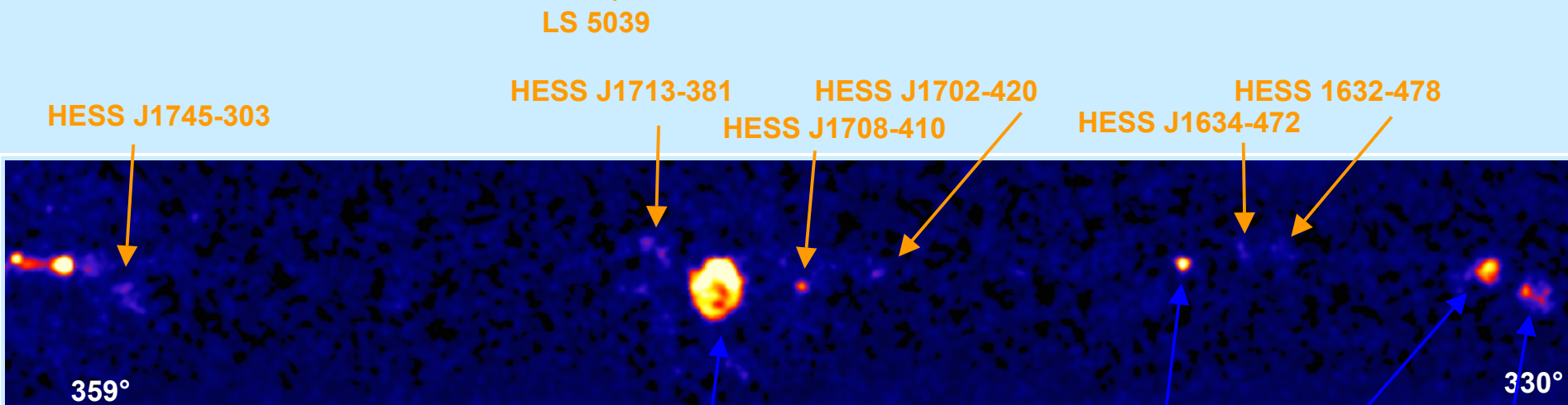
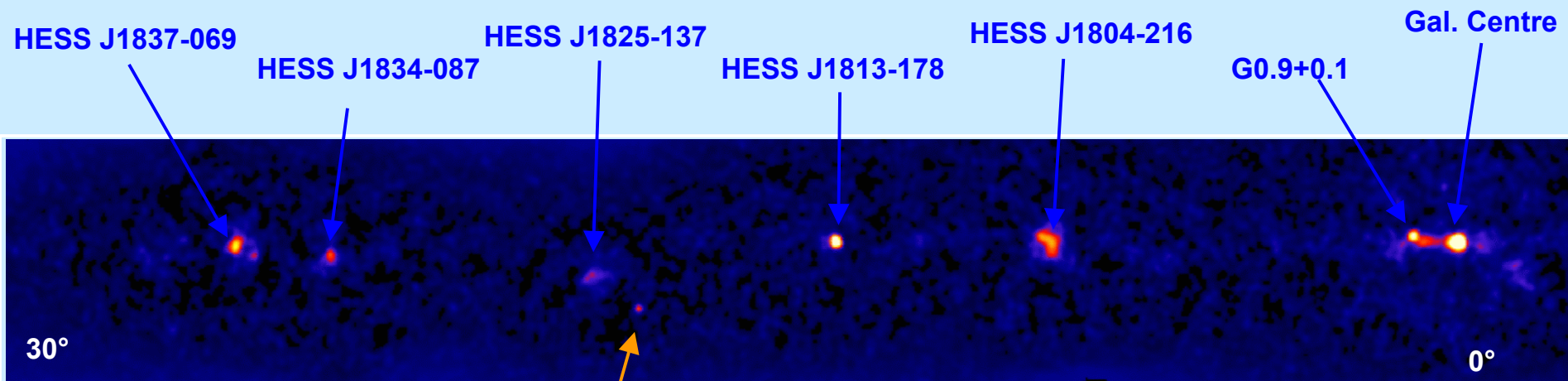
RX J1713.7-3946

HESS J1640-485

HESS J1614-518

HESS J1616-508

HESS Survey: New Sources



Sources > 6 sigma (9 new, 11 total)

Sources > 4 sigma (7 new)

RX J1713.7-3946

HESS J1640-485

HESS J1614-518

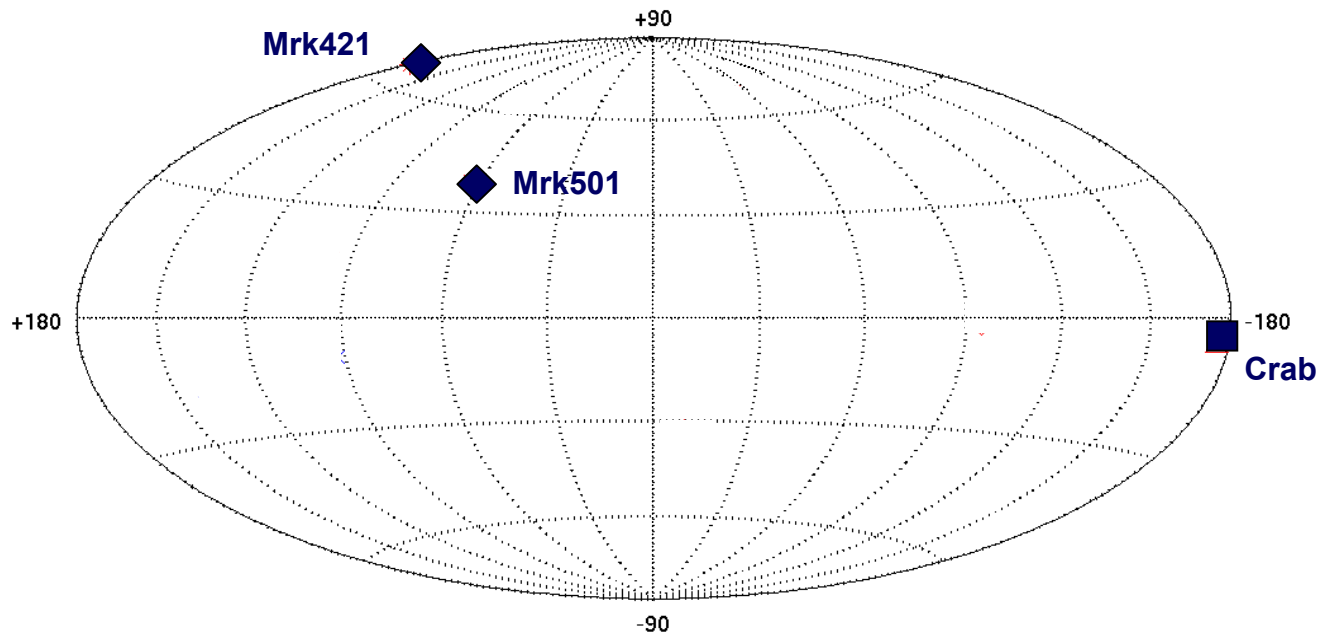
HESS J1616-508

HESS GP Source Associations

Source	Possible Association	Comments
G0.9+0.1	PWN/SNR	Firm Association
LS 5039	Micro-Quasar	Firm Association
J1616-508	PWN ?	Tentative
J1614-518		Unid ?
J1640-465	SNR ? (G338.3)	Tentative
J1804-216		Unid ?
J1813-178	SNR ? (G12.8-0.0, Integral)	Tentative
J1825-137	PWN ? (PSR J1826)	Tentative
J1834-087	SNR ? (G23.3)	Tentative
J1837-069	(Integral)	Unid ?

Study spatial association and correlation with HI/CO maps [Lemiere].

The VHE Sky - 1995



■ Pulsar Nebula

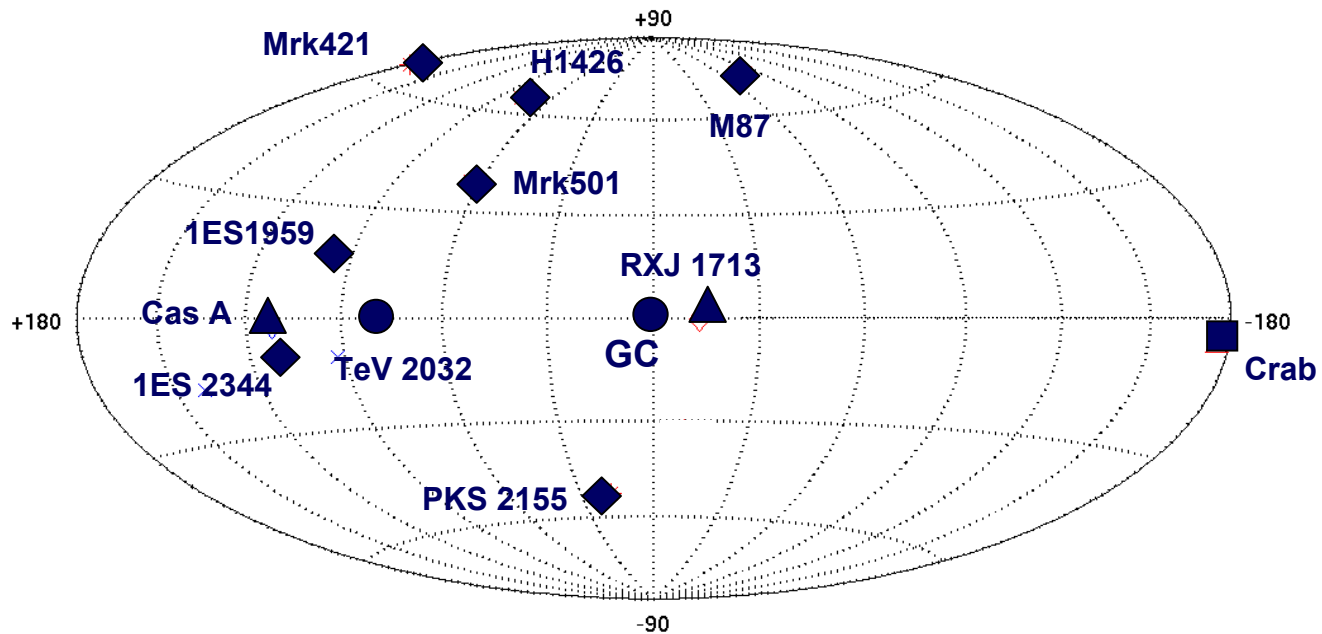
◆ AGN

▲ SNR

● Other, UNID

R.A.Ong
Aug 2005

The VHE Sky - 2003



■ Pulsar Nebula

◆ AGN

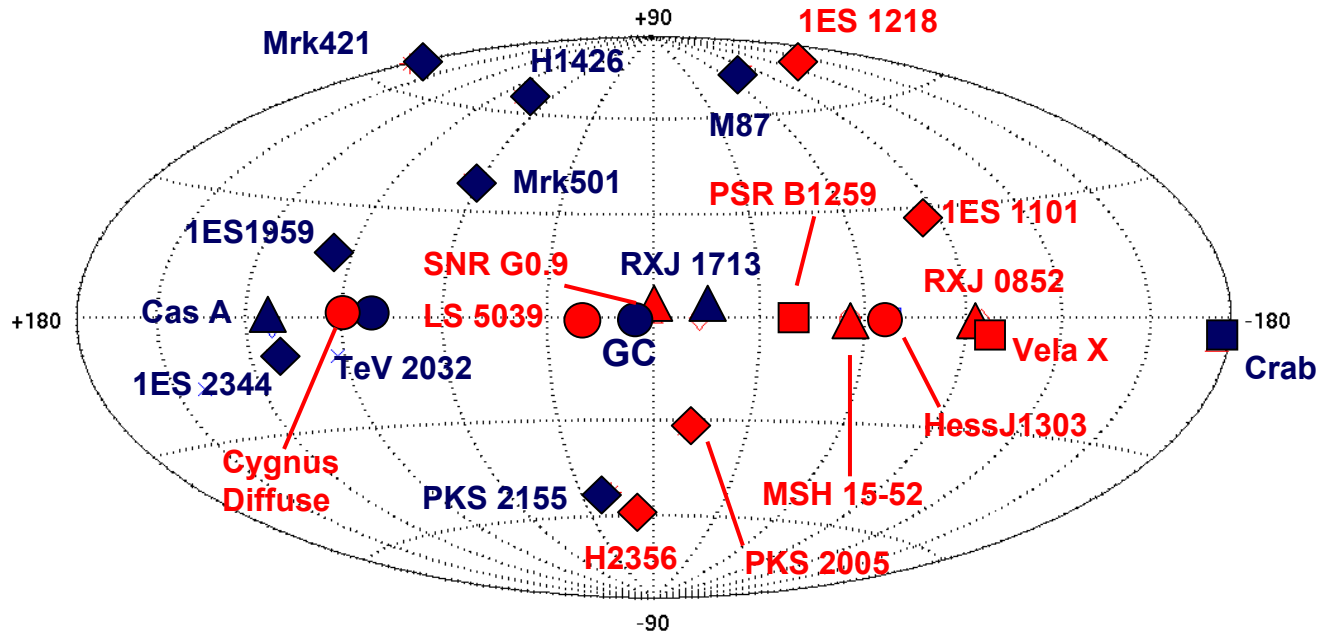
▲ SNR

● Other, UNID

R.A.Ong
Aug 2005

The VHE Sky - 2005

+ 8-15 add. sources
in galactic plane.



- Pulsar Nebula
- ◆ AGN
- ▲ SNR
- Other, UNID

R.A.Ong
Aug 2005

Source Counts

Source Type*	2003	2005
Pulsar Wind Nebula (e.g. Crab, MSH 15-52 ...)	1	6
Supernova Remnants (e.g. Cas-A, RXJ 1713 ...)	2	6
Binary Pulsar (B1259-63)	0	1
Micro-quasar (LS 5039)	0	1
Diffuse (Cygnus region)	0	1
AGN (e.g. Mkn 421, PKS 2155 ...)	7	11
Unidentified	2	6
TOTAL	12	32

* Includes likely associations of HESS unid sources.

→ **Explosion in the number of VHE sources.**

SCIENCE RESULTS

Session by Session

OG 2.2 – Galactic Source

OG 2.1 – Diffuse Sources

OG 2.4 – Gamma-Ray Bursts

OG 2.3 – Extragalactic Sources

OG 2.2

Galactic Sources

OG 2.2: Galactic Sources

1a. Pulsar Nebulae (PWN)

- Pulsar wind powered. Known e^- accelerators (synch and IC).
- “Standard Candle” (Crab).

Known (2003): Crab

HESS [Masterson], MAGIC [Wagner],
STACEE [Kildea], TACTIC [DHAR],
CANGAROO [Nakamori]

New Sources: MSH 15-52

HESS [Khelifi]

G0.9+0.1

“ “

Vela X

“ “

J1616-508

HESS association [Funk]

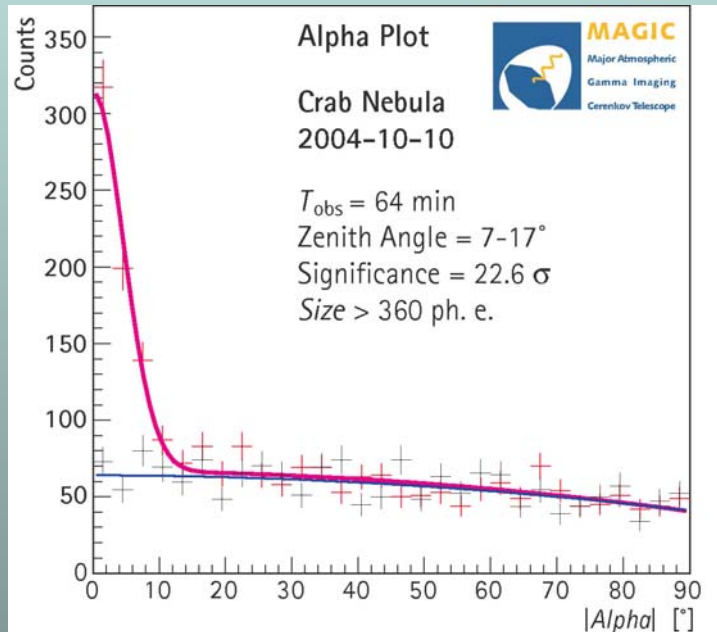
J1825-137

HESS association [de Jager]

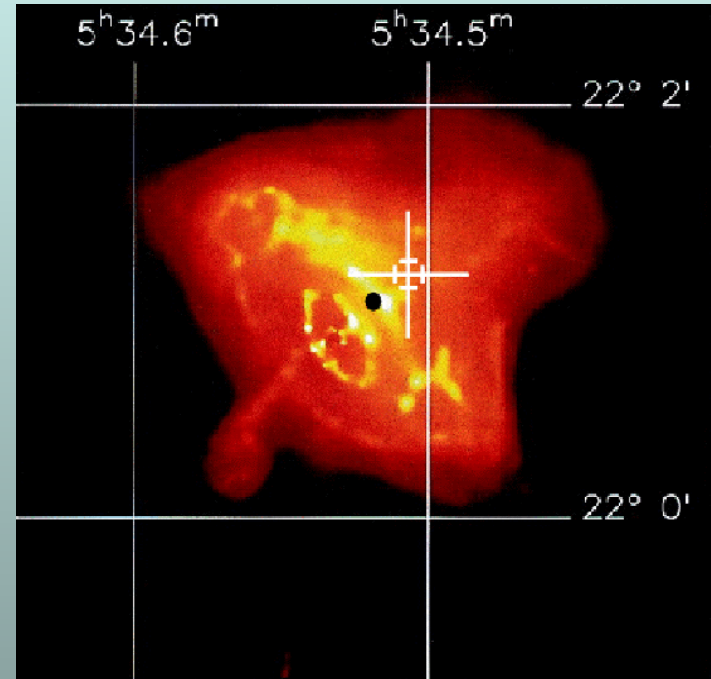
6 established VHE pulsar nebulae.

OG 2.2: Galactic Sources

CRAB



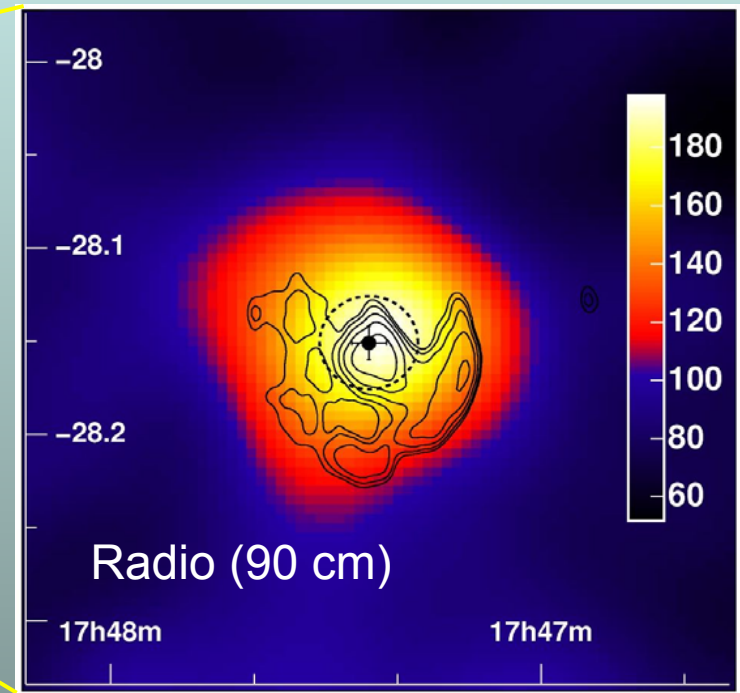
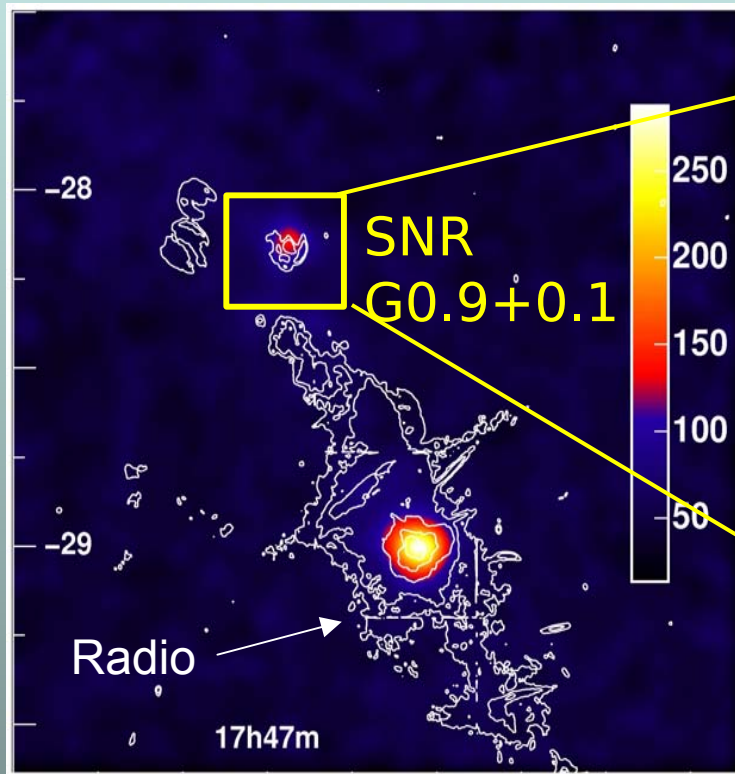
MAGIC [Wagner]:
 $22.6 \sigma/\text{hour}$
Spectrum from 300-3000 GeV



HESS [Masterson]:
Consistent with pt. source,
< $2'$ and X-ray center.

OG 2.2: Galactic Sources

G0.9+0.1



HESS [Khelifi]:
2004: 50 hrs, 4 Telescopes
13 σ

Consistent with pt. source
at position of pulsar wind.
(but not SN shell).
2% Crab flux (50% luminosity).

OG 2.2: Galactic Sources

1b. Pulsars

- 8 pulsars detected by EGRET, but none at VHE.
- Pulsed emission mechanism not known.

Limits:	Crab	HESS [Konopelko], MAGIC [Lopez], STACEE [Kildea], PACT [Singh]
	Vela, 1706-44	HESS [Konopelko]
	msec pulsars	MAGIC [Ona-Wilhelmi]
Unconfirmed:	Geminga	Shalun [Sinitsyna] PACT [Vishwanath]

→ No convincing pulsar detections at VHE.

OG 2.2: Galactic Sources

2. Supernova Remnants (SNRs)

- Origin of CR's ? Based on energetics & spectrum.
- Known to accelerate TeV electrons (synchrotron).
- 2003: situation is unclear (especially SN1006).

Known (2003):	RXJ 1713 Cas-A	HESS [Berge]
New Sources:	RXJ 0852 J1813-178 J1834-087 J1640-465	CANGAROO [Tanimori], HESS [Komin] HESS [Funk], MAGIC [Bartko] HESS association [Funk] HESS association [Funk]
Limits:	SN 1006 5 SNRs Monogem Loop G40.5-0.5	CANGAROO [Tanimori], HESS [Rowell] HESS [Rowell] TIBET [Saito] TIBET [Zhang]

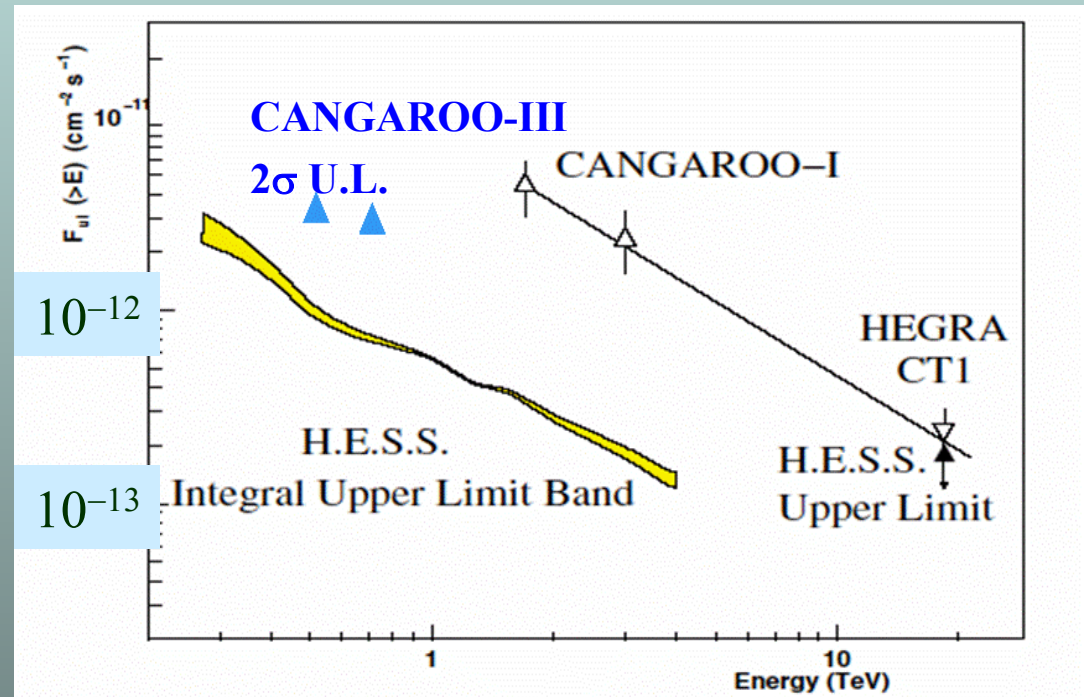
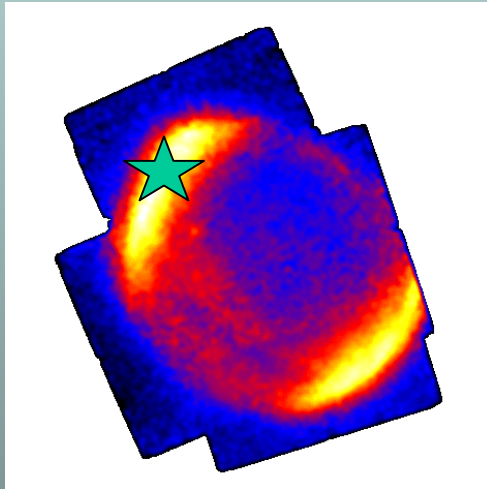
OG 2.2: Galactic Sources

SN 1006

Earlier detection by CANGAROO-I from NE rim. Also HEGRA CT1.

2004: HESS Upper limit.

New data from CANGAROO-III [Tanimori].
~ 25 hrs, 3 telescopes

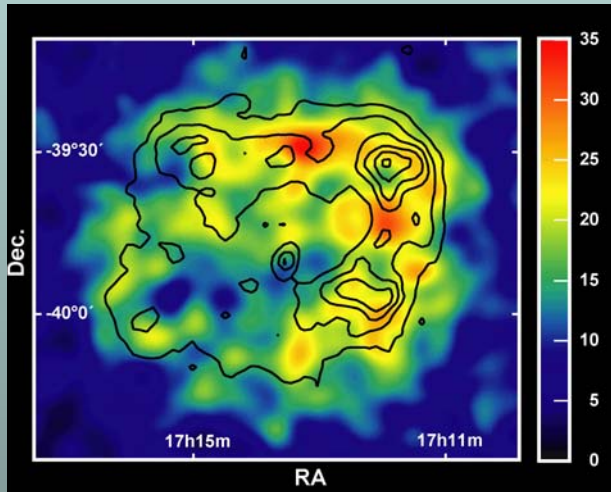


→ SN1006 not detected at previous levels.

OG 2.2: Galactic Sources

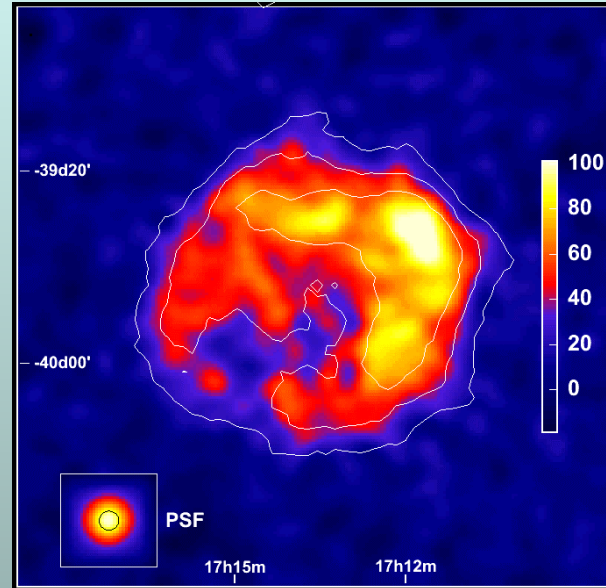
RXJ 1713-394

2003: CANGAROO
detection $\sim 7\sigma$.



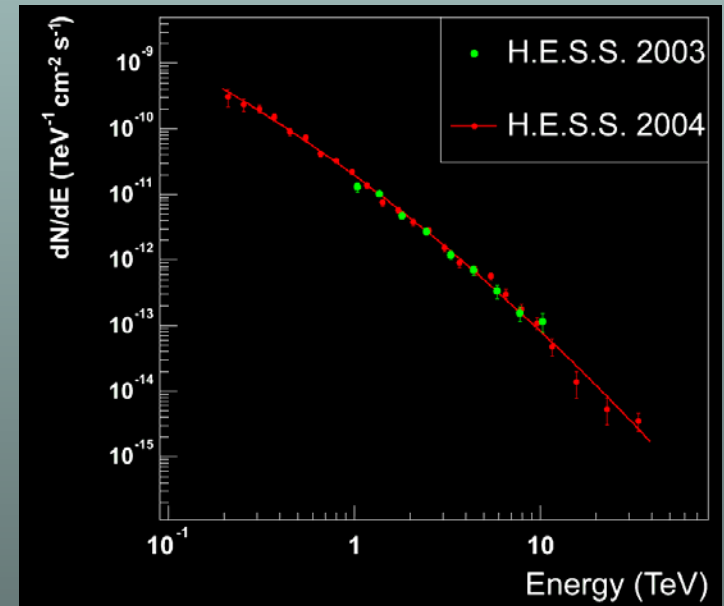
HESS [Berge]:
2003 data: 2 tels,

Hard spectrum $\Gamma \sim 2$
Not a simple power-law.



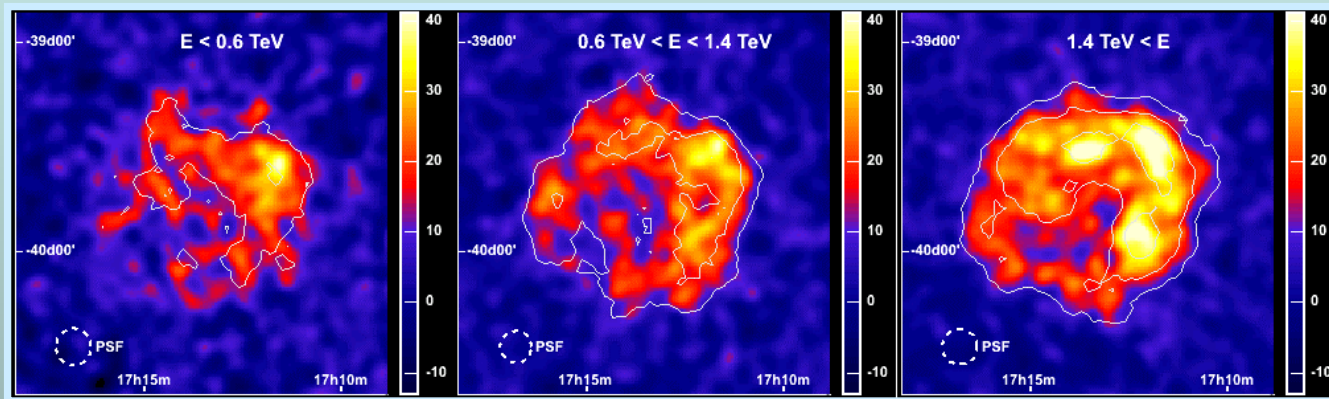
HESS [Berge]:
2004 data:
4 telescopes
33 hrs
 $> 40\sigma$

Shell structure
clearly resolved.



OG 2.2: Galactic Sources

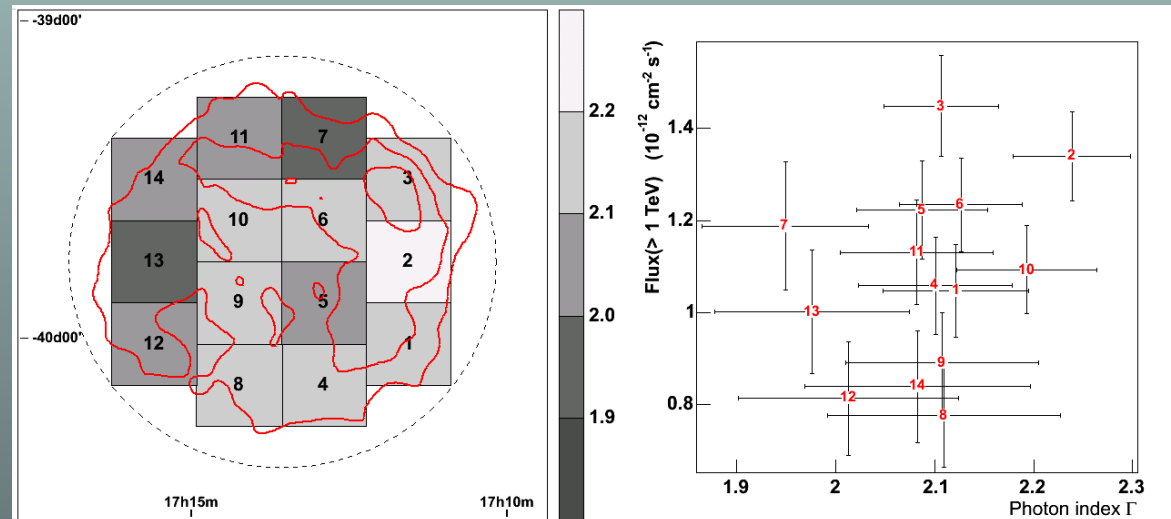
RXJ 1713-394



HESS: Morphology does not change with energy.

Spatially resolved spectra:

Spectra do not appreciably change over remnant !



We are now confident of VHE emission from numerous shell-type SNRs.

We have much more precise VHE data on SNRs than previously.

OG 2.2: Galactic Sources

SNR Theory

- SNRs clearly accelerate particles to > 10 TeV. Are they electrons or protons?
- Detailed diffusive shock models designed to explain broad-band emission from SNRs.

Ksenofontov:

Non-linear model for SN1006, key importance of gas density.

Dominant γ -ray emission from π^0 decays.

HESS result requires $N_H < 0.1 \text{ cm}^{-3}$.

Ellison:

CR ions efficiently accelerated to beyond knee. Large compression ratio required – evidence in Kepler, Tycho.

OG 2.2: Galactic Sources

SNR Theory

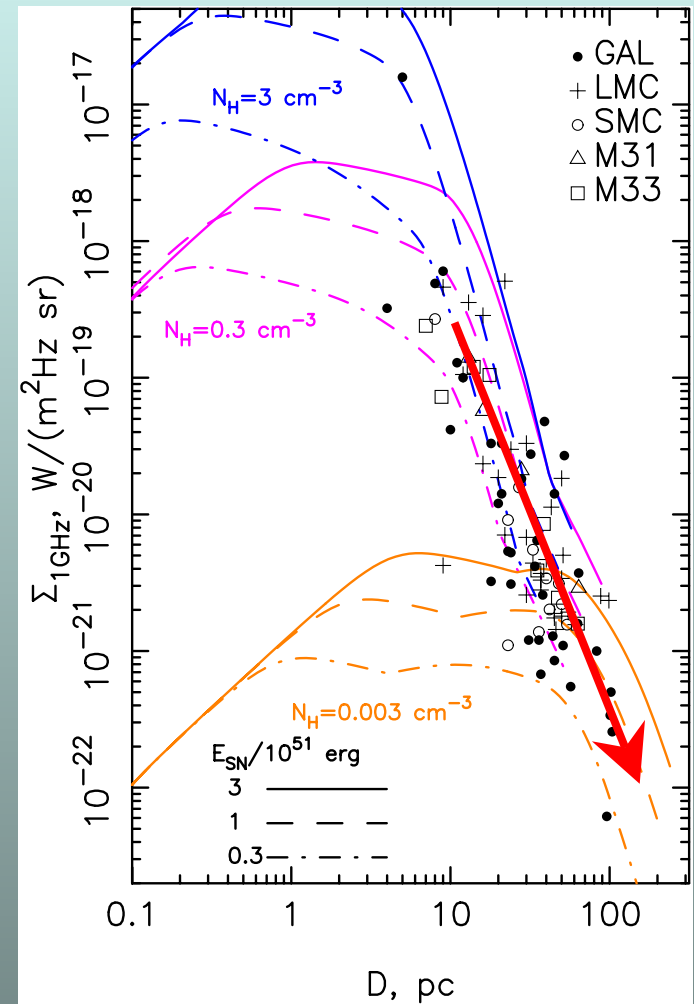
Berezkho:

Evidence for B field amplification in young SNR – steep synchr. spectra and sharp X-ray features.

Surface brightness fall rapidly shell diameter – in Sedov phase?

Q: Are SNRs the origin of HE CR's ?

e^{\pm} could be source of TeV γ -ray emission in SNRs, but would require low B fields – not generally supported by data and models.

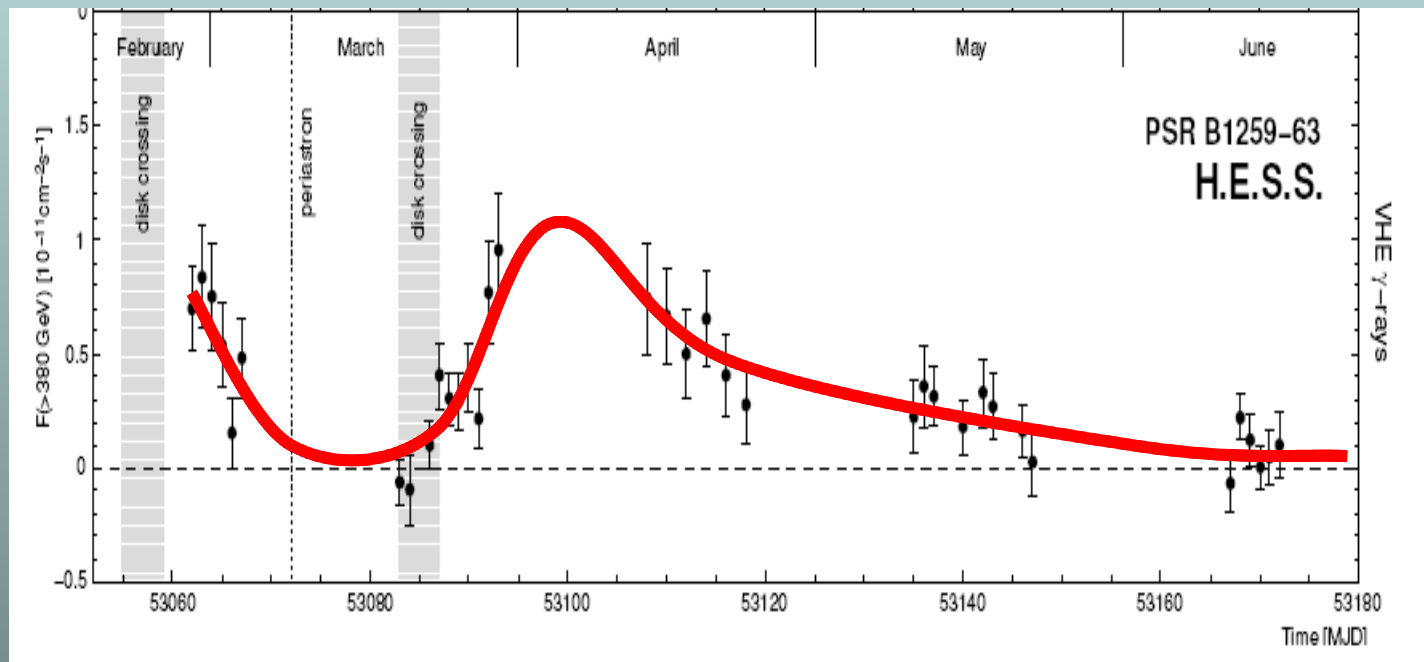


OG 2.2: Galactic Sources

C. Binary Pulsar B1259-63

HESS [Schlenker]

Theory [Kirk]



Variable light curve on daily timescale.
Complex interaction of pulsar wind and disk.
→ First variable galactic source !

OG 2.2: Galactic Sources

D. Unidentified Sources

Known (2003):

TeV 2032

Galactic Center

HESS [Rolland], MAGIC [Bartko]

New Sources:

J1303-631

HESS [Raue]

J1614-518

HESS [Funk]

J1804-216,

“ “

J1804-216

“ “

Unconfirmed:

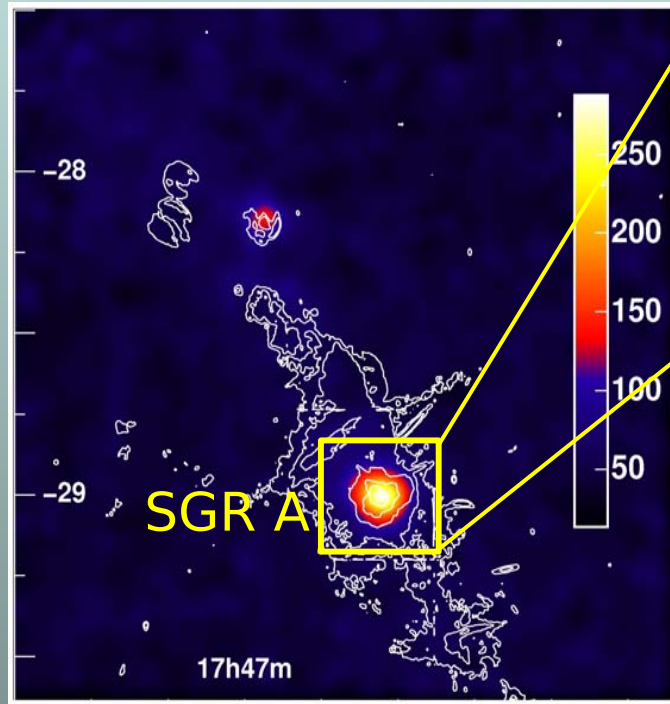
3EG J0520+2556

Milagro [Dingus]

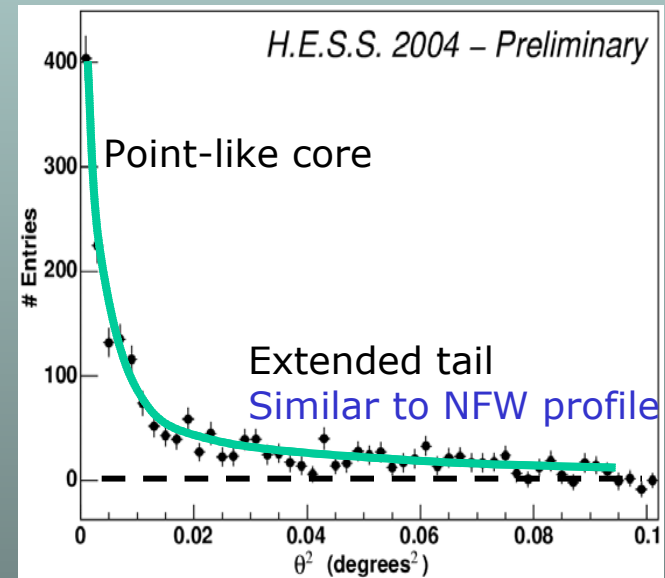
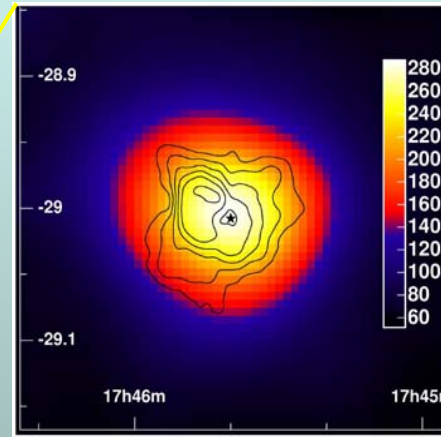
Q: Is there are new population of “dark” TeV emitters ??

OG 2.2: Galactic Sources

Galactic Center



HESS [Rolland]



Consistent with SGR A* to 6" and slightly extended.

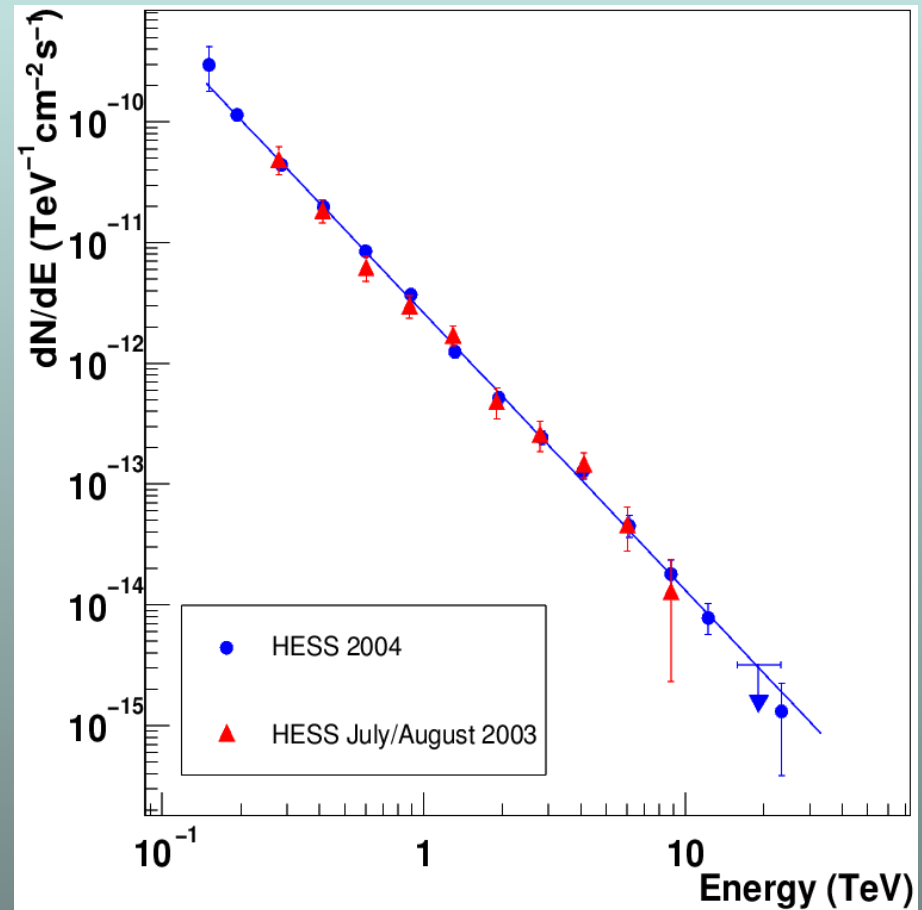
OG 2.2: Galactic Sources

Galactic Center

HESS Spectrum

- Unbroken power-law.
- Hard spectrum $\Gamma = 2.2$.
- No evidence for variability on a variety of time scales.

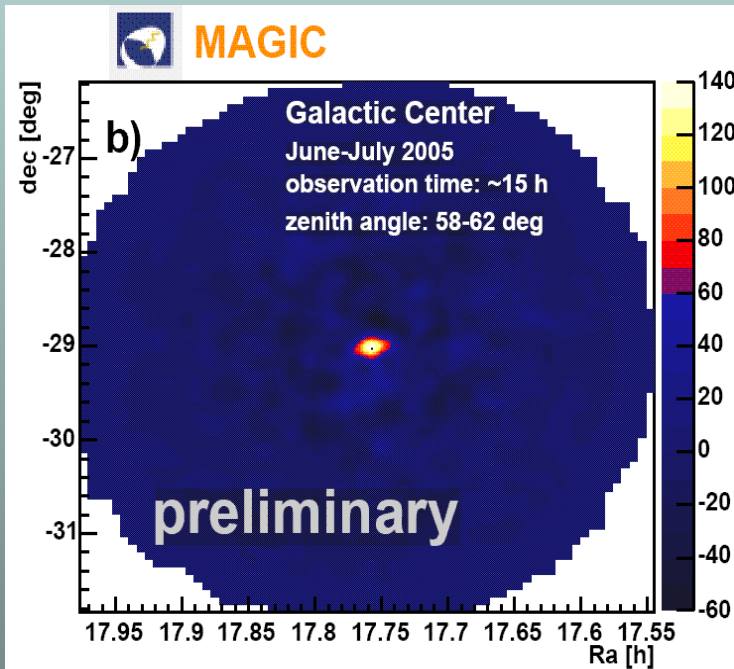
Very unlikely to be dark matter [Ripken]. Presence of strong γ -ray source makes this region difficult area to search for DM.



OG 2.2: Galactic Sources

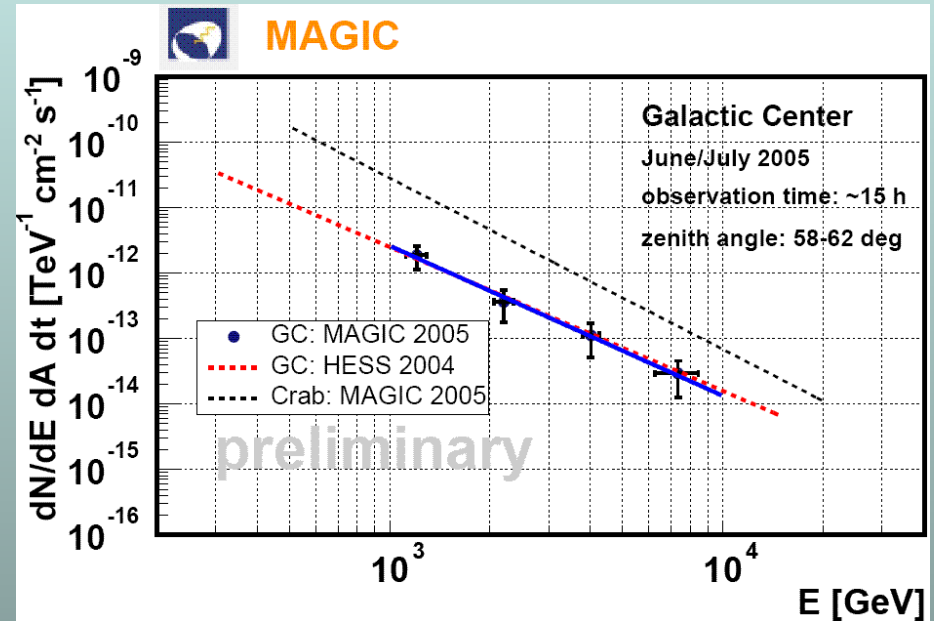
Galactic Center

Also seen clearly by MAGIC.



MAGIC (Bartko)

Spectrum



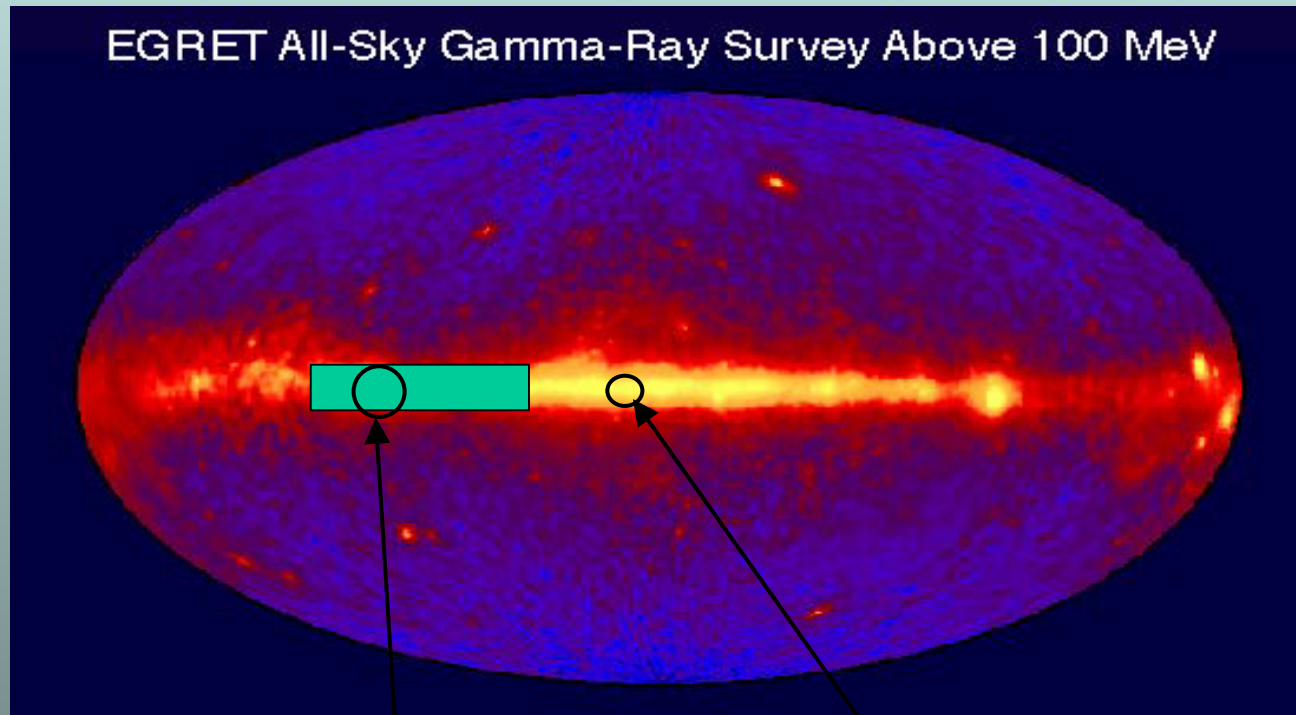
Good agreement between HESS and MAGIC.

OG 2.1

Diffuse γ -ray Sources

OG 2.1: Diffuse γ -ray Sources

Galactic Diffuse Emission



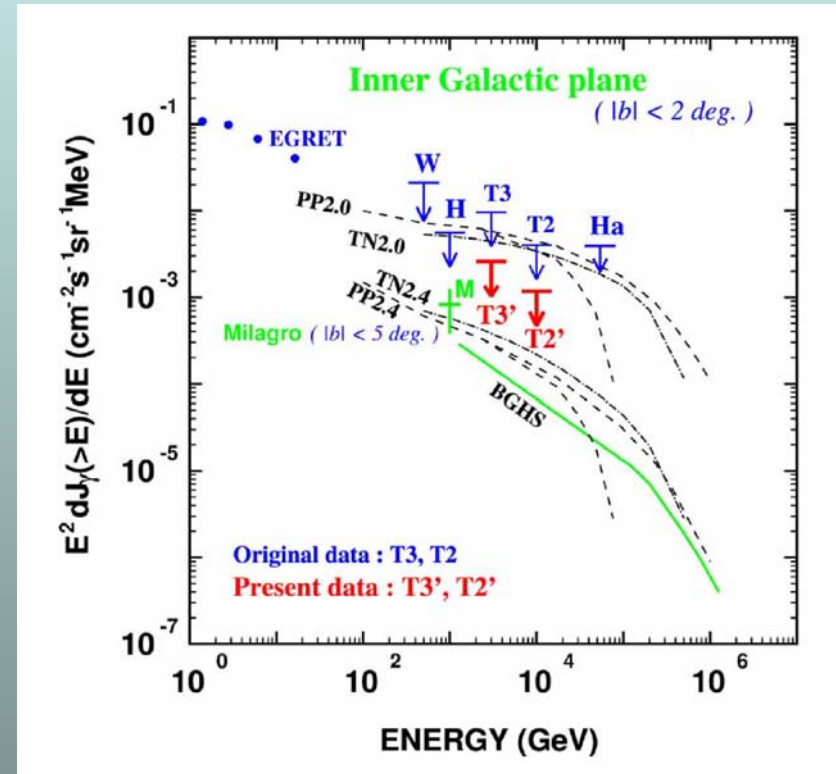
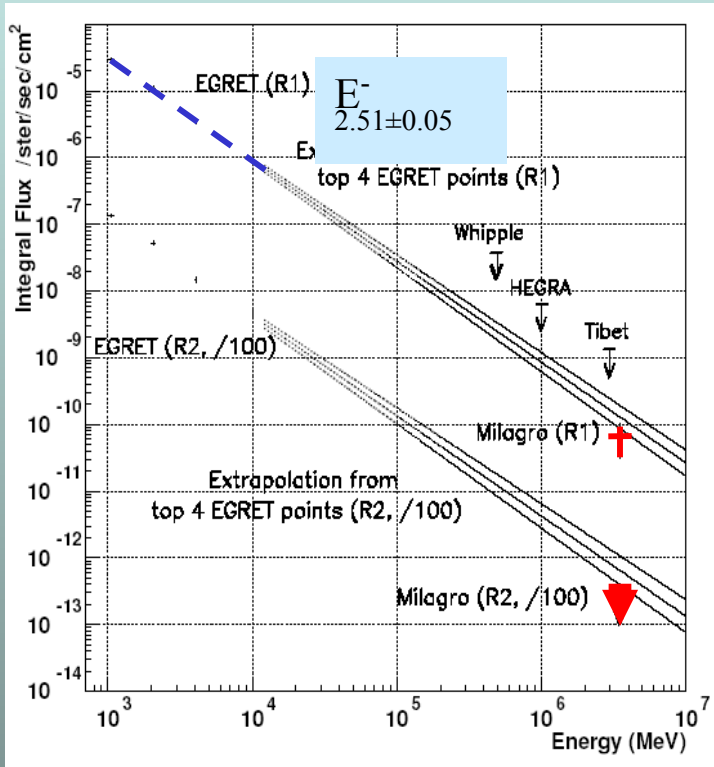
Gal. Plane
Emission
(Milagro)

Cygnus
Region
(Milagro)

Galactic Center
Region (HESS)

OG 2.1: Diffuse γ -ray Sources

1. Galactic Plane



Milagro [Sinnis]

~ 4.5 σ detection in 3 yr data set.

Inner region:

40-100° in longitude, $\pm 5^\circ$ in latitude

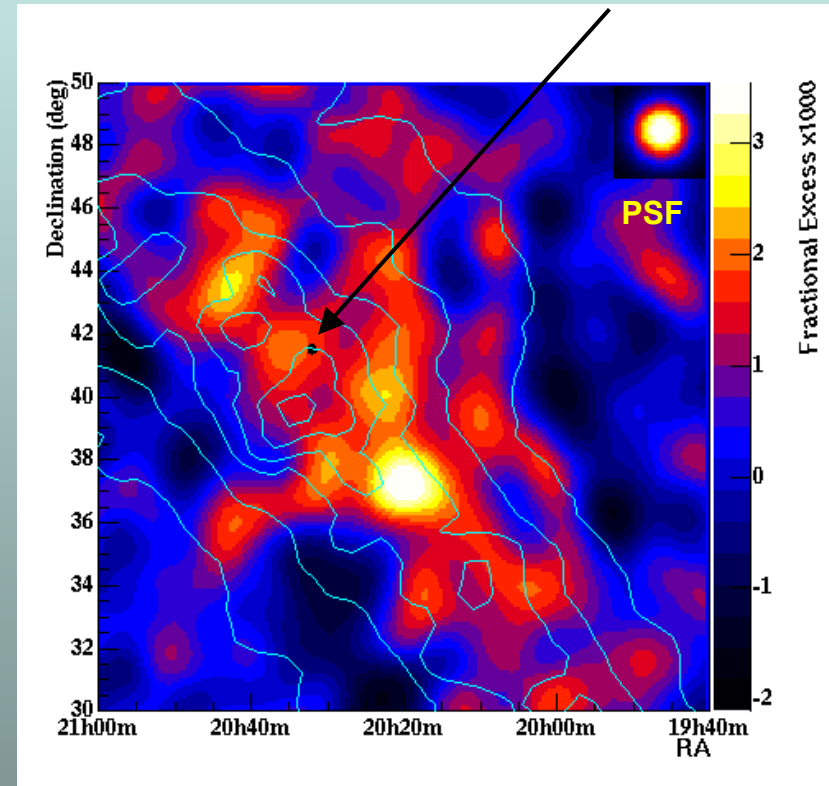
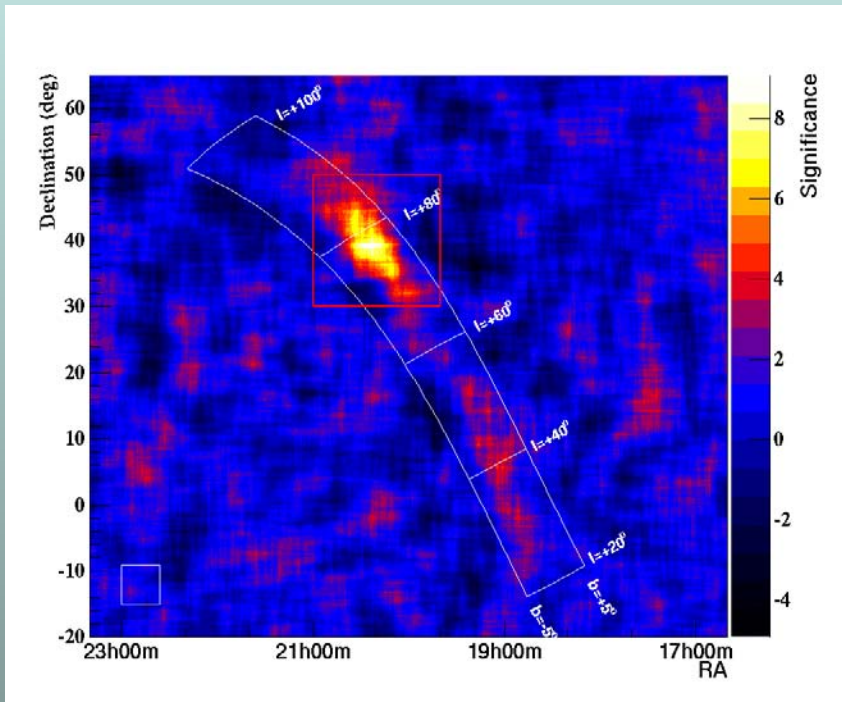
Tibet [Ohnishi]

Flux upper limit, consistent with Milagro detection.

OG 2.1: Diffuse γ -ray Sources

2. Cygnus Region

HEGRA
TeV 2032+4130



Milagro [Smith]
All-sky survey – for extended sources.
Cygnus region: most luminous in N. sky.
Significance $\sim 7\sigma$.

Milagro excess (color).
EGRET diffuse model (contours).

OG 2.1: Diffuse γ -ray Sources

Galactic Diffuse Emission

New Sources:

Galactic Center

HESS [Hinton]

Galactic Plane

Milagro [Sinnis]

Cygnus Region

Milagro [Smith]

Limits:

Galactic Plane

CANGAROO [Ohishi]

TIBET [Ohnishi]

Cygnus Region

Whipple [Atkins]

Dwarf (M32, M33, etc.)

Whipple [Hall], MAGIC [Bartko]

OG 2.4

Gamma-Ray Bursts

OG 2.4: Gamma-Ray Bursts

→ No evidence for VHE emission from GRB's.

Limits:

Cherenkov

STACEE [Jarvis]

MAGIC [Bastieri]

Air Shower

Milagro [Parkinson] – satellite triggered GRBs.

Milagro [Noyes] – short duration events.

ARGO-YBJ [Girolamo]

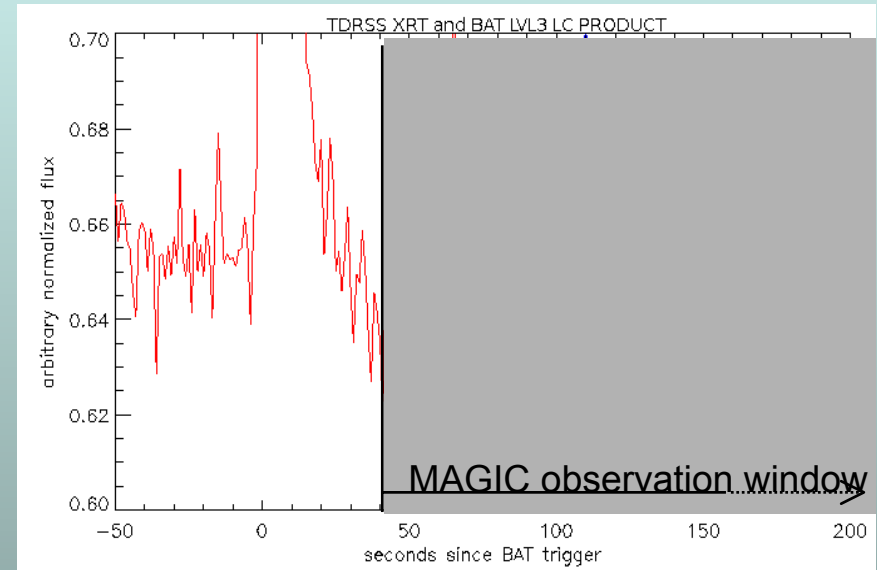
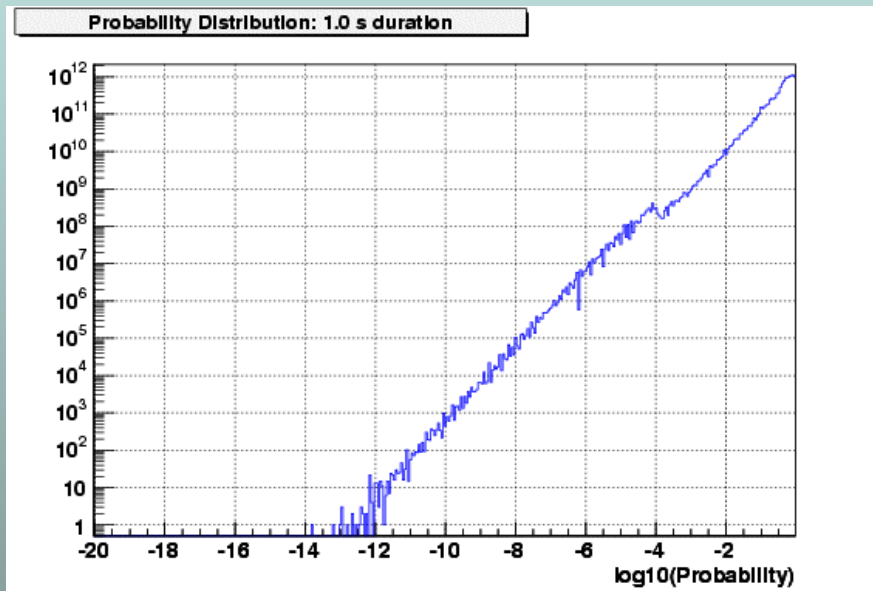
Andyrchy, Baksan [Smirnov]

Sierra la Negra [Alvarez]

Auger Observatory [Allard]

OG 2.4: Gamma-Ray Bursts

GRB Observations



MAGIC [Bastieri]

GRB050713A: data-taking started 40s after burst.

Preliminary analysis: no signal.

STACEE [Jarvis] will similarly reach bursts in < 2 min.

Milagro [Noyes]

Untriggered GRB search.

Search at all times/positions/durations.

No bursts detected.

OG 2.3

Extragalactic Sources

OG 2.3: Extragalactic Sources

Active Galactic Nuclei (AGN)

- Blazars – wideband non-thermal spectra, highly variable.
- VHE sources of the HBL (“extreme BL Lac”) type and one radio galaxy (M87).
- Models: Leptonic versus Hadronic cascades.

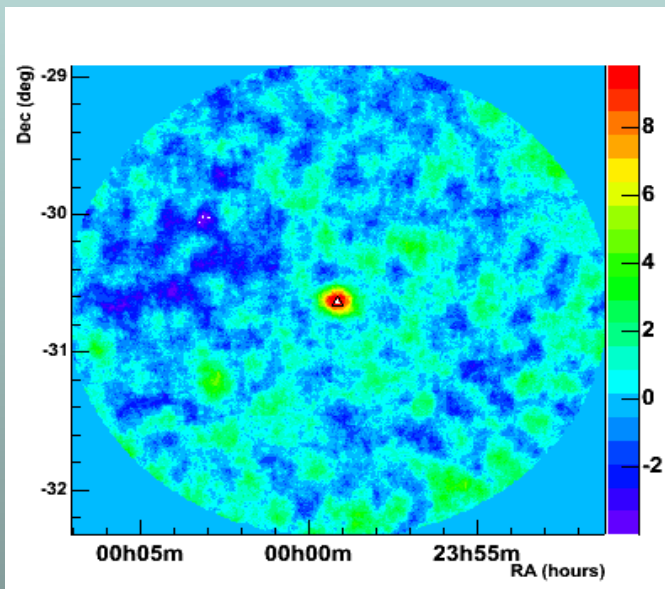
NEW SOURCES:

PKS 2005-489	$z=0.071$	HESS [Benbow]
H2356-309	$z=0.165$	HESS [Pita]
1ES1218+304	$z=0.182$	MAGIC [Meyer]
1ES1101-232	$z=0.186$	HESS [Tluczykont]

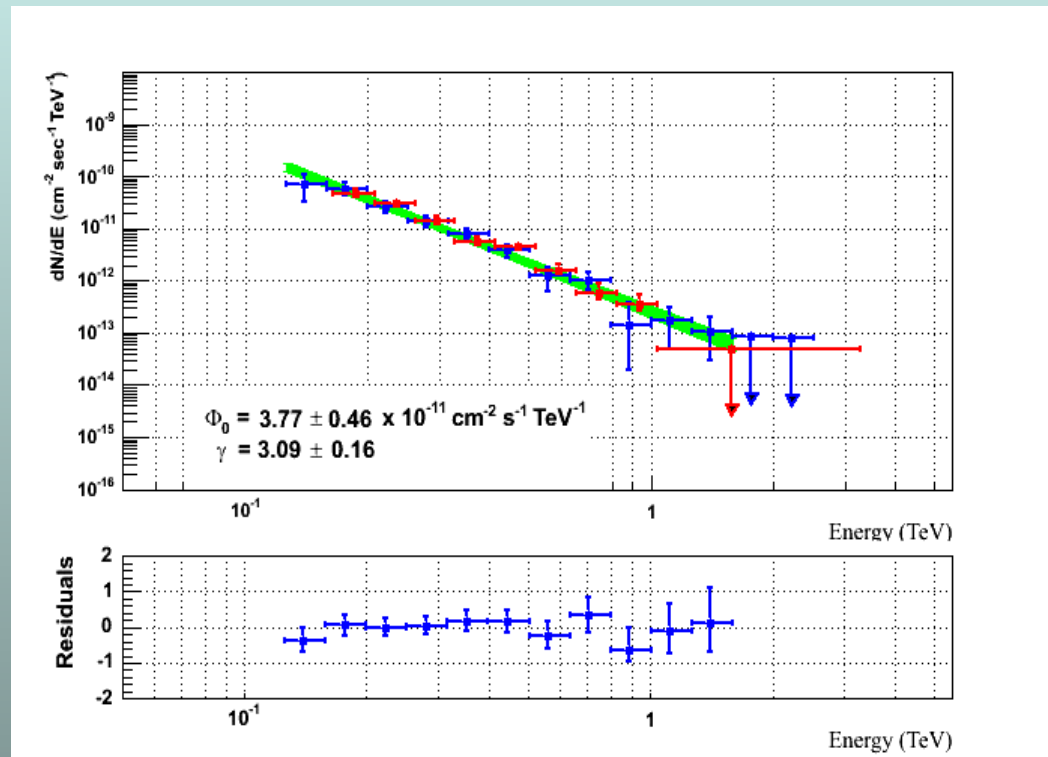
OG 2.3: Extragalactic Sources

H2356-309

($z=0.165$)



HESS [Pita]:
2004 data:
~39 hrs, 9σ , pt. source

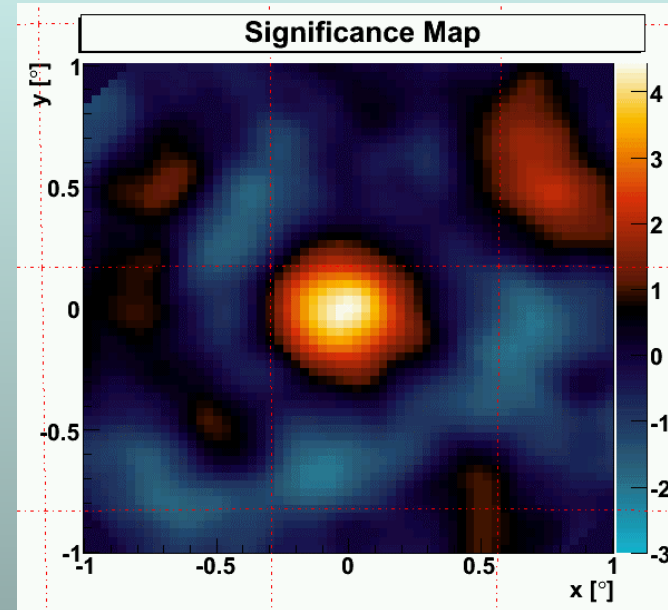
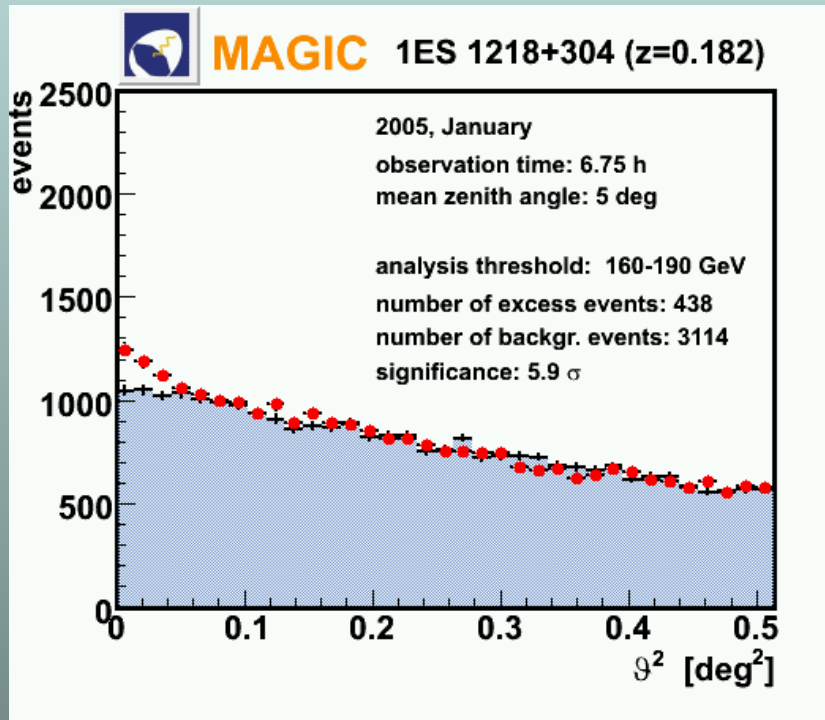


Spectrum:
Good power-law fit, $\Gamma \sim 3.1$
Flux $\sim 3\%$ Crab.

OG 2.3: Extragalactic Sources

1ES1218+304

($z=0.182$)



MAGIC [Meyer]:
2005 data:
 ~ 7 hrs, 6σ
no spectrum yet.

OG 2.3: Extragalactic Sources

AGN Summary

Source	Redshift	Type	First Detection	Confirmation
M87	0.004	FR I	HEGRA	HESS
Mkn 421	0.031	BL Lac	Whipple	Many
Mkn 501	0.034	BL Lac	Whipple	Many
1ES 2344+514	0.044	BL Lac	Whipple	HEGRA
1ES 1959+650	0.047	BL Lac	Tel. Array	Many
PKS 2005-489	0.071	BL Lac	HESS	
PKS 2155-304	0.116	BL Lac	Mark VI	HESS
H1426+428	0.129	BL Lac	Whipple	Many
H2356-309	0.165	BL Lac	HESS	
1ES 1218+304	0.182	BL Lac	MAGIC	
1ES 1101-232	0.186	BL Lac	HESS	

→ Reaching further out in redshift.

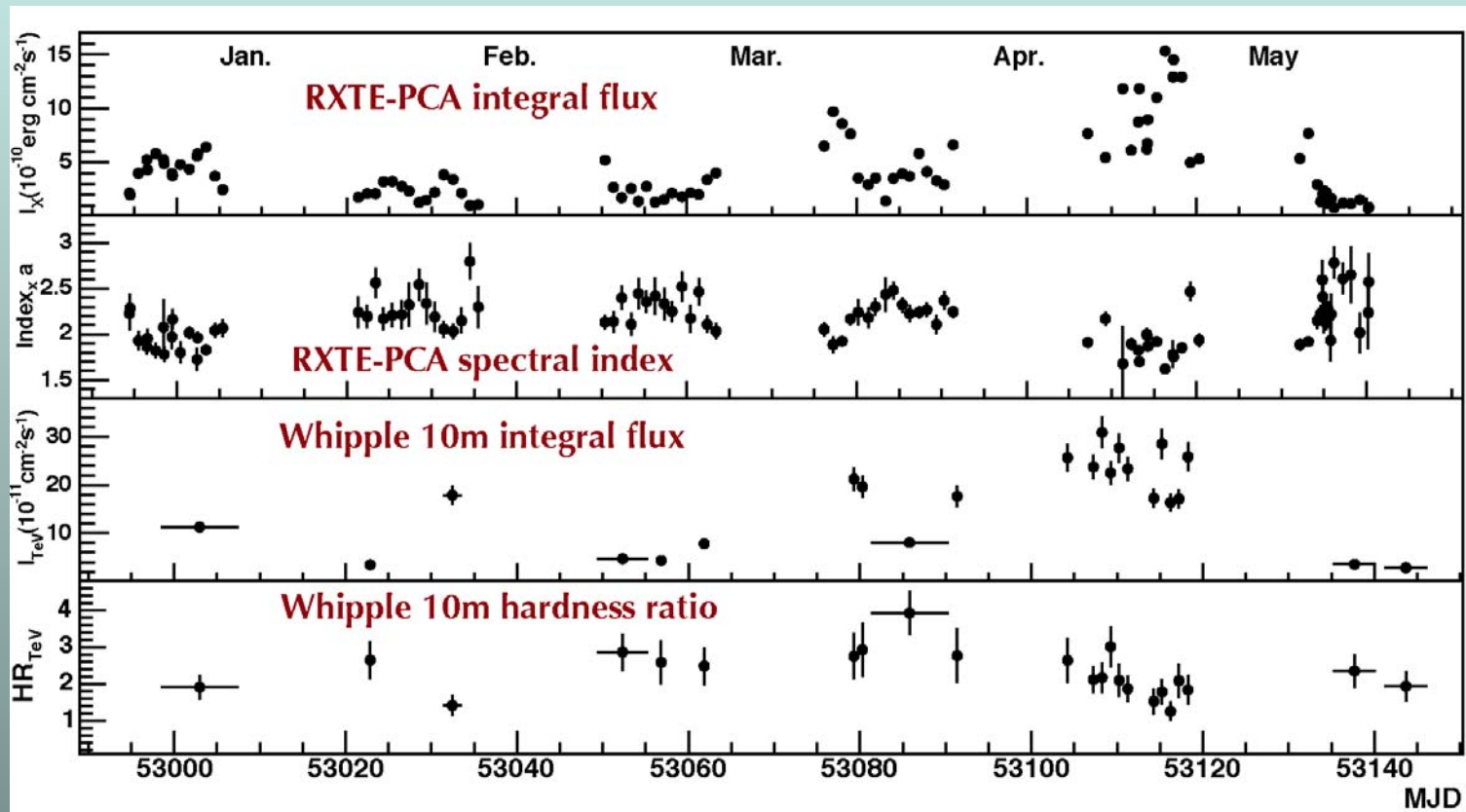
OG 2.3: Extragalactic Sources

- Detailed studies of established AGNs.

Mkn 421	Whipple [Grube], STACEE [Carson], MAGIC [Mazin], CANGAROO [Sakamoto], HESS [Horns], PACT [Bose], VERITAS [Cogan], TACTIC (Rannot[
Mkn 501	MAGIC [Mirzoyan]
PKS 2155-304	HESS [Fontaine], CANGAROO [Sakamoto]
1ES1959+650	MAGIC Tonello]
M87	HESS [Beilicke], MAGIC [Albert]

OG 2.3: Extragalactic Sources

Mkn 421

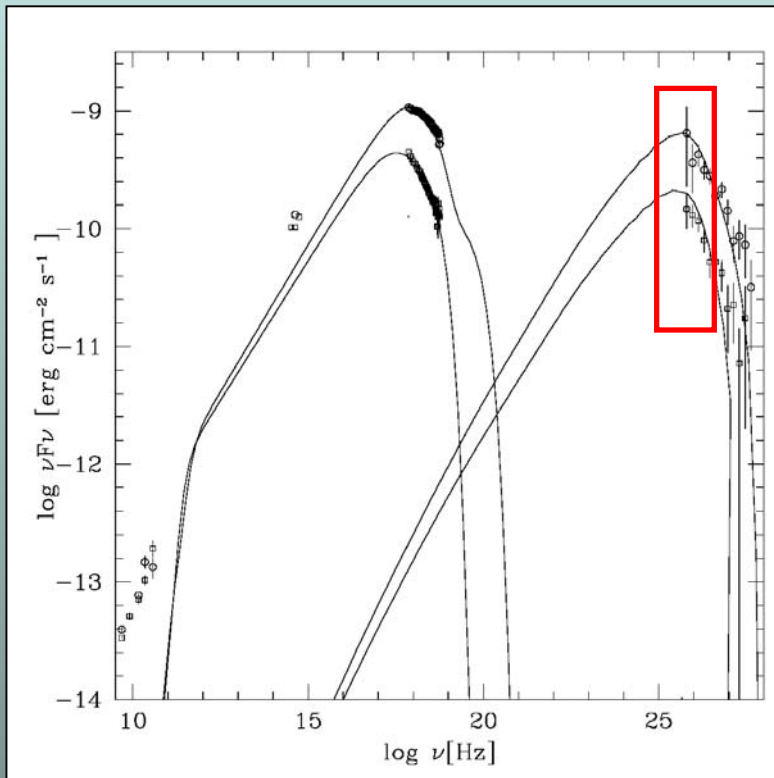


Whipple [Grube]:

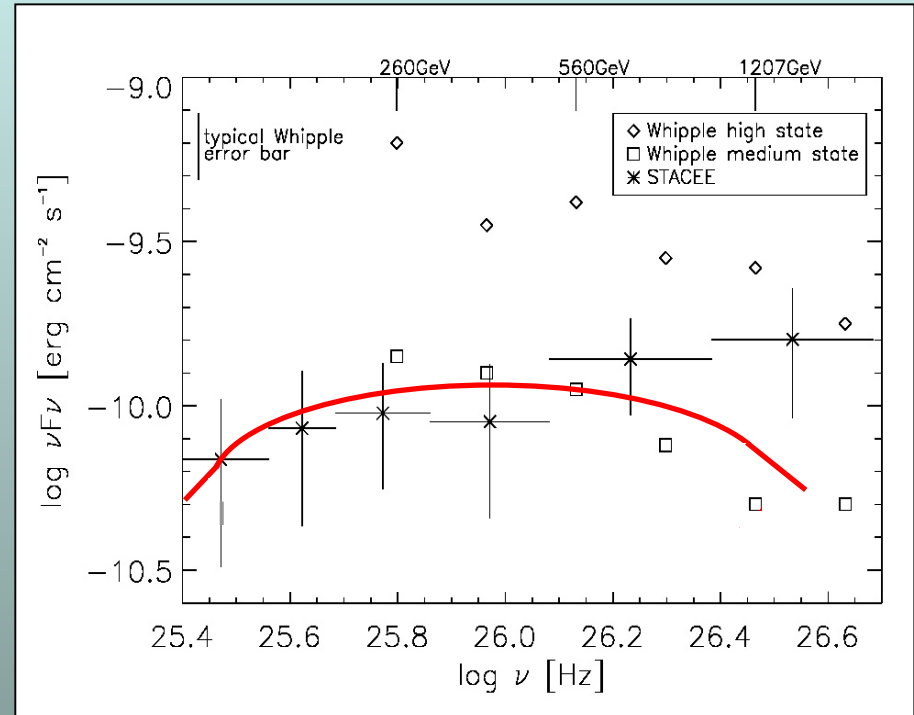
2004: Five month multi- λ campaign: X-ray, TeV γ -ray.

OG 2.3: Extragalactic Sources

Mkn 421



Whipple and RXTE:
Spectral Energy Distribution.
“Low” and “high” flux states.

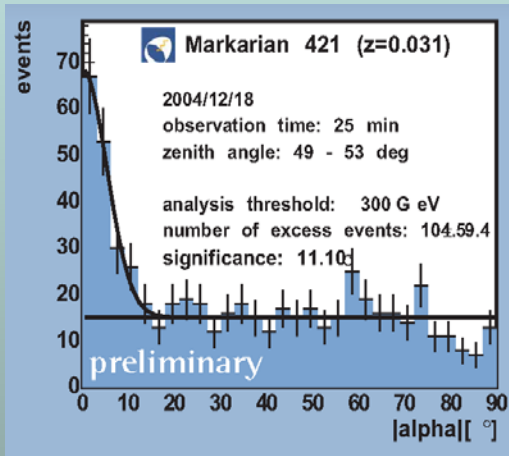


STACEE [Carson]:
Spectrum extends down to 100 GeV.
See peak of IC component ?

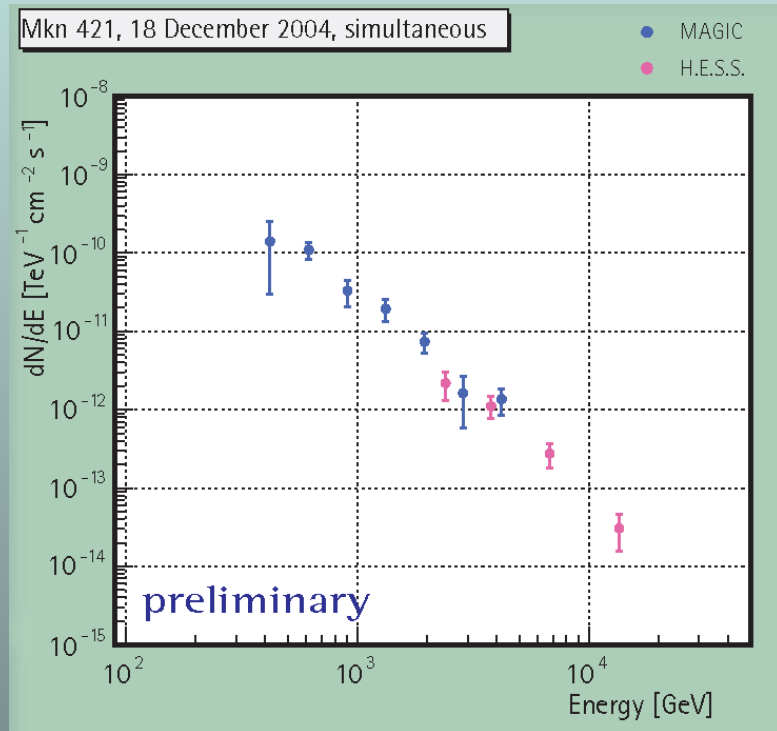
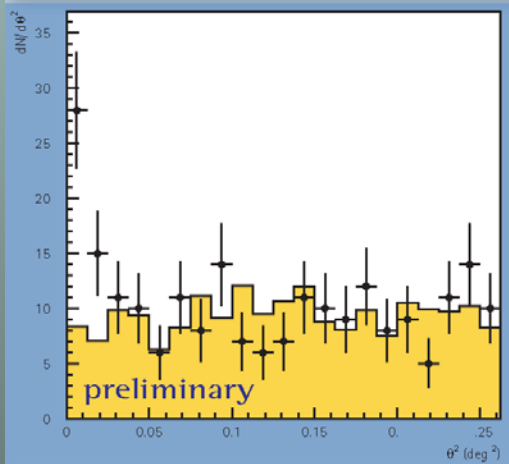
OG 2.3: Extragalactic Sources

Mkn 421: Simultaneous Observations [Mazin]

MAGIC



HESS

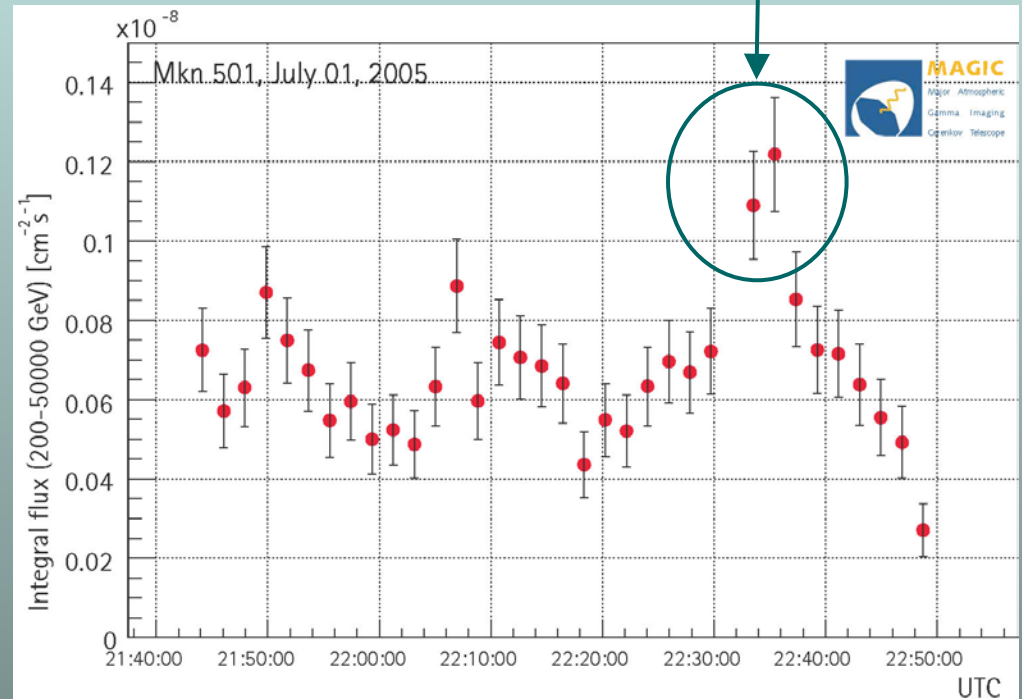
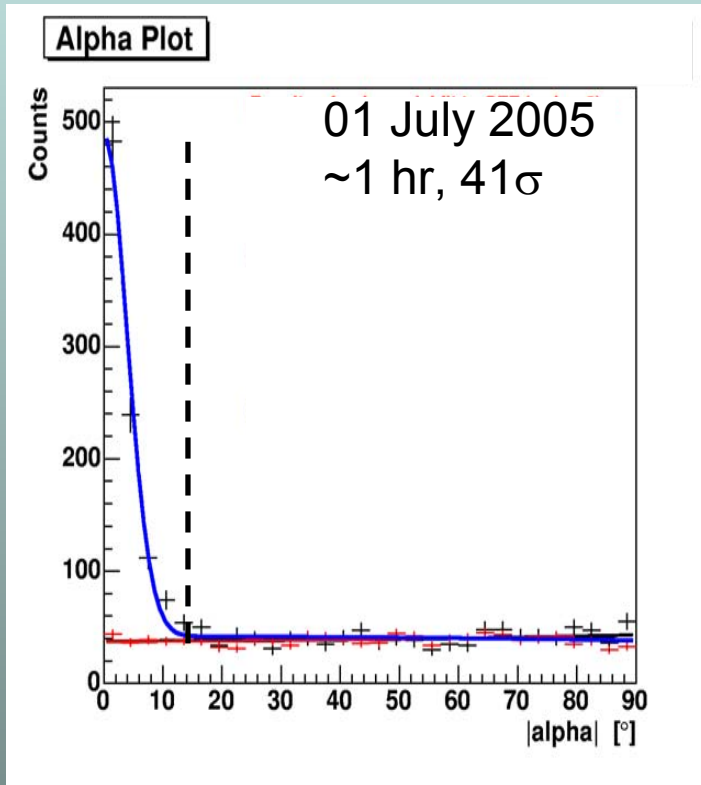


Combined spectrum – agreement !

OG 2.3: Extragalactic Sources

Mkn 501: Giant flare July 2005

V. rapid variation
(< 4 min)

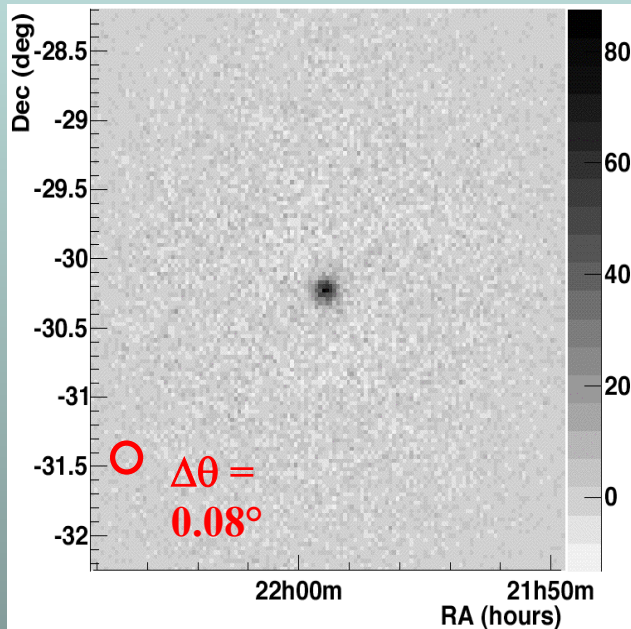


MAGIC [Mirzoyan]:

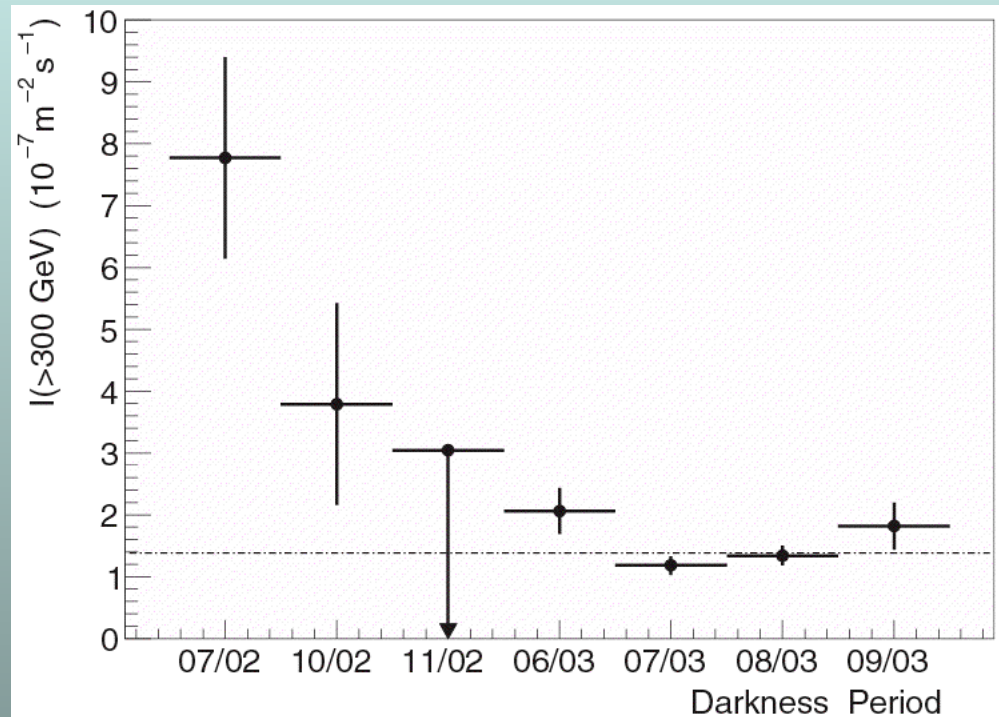
Light Curve: 2 min time bins.

OG 2.3: Extragalactic Sources

PKS 2155-304



HESS [Fontaine]:
2003 data: 34σ



Flux steadily decreasing.
Spectral index does not vary with flux level.

Reached quiescent level ?

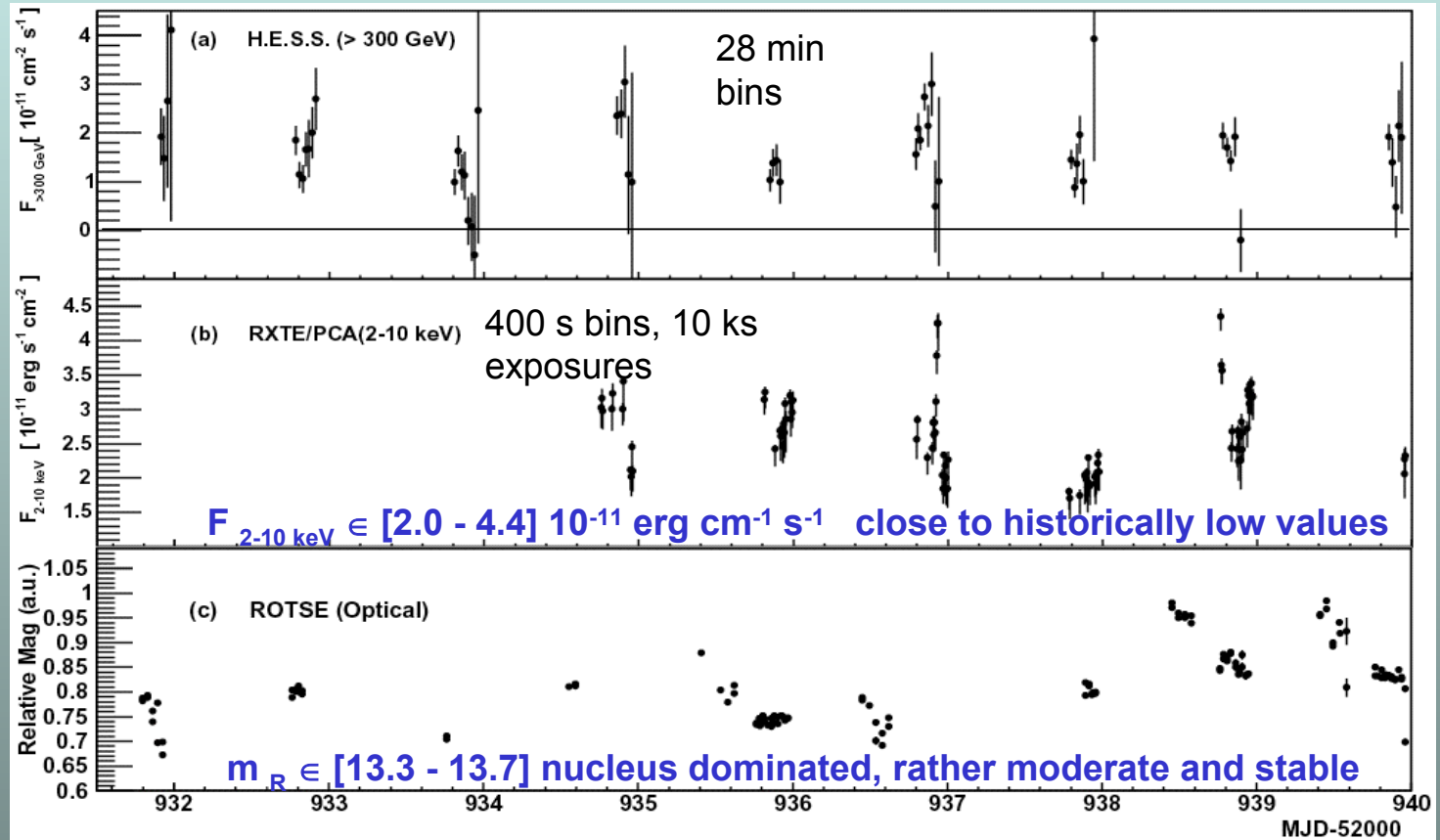
OG 2.3: Extragalactic Sources

PKS 2155-304

VHE γ -rays
(HESS)

X-rays

Optical



→ Low state detected in γ -rays ?

OG 2.3: Extragalactic Sources

- Limits from new extragalactic observations.

AGN:

Surveys

W Comae, 3C 66A

1ES2344+514

H1426+428

Cen-A

HESS [Benbow], MAGIC [Meyer]

STACEE [Mukherjee]

TACTIC [Godambe]

TACTIC [Thoudam]

CANGAROO [Kabuki]

Non-AGN:

NGC 253 (starburst)

Galaxy Clusters

HESS [Lemoine-Goumand]

Whipple [Perkins]

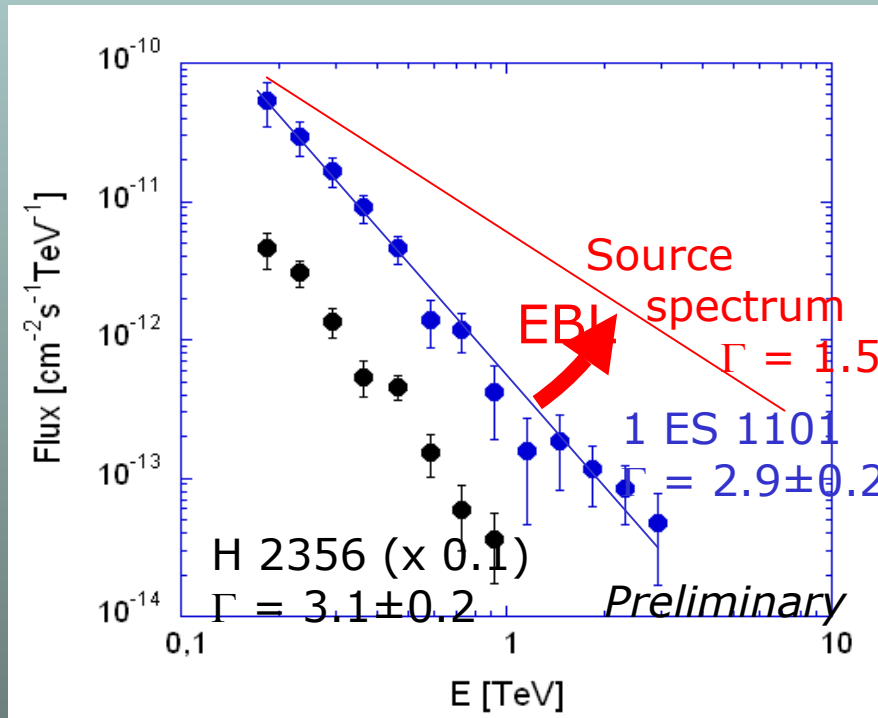
OG 2.3: Extragalactic Sources

Extragalactic Background Light (EBL)

- Diffuse O/IR radiation from star formation and evolution.
- Spectral signature from $\gamma\gamma$ absorption for $E_\gamma \sim 50\text{-}2000$ GeV.
- Use measured AGN spectra to constrain EBL.

HESS
[Tluczykont]

H2356 and
1ES 1101



Appears to be less
EBL than expected.

Close to lower bound
from galaxy counts.

OG 2.7

New Experiments and Instrumentation

Will say very little regarding instrumentation: many papers on detectors, techniques, and simulations.

See also Cherenkov Workshop in Palaiseau (May 2005).

OG 2.7: New Experiments

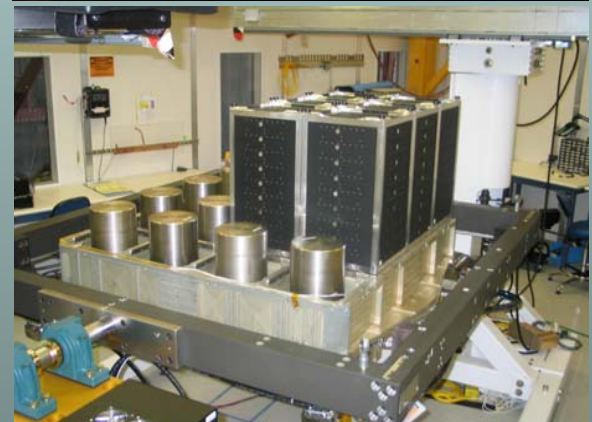
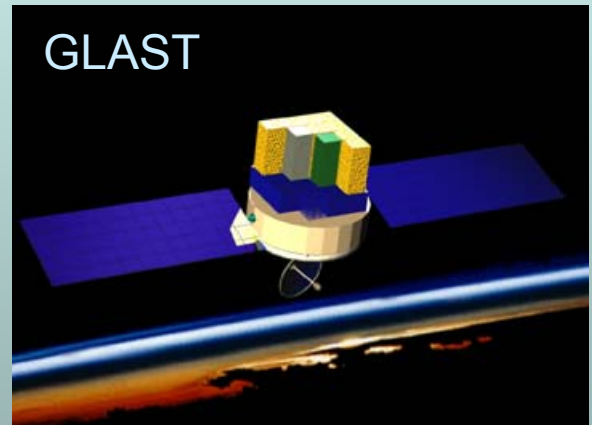
Satellite

1. ASTROSAT [Agrawal]

- UV/X-ray mission.
- 5 instruments: 0.3-100 keV range.
- Scheduled for 2007.

2. GLAST [Moiseev]

- New, major HE g-ray mission.
- LAT: coverage from 20 – 300 GeV.
- Scheduled for 2007.



LAT construction at SLAC
June 2005.

OG 2.7: New Experiments

Cherenkov Telescopes

1. VERITAS [Holder]

- 4x12m telescopes for Kitt Peak site → (2006).
- 2 telescopes operated this fall at Whipple Base Camp.



VERITAS
Telescope 1



OG 2.7: New Experiments

Cherenkov Telescopes

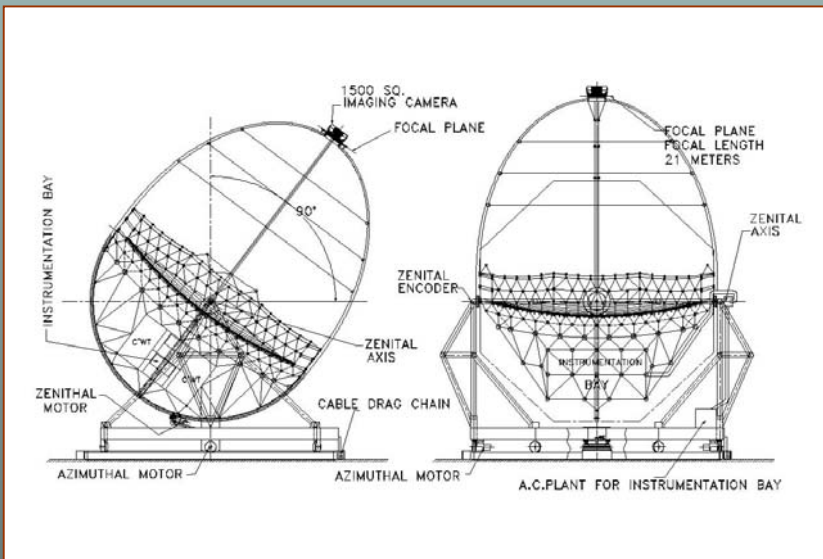
2. HAGAR [Chitnis]

- 7 telescopes by late 2006.
- Hanle site, 4200m a.s.l.



3. Mace [Koul]

- 2 x 21m imaging telescopes.
- 4° x 4° camera.
- Hanle site.



OG 2.7: New Experiments

Cherenkov Telescopes

4. HESS-II [Vincent]

- New 28m telescope.
- 2048 pixel camera.
- Lower energy 40-50 GeV.



5. MAGIC-II [Teshima]

- New 17m telescope.
- Possible high-QE camera.
- 2007 schedule.

OG 2.7: New Experiments

Air Shower Experiments



1. ARGO-YBJ [Girolamo]

4300m ASL
6,000 m² RPC detector
Scalers sensitive \sim GeV energies.



2. Pierre Auger [Allard]

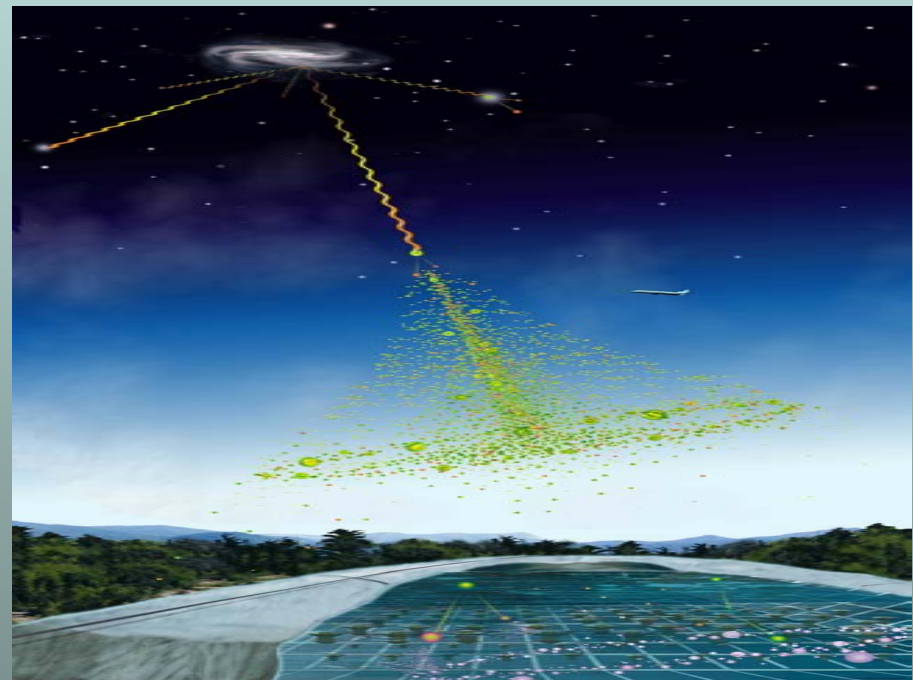
Single particles in Auger water tanks sensitive to GRBs.
A $\sim 1600 \times 10$ m².

OG 2.7: New Experiments

Air Shower Experiments

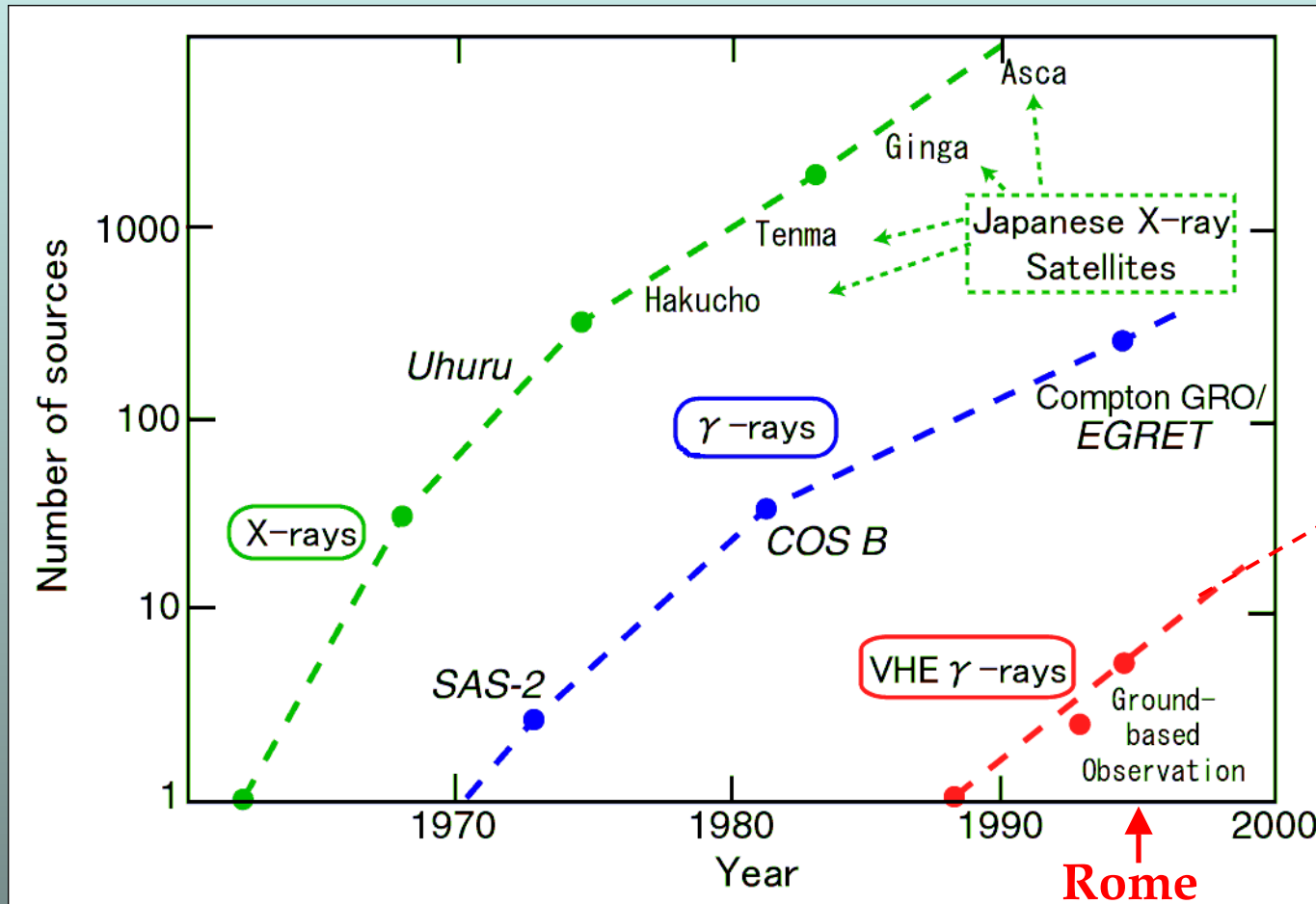
3. HAWC [Sinnis]

- 300m x 300m pond.
- Wide FOV ~ 2 SR.
- High altitude site to go lower in energy.
- In R&D phase.



HAWC Concept

“Kifune Plot”



Source count versus year
[T. Kifune]

Pune
2005

SUMMARY

- **New generation of Cherenkov telescopes has yielded outstanding results.**
- **Many new VHE sources discovered in last two years ... an almost unprecedented increase.**
- **Galactic Plane is rich in the number and type of VHE emitter. Pulsar nebulae and SNRs are both firmly established in the TeV band.**
- **New discoveries increase the number of known TeV blazars and push further out in redshift.**
- **Future experiments on ground and in space should continue the rapid development of VHE astrophysics.**