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**A Computer Code to Calculate Line by Line
Atmospheric Transmission Spectra
on a Microcomputer**

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INTRODUCTION

To support work being done in the Instrument Electro-Optics Branch in remote sensing a decision was made to set up an in-house capability to compute line by line atmospheric transmission spectra.

An available DEC 11/23 microcomputer was dedicated to this task; its VT100 terminal had been modified by another manufacturer to greatly enhance its graphics capability, so that plots of the computed spectra could be displayed directly on the screen.

The computer program, called MICTRA (for MICrocomputer TRANsmission calculations) is based on the LASER code published by the Air Force Geophysics Laboratory in 1978 (Reference 1); it uses the 1982 edition of the AFGL Atmospheric Absorption Line Parameters Compilation (Reference 9).

The Air Force code was extensively reworked to allow execution on a desk top microcomputer and to generate transmission spectra rather than tables of extinction coefficients. The program treats absorption due to lines of water vapor, carbon dioxide and ozone throughout the spectrum, from the visible to the microwave region, plus certain "continuum" absorption effects, Rayleigh (molecular) scattering and aerosol absorption and scattering. The program does not include radiance calculations.

The program described in this report, along with another program and a plot package developed by the author (to be published — see References 2 and 3) have proven useful in supporting work being done by the Branch in the development of a submillimeter heterodyne radiometer, in lidar studies in the five micrometer region and in feasibility studies for near infrared laser communications.

The MICTRA program was developed by the author from a previous MICTRA routine developed in the Instrument Electro-Optics Branch; this earlier routine was based on the Air Force LASER routine. For clarity these three routines will be referred to as the "new" MICTRA routine, the "old" MICTRA routine and the Air Force routine, respectively.

The old routine was constructed primarily by deleting unneeded portions of the Air Force routine. The most important deletions were:

1. Restriction of the atmosphere and aerosol models to a maximum altitude of 15 km (as compared to the original 100 km) and deletion of the Voigt profile calculation;
2. Restriction of the aerosol models, as functions of wavelength, to a small wavelength range centered at 5 micrometers;
3. Deletion of four of the seven atmospheric absorbing species handled by the Air Force routine — only water vapor, carbon dioxide and ozone were retained;
4. Deletion of molecular (Rayleigh) scattering;
5. Deletion of the 3.5-4.2 micrometer water vapor continuum and the 4-5 micrometer nitrogen continuum.

These changes were made because MICTRA was needed primarily to support lidar work in the 5 micrometer region of the spectrum, at altitudes not exceeding about 15 km and involving water vapor, carbon dioxide and ozone. Below 15 kilometers, and for wavelengths of 1 micrometer or more, the lines are mainly pressure broadened (so the Voigt profile is not needed for a reasonably accurate calculation), and Rayleigh scattering is significant only in the visible part of the spectrum. In the new version of MICTRA all but the first and third of the above features have been restored.

It should be noted that, for wavelengths shorter than 5 micrometers, the use of the Lorentz line profile in place of the Voigt profile becomes less accurate. At 5 micrometers, the ratio of the Voigt halfwidth to the Lorentz halfwidth (V/L) ranges from about 1.02 at 15 km altitude to 1.0006 at sea level. At 1 micrometer the V/L ratio ranges from 1.38 at 15 km to 1.02 at sea level, and at 0.5 micrometers it ranges from 2.02 at 15 km to 1.06 at sea level. (All of these ratios hold to a good approximation for any of the three species included in the program: water vapor, carbon dioxide or ozone.) Because of this decrease in accuracy of the Lorentz approximation at shorter wavelengths it is planned to include the Voigt profile in the transmission calculation in the near future.

Neither the old routine nor the Air Force routine calculated transmission spectra; rather, they calculated and printed out tables of absorption and extinction coefficients for all of the atmospheric layers and for the various atmosphere and aerosol models.

The author's main purpose in modifying the old MICTRA routine was to make it capable of calculating transmission spectra over some given range of wavelength or frequency.

The basic strategy used to accomplish this was to use the old MICTRA routines as subroutines called by a newly-created executive routine, which directs the calculation of the spectrum. The executive routine was also given extensive capability to prompt the user to type in information at the keyboard which is used to set up the run.

In addition, the full tables of aerosol coefficients as functions of wavelength were restored, because new Branch uses for the program involved other regions of the spectrum than just the 5 micrometer region. It is planned to replace the aerosol models by newer models which take account of the effect of condensation of water vapor on the haze particles, as a function of relative humidity. The full 100 km atmosphere and seven absorbing species may also be restored. As mentioned above, several effects included in the Air Force routine but deleted in the old MICTRA routine were also restored: Rayleigh scattering and the continuum absorption effects for water vapor at 3.5-4.2 micrometers and for nitrogen at 4-5 micrometers.

An auxiliary plotting routine (PLOTSP) was written to display the spectra computed by the new MICTRA; this routine uses a plot package designed and written by the author specifically for the microcomputer configuration used to run MICTRA (see Reference 3).

Numerous other changes connected with the coding of the routines, were made, but these are too detailed to merit description here. (The various line files used in the calculation of the spectrum are described in the section of this report dealing with the structure of the program.)

WHAT THE PROGRAM DOES

The program calculates atmospheric transmission spectra on a line-by-line basis over a variety of paths, spanning an arbitrarily chosen wavelength or frequency range. Radiance effects are not included in the calculation.

The path may be chosen to be horizontal, vertical or slanted at any zenith angle, and may extend between any two altitudes contained in the atmosphere model, which is a layered, flat model.

The range covered by the spectrum may be chosen freely and may be specified in angstroms, micrometers, hertz or wavenumbers, depending on the preference of the user.

The atmosphere models used are three of the models used in the original Air Force program: the U.S. Standard Atmosphere, 1962 and two mid-latitude models for average summer and average winter conditions. The aerosol models used are those devised by Shettle and Fenn in 1976 (Reference 4). (A more recent version of these models is available, which includes the effects of relative humidity on the size of the aerosol particles; it is planned to include these models in the future.) The atmosphere models, which originally extended up to 100 km altitude, were modified to extend only to 15 km altitude, because higher altitude effects either were not needed for our work or were not significant. Also, of the seven main atmospheric absorbing species included in the original program, only three were retained: water vapor, carbon dioxide and ozone, because the other species were not essential for our work.

The absorption line data used by the program are taken from the 1982 edition of the main tape compiled by the Air Force Geophysics Laboratory, which lists all the known absorption lines for the seven most important atmospheric absorbing species, from the visible to the microwave region (References 5-9). Segments of line data needed are read from the Air Force tape, preprocessed into the form used by the program and stored on floppy disks.

MICTRA computes a transmission spectrum by computing the transmission over a net of points spanning the chosen wavelength or frequency range; at each point the extinction coefficients are computed by adding up several effects:

- The absorption due to all atmospheric lines (of water vapor, carbon dioxide and ozone) in a neighborhood of the current wavelength, chosen so that absorption due to the tails of more distant lines is negligible;
- Molecular continuum absorption: the 3.5-4.2 and the 8-14 micrometer water vapor continua and the nitrogen continuum at 4-5 micrometers;
- Rayleigh (molecular) scattering;
- Aerosol absorption and scattering (at present the 1976 models of Shettle and Fenn are used; their newer models, which include the effects of relative humidity, will be incorporated in the near future).

All these effects are computed for each of the atmospheric layers (there are fifteen 1 km-thick layers) which the chosen path traverses, and are combined to form the total extinction for the given wavelength, over the given path. The transmission values thus calculated over the chosen wavelength or frequency range are then stored in a disk file, for later display by the auxiliary plotting routine PLOTSP.

The plotting routine allows the user the option of displaying the computed transmission values vs wavelength in either angstroms or micrometers or vs frequency in either hertz or wavenumbers, depending on the region of the spectrum and on the preference of the user. Further, whatever scale is chosen may be displayed as increasing toward either the right or the left of the plot.

Two types of spectral plots may be computed by MICTRA and displayed by the plotting routine: a "continuous" spectrum which consists of the transmission over an equally-spaced net of points, joined by line segments to give a continuous appearance, or an "overview" spectrum in which the transmission is computed only at the exact center of each line in the chosen range. In the overview type of spectrum the plot shows the continuous background transmission due to aerosol and molecular continuum effects as a straight line (a good approximation over small wavelength ranges), with the line-center values displayed as vertical lines coming downward from the continuum; the value at the bottom of each line is, of course, the correct total transmission at that wavelength. To avoid the problem of the equally-spaced grid of points in a continuous spectrum skipping over the centers of absorption lines, thus not showing the true depth of the line or even missing it entirely, the line-center wavenumbers are intercalated into the equally spaced grid; thus the bottoms of all lines are correctly shown.

The overview type of spectrum is intended to give a picture of the gross properties of the spectrum over a large range (but small enough to contain at most a few hundred lines, and to ensure that the linear approximation to the background continuum is good), while the continuous type of spectrum is intended to show the detailed structure of the spectrum, including the shapes of the lines, over a smaller range.

The program allows the user to interactively control or select (at the terminal keyboard) many of the parameters and data sets used by the program and many of the functions provided by the program. For example, the user may: select any of the three atmosphere models and any of the aerosol models provided; select the path; change the sea-level values of temperature, pressure or other parameters in the program; specify the wavelength or frequency range of the spectrum; specify the fineness of the grid for a continuous spectrum; etc. The program also allows the user to interactively retain or reject any or all of the extinction effects, in order to see clearly how great each effect is, taken by itself. For example, the user may decide to compute a spectrum showing only the absorption due to lines, with no aerosol or molecular extinction considered; to see just how strong each line is by itself, the user could choose the neighborhood of surrounding lines to be so small that only the absorption due to each line by itself is computed. Further, any or all lines in the chosen range may be independently selected or rejected; this is accomplished by commanding the program to display each line on the screen, and telling the program whether to include it. Provision is made to accept or reject not just individual lines, but also all lines of a given species or all lines in a freely chosen wavelength range; thus a spectrum could be computed which shows only ozone lines, for example, or even a single line.

Any number of spectra may be set up and computed in a single run; this is made possible by the BATCH compiler, which functions for the microcomputer used as a kind of simplified job control language. Each spectrum is stored in its own disk file, and may later be displayed by the plotting routine.

HOW TO USE THE PROGRAM

The program was designed to use the interactive capability of microcomputers; the user communicates with the program through the terminal keyboard and the video screen.

The user only needs to execute the command file RNMICB (which stands for "run the batch version of MICTRA"), which then takes over and directs the user to execute other command files, depending on the type of run. Three types of runs are provided: create a BATCH file which can be executed later to compute a spectrum; compute a spectrum by executing a previously created BATCH file; or both create and execute a BATCH file.

If one of the last two run types is selected, the command file will automatically call MICTRA, after instructing the user to mount the proper floppy disk in the proper device; MICTRA will then take over and lead the user step-by-step through the run setup and execution.

All data sets needed by the program are stored internally, with one exception: the line file must be supplied by the user. The line file is a segment of data condensed from the Air Force Geophysics Laboratory main tape, which is a compilation of about 190,000 atmospheric absorption lines spanning the spectrum from the visible to the microwave region. Because MICTRA is implemented on a microcomputer with no tape drive, it is necessary to take the needed segment of data from the Air Force tape and store it on a floppy disk. The segment of data also must be "condensed" before MICTRA can use it; this is accomplished by using the auxiliary routine LNEDIT. The condensing process consists of taking the lines from the tape, which are coded as character strings in ASCII, stripping away certain parameters not needed by MICTRA (primarily quantum identifications), and converting the remaining numbers from character strings to binary form. The resulting file is about one fourth the size of the unprocessed file; this allows rather large segments of line data to be stored on a floppy disk.

Each computed spectrum is stored in a disk file, specified by the user upon prompting by MICTRA. The spectrum may be displayed by using the auxiliary routine PLOTSP, which instructs the user how to use it and prompts for any needed information or choices.

Because MICTRA leads the user through the run setup process, it is not necessary to list the required input here; the information and choices required by MICTRA may be clearly seen by examining the listing, which is well documented with explanatory comments. It might be noted that the sequence of items needed for setting up a run is not fixed; many choices are provided, and different choices will result in different sequences of input items.

To eliminate the possibility that the user might wait for a long run to finish without any knowledge of what is going on, the program types each spectrum point (wavelength or frequency and transmission) on the screen as it is computed; thus the user can abort a run that looks useless, instead of waiting blindly for it to finish.

STRUCTURE OF THE PROGRAM

The basic organization of the program is shown in Figure 1. MICTRA is the driver and main computing module; it leads the user through the run setup process and directs the calculation of the spectrum.

The logical flow of the spectrum calculation involves several nested loops. The outside loop steps through the sequence of frequency points at which the transmission is to be calculated. For a continuous type spectrum this will be an equally-spaced grid of frequency points, spanning the desired frequency range; for an overview (line-center) type of spectrum the set of frequency points will be the exact centers of all the absorption lines contained in the frequency range.

The next inner loop steps through the atmospheric layers contained within the altitude range spanned by the chosen atmospheric path. The innermost loop steps through the absorption lines contained in the frequency range of the spectrum.

Thus, for each frequency the program calculates the total absorption coefficient for all the lines in each atmospheric layer, as well as the continuum absorption coefficients, the Rayleigh scattering coefficient and the aerosol absorption and scattering coefficients for each layer. These are all added up to give the total extinction coefficient for each layer. Then the transmission for each layer is computed using the usual decreasing exponential law for extinction ($e^{-\text{depth}}$, where depth is the product of the total extinction coefficient and the path length through the layer), and finally the transmission over the entire path is computed by multiplying together the transmissions for the various layers.

The subroutines communicate with one another through the three COMMON blocks, which hold variables needed by more than one routine, as well as by arguments in the call sequences.

The program needs two external data sets: a file containing the atmosphere and aerosol models and a file containing a segment of data from the Air Force line tape. The atmosphere model file is located on a floppy disk along with the program load module; it is automatically read into the program by subroutine SETUP and stored in one of the COMMON blocks. The line file must be supplied by the user; it is created by taking the needed segment of data from the Air Force tape and using the auxiliary routine LNEDIT to condense it into the form used by MICTRA.

After the MICTRA routine has obtained the required information from the user (via the keyboard), it forms the run line file, which consists of the lines needed for the current run, by taking from the line file only those lines selected by the user in the run setup process. The run line file may contain all of the lines in the line file, some of them or none of them; the lines to be included may be selected individually, by species or by sub-ranges of the spectrum frequency range.

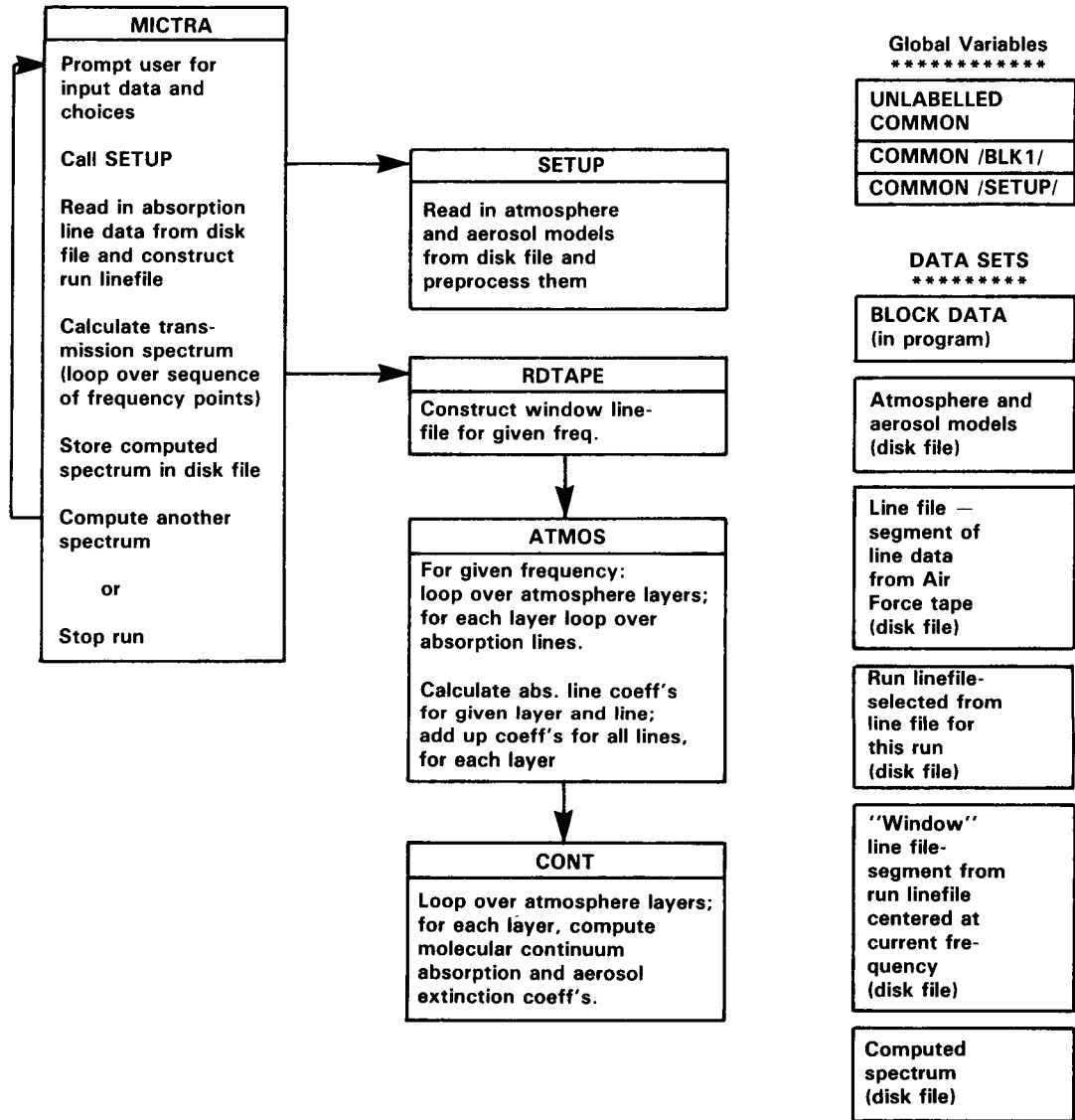


Figure 1. Basic Flow of MICTRA Program

During the calculation of the spectrum subroutine RDTAPE forms a "window" line file for each frequency at which the transmission is to be computed; this consists of all the lines in the run line file which are contained in a window of specified frequency width centered at the current frequency (the width is chosen by the user during the run setup process). The purpose of the window line file is to include, for the calculation of the absorption coefficient at the current frequency, all those lines within some neighborhood of the current frequency which will affect the absorption appreciably. If the user wants to calculate the absorption due to a single line as if it were isolated (even though it may actually have close neighbors which affect it), the window width may be chosen small enough to exclude all nearby lines.

In the implementation of MICTRA used by the author on the DEC 11/23 microcomputer, the line file is located on a floppy disk, while the run file and the window file are located in a part of the computer memory which is treated as a disk (logically) by the system software; thus, the operation of the program is much faster than it would be if these files were located on floppy disks. (In the event that the run file or the window file is too large to fit into that part of the memory, it is automatically put in a file located on a DEC RL02 disk drive, which holds 10.4 megabytes and is much faster than a floppy disk drive.)

MEMORY REQUIREMENTS AND OPERATING SPEED

The entire program (i.e., the load module) requires about 42 kilobytes of memory; this includes all the atmosphere and aerosol models, which are read into the program and stored internally. The line files should be small enough to fit onto a floppy disk. In the DEC computer being used by the author, which handles 8-inch double density floppy disks which hold 512 kilobytes, this allows the line files to hold up to about 20,778 lines, in the condensed form used by MICTRA; this should be large enough for most purposes. The auxiliary programs LNEDIT and PLOTSP, which respectively construct condensed line files and display the spectra, require about 13 and 32 kilobytes of memory, respectively.

The time required to compute a spectrum varies greatly, depending on the type of spectrum (continuous or only line centers) and on other factors such as the wavelength or frequency range covered, the number of absorption lines contained in the range, the width of the window of lines which significantly affect the absorption at the current frequency, the fineness of the frequency grid, etc. In general, the calculation of a line-center plot is much faster than the calculation of a continuous plot; the times may range from a few seconds for a line-center plot over a small interval to several hours for a continuous plot over a large interval with many lines and a wide window.

SOME EXAMPLES OF COMPUTED TRANSMISSION SPECTRA

In this section, several examples of transmission spectra are shown exactly as they appear on the terminal screen when displayed by the auxiliary plotting routine PLOTSP; the plots were copied directly from the screen by a TEKTRONIX hardcopy device.

Figure 2 shows a line-center spectrum in the near infrared, from 8250 to 8300 angstroms. The spectrum was computed for a slanted path extending from the ground to the top of the atmosphere model (at 15 km altitude); the zenith angle was taken to be 51 degrees. The midlatitude summer atmosphere model was used, along with a very hazy aerosol model; many of the lines show complete absorption (all the lines in this region happen to be water vapor lines). The plot gives a very clear overall picture of the absorption by water vapor and aerosols in this region. Figure 3 shows a continuous plot of a part of the region, from 8270 to 8280 angstroms; from this plot it is clear that the lack of line shapes in Figure 2 does not detract much from the qualitative picture of the spectrum, at least in this region. Both of these plots were computed with a window halfwidth of 20 wavenumbers (i.e., all lines within 20 wavenumbers to each side of the current frequency were taken into account in computing the absorption), which gives a very accurate spectrum.

Figure 4 shows a line-center spectrum in the submillimeter region, from about 180 to 200 micrometers. This spectrum was computed for a vertical path extending from 10 to 15 km altitude (if computed from the ground up, the water vapor absorption is so strong and the lines so broadened that the spectrum is virtually flat — almost everything gets absorbed). The midlatitude summer atmosphere model was used; the aerosol model does not extend to the submillimeter region, because the effect is negligible. To allow a very quick look — i.e., a fast (about 10 seconds) calculation of the spectrum — the lines were effectively considered separately by taking the window halfwidth to be only 0.001 wavenumbers; thus each line is shown as it would appear if no other lines were present. In this region both water vapor and ozone lines appear, but only water vapor lines were included in the spectrum. For comparison, the spectrum in Figure 5 was computed. This shows the same spectrum as Figure 4 does, but in a continuous form, to show its actual appearance. Also, the window halfwidth for this spectrum was taken to be 10 wavenumbers, so that it is more accurate. Because of this, and because a fine grid of points was used to span the spectral range, this spectrum took about 2 hours to compute. (It should be recalled that the exact line centers of all lines are intercalated into the grid, so that the line depths are accurately shown — there is no skipping over line bottoms due to grid points missing the line.)

It is interesting to note, from Figure 5, that the water vapor line at about 187 micrometers is extremely strong; the plot clearly shows other, weaker lines extending down from its very broad wing on the high wavelength side; on its low wavelength side there is another very strong line, which is mostly off the plot. Inspection of the line file shows that the strong line at 187 micrometers is in fact unusually strong; its intrinsic line strength is about 45 times greater than the strength of the next strongest line near it (except for the line only partly shown, which is even stronger), and is from two to six orders of magnitude greater than the strength of most of its neighboring lines.

Lastly, Figure 6 shows a spectrum in which only the aerosol effects were computed, over the entire range of the aerosol model from 0.2 to 40 micrometers. This continuous plot shows very clearly how the aerosol absorption and scattering effects decrease markedly outside the visible region of the spectrum. The spectrum was computed over a vertical path extending from the ground to the top of the atmosphere, with a midlatitude summer atmosphere model and rather clear visibility (not much haze). The wavelength grid used was very fine, so that the spectrum shows all details clearly.

TESTING OF PROGRAM

The MICTRA routine was initially tested by comparing its computed transmission values at selected wavelengths with hand calculated points; in all cases the results agreed very closely.

The routine was then tested against an independent program developed by another group at the Goddard Space Flight Center (Reference 10). This program takes account of all lines within 20 wavenumbers of the frequency at which the transmission is being computed, and it uses a Voigt line profile. Transmission spectra were calculated, using both programs, in the 5 micrometer region of the spectrum, from 2045 to 2057 wavenumbers, using the 1982 version of the Air Force absorption line tape. Both programs computed the spectrum with good resolution: every 0.01 wavenumber for the other program and every 0.04 wavenumber for MICTRA. The two calculated spectra were virtually identical; any slight differences in transmission values, at most a few percent, could be attributed to the fact that the two programs were completely independent and thus could not give identical results.

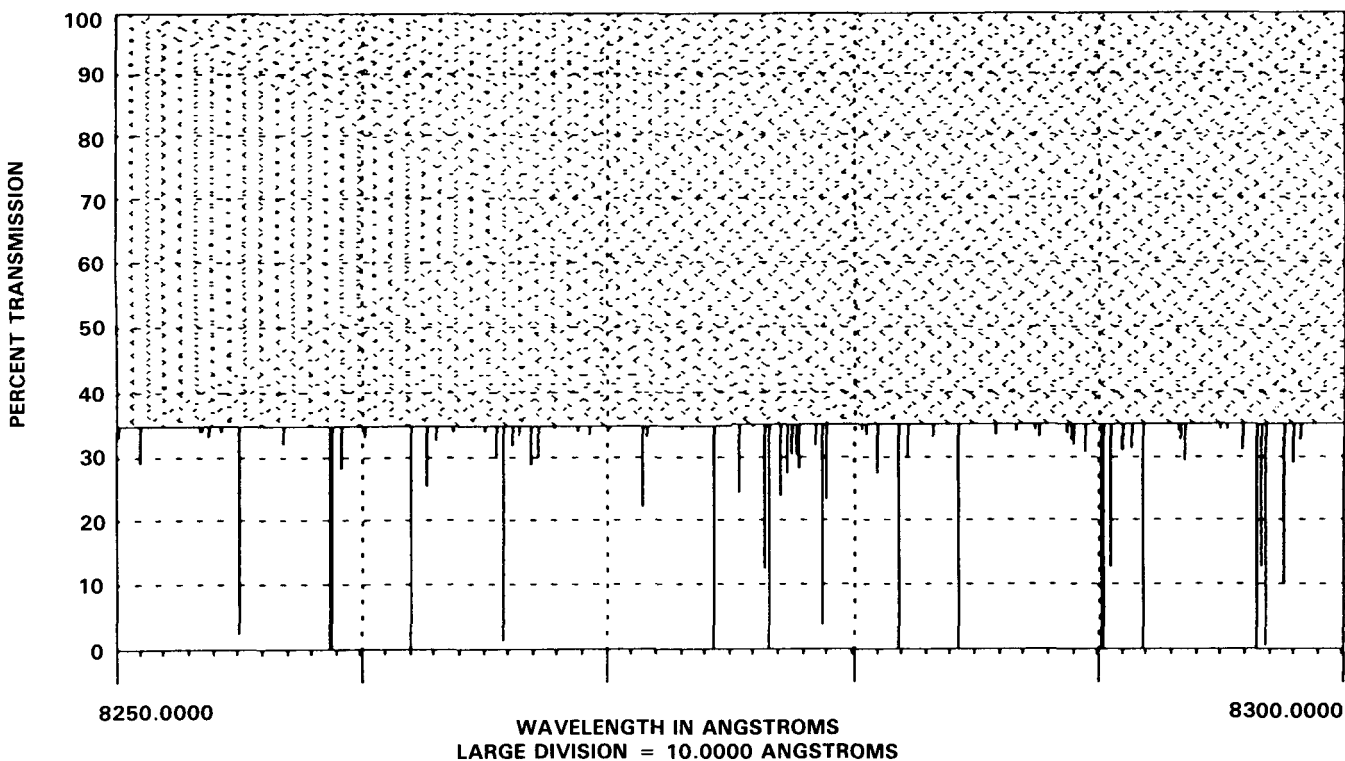


Figure 2. Line Center Spectrum in Near Infrared.

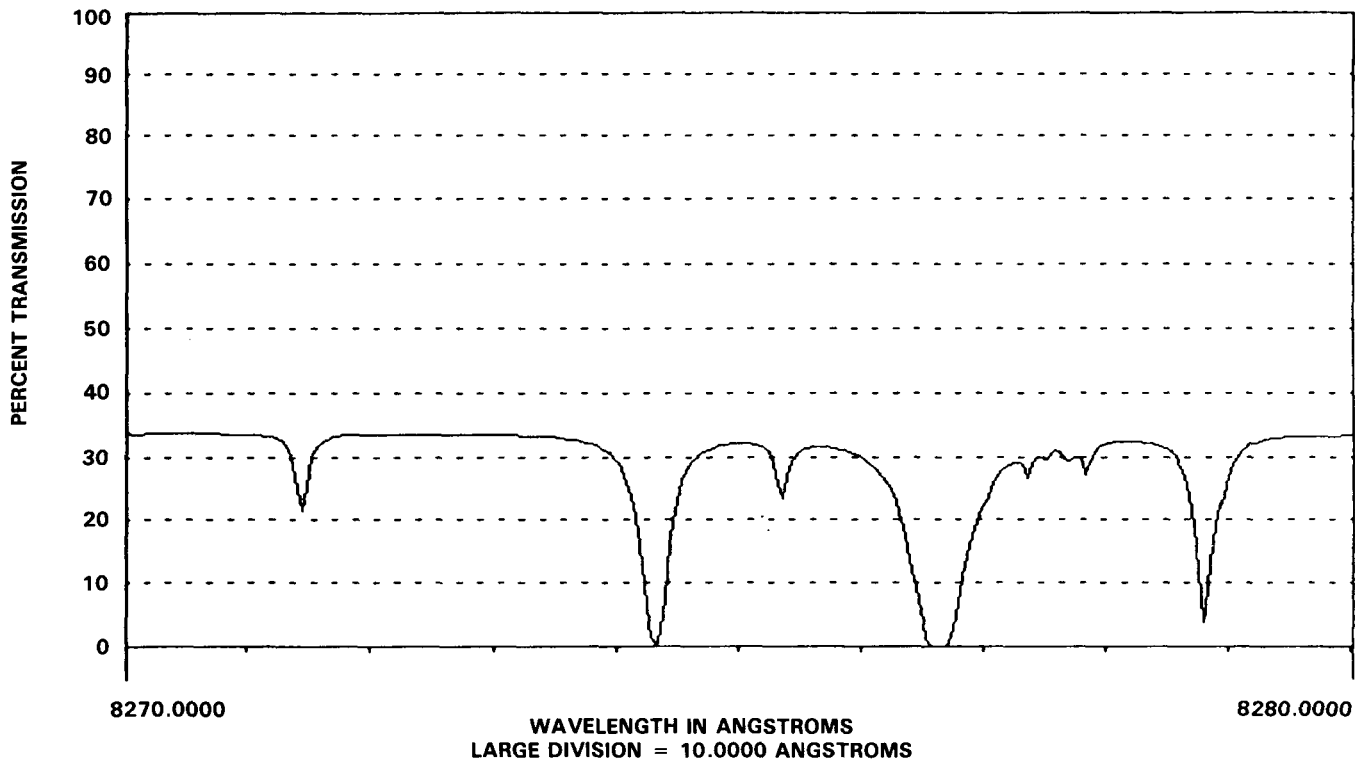


Figure 3. Continuous Plot of Sub-region of Figure 2.

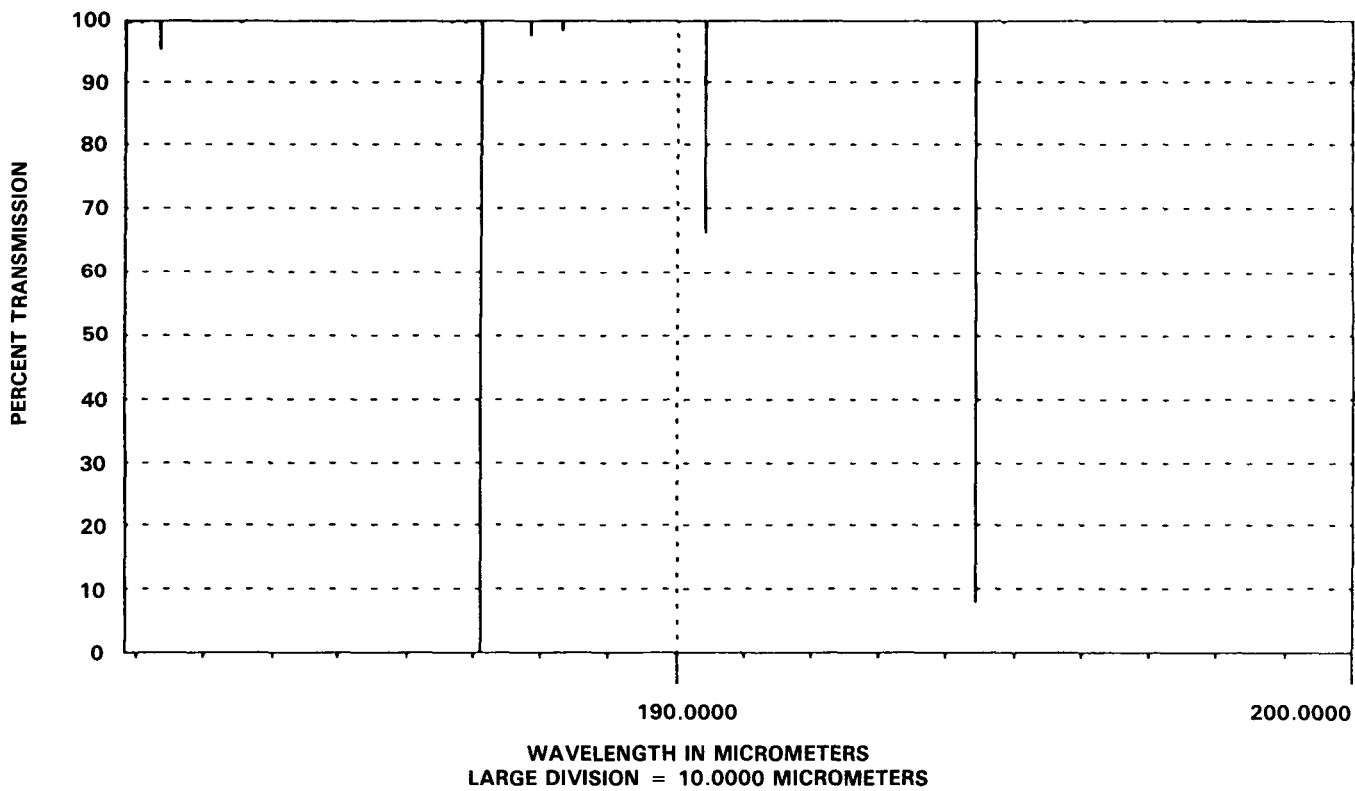


Figure 4. Line Center Spectrum in Submillimeter Region; Water Vapor Lines Only.

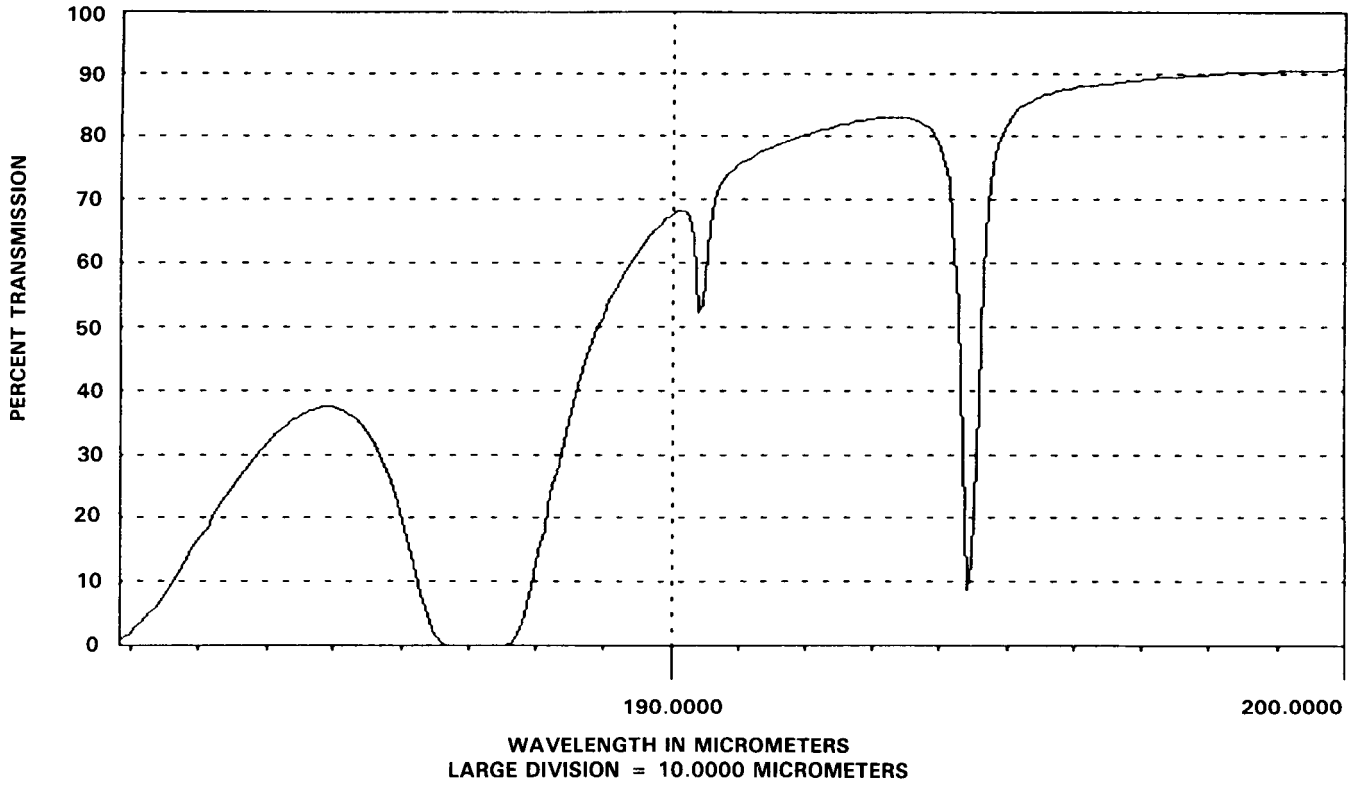


Figure 5. Continuous Plot of the Spectrum in Figure 4.

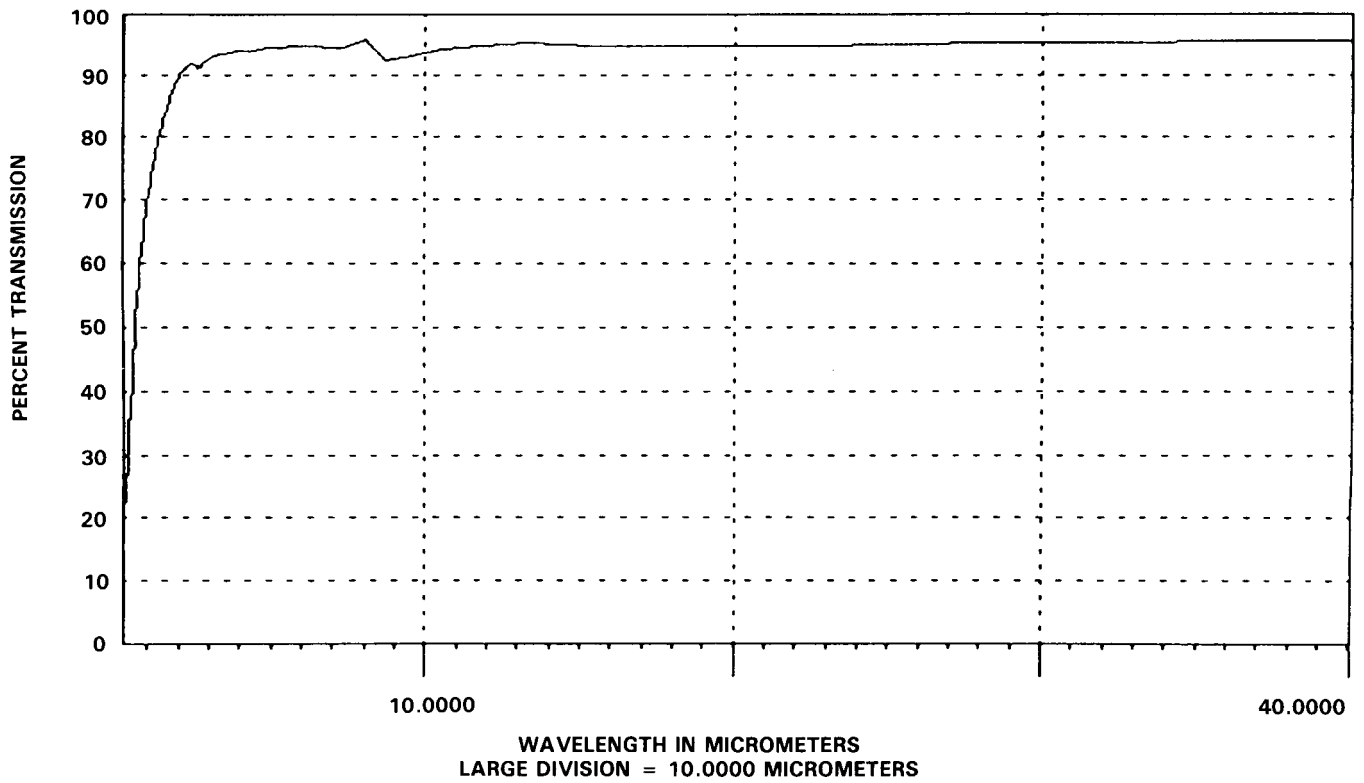


Figure 6. Aerosol Absorption and Scattering Over a Vertical Path Through the Atmosphere, From the Visible to 40 Micrometers.

PROGRAM LISTINGS

PROGRAM MICTRA

C
C This version of the MICTRA routine is designed to run in the BATCH mode;
C the purpose is to be able to calculate many transmission spectra with no
C operator attention.

C
C Three modes of operation are provided:

- C
C 1. Setup mode - the routine is run interactively to create
C a control file for the BATCH program. The
C file includes all input to MICTRA which is
C normally input interactively at the keyboard,
C using ACCEPT statements.
- C
C 2. Batch mode - the routine is run under the control of the
C BATCH program to calculate the spectra,
C using the control file set up in a previous
C run in the setup mode.
- C
C 3. Combined mode - the routine creates a BATCH file and executes
C it to compute any number of spectra, all in
C one run.

C
C This routine is mainly a driver for the subroutines that actually
C do the calculations of the transmission spectrum. It contains the
C following options:

- C
C * The following data, options and parameters may be specified:
- C * Atmosphere and aerosol models;
 - C * Sea-level temperature and pressure;
 - C * Sea-level concentrations of water vapor, CO2 and
C ozone;
 - C * Absorption line file;
 - C * Spectral range of the calculated transmission spectrum;
 - C * "Overview" spectrum (large spectral range) showing
C transmission at line centers only, with aerosol and
C molecular continuum effects also shown, or "continuous"
C spectrum (small spectral range) showing line shapes;
 - C * Wavenumber increment between calculated points
C (may be as small as 0.01 cm⁻¹, or even less);
 - C * Width of the "window" over which absorption lines are
C included in the calculation of the absorption at the
C window center.
- C
C * The spectra are not plotted as they are calculated, because BATCH
C takes so much space. (The spectra can be later plotted from the
C disk spectrum files, using PLOTSP.)
- C
C * The computed spectra are saved on disk files.
- C
C * A separate routine (PLOTSP) is available to plot the spectra.

C*****

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C                                     SPECIFICATION STATEMENTS
C                                     *****
C   BYTE      reply, linfil, spcfil, edmode, disp, ff,
*             htunit, pthtyp, batmod, batmd,
*             grdmod, eqspce, centrs, replys, replyp
C   INTEGER   especs, pcount
C   REAL*4    lownu, lastwn, molnam
C   DIMENSION linfil(15), spcfil(15), keepsw(9), molnam(3)

C   COMMON    z(16),p(3,16),
*             t(3,16),w(3,3,16),cay(3,16),wg(3,3),
*             ta2(16),fa(6,61),asc(6,61),aab(6,61),
*             has(16)

C   COMMON/BLK1/  dnu(16),chi(16),hz1(16),hz2(16),itp,jt,ksam

C   COMMON/RSETUP/ bound,lownu,highnu,delnu,v,
*                 iatmod,iarmdg,iarmdt,iarmds,
*                 numpts,nlines,keeps,iblayr,itlayr,
*                 irsize, hzint(16),iaermd

C   DATA      molnam(1)/'H2O '/, molnam(2)/'CO2 '/, molnam(3)/'O3 '/
C   DATA      ff/'014/, linfil/15*'000/, spcfil/15*'000/, batmd/'C'/
C   DATA      eqspce/'E'/, centrs/'C'/
C   DATA      C/2.997925E10/

C *****
C                                     RUN SETUP
C                                     *****
C   SET UP MODE OF OPERATION: COMMAND FILE SETUP OR CALCULATION UNDER BATCH
C   ////////////////////////////////////////////////////////////////////
C
C   type 4
C   4 format(//t2,'Is this run to set up a BATCH command file'/
*       t2,'or to calculate spectra under BATCH control? (F or C) ', $)
C   accept 5, batmod
C   5 format(a1)

C   IF (batmod .eq. 'F') go to 6
C   IF (batmod .eq. 'C') go to 9

C   Open BATCH command file for input
C   -----

C   6 OPEN (unit=9, name='SD0:BATFIL.BAT', type='NEW', disp='KEEP',
*       form='FORMATTED', initialsize=200)

C   write (9,8)
C   8 format('$JOB'/'$RUN SD2:MICTRA.SAV'/'$DATA')
C   write (9,5) batmd

C   OPEN FILE TO HOLD CALCULATED TRANSMISSION SPECTRUM
C   ////////////////////////////////////////////////////////////////////
C
C   9 type 2
C   2 format(//t2,'Type name of file (including device) to store'/
*       t2,'the computed transmission spectrum: ', $)
C   accept 3, spcfil
C   3 format(15a1)
C   IF (batmod .eq. 'F') write(9,3) spcfil

```

```

        IF (BATMOD .EQ. 'C')
        * OPEN (unit=8, name=spcfil, type='NEW', disp='KEEP',
        *       access='SEQUENTIAL', form='UNFORMATTED', recordsize=2,
        *       initialsize=300)
C
C
C   CHOOSE PATH OVER WHICH TRANSMISSION IS TO BE COMPUTED
C   ///////////////////////////////////////////////////
C
        type 210
210  format(/t2,'Set up path over which transmission is to',
        *      t44,'be calculated: '//)
C
C       Heights and path length in feet or meters?
C       -----
C
214  type 211
211  format(t10,'Specify heights above sea level and path',
        *      t51,'length'/
        *      t10,' in feet or meters? (F or M): ', $)
        accept 212, htunit
212  format(a1)
        IF (batmod .eq. 'F') write(9,212) htunit
C
        IF (htunit .eq. 'F') htconv = 3.048E-04      ! convert ft to km
        IF (htunit .eq. 'M') htconv = 1.E-03        ! convert m to km
C
        IF (htunit .ne. 'F' .and. htunit .ne. 'M') go to 214 ! error corr.
C
        Is the path horizontal, vertical or slanted?
C       -----
C
215  type 213
213  format(/t10,'Is the path horizontal, vertical or slanted?',
        *      t54,' (H,V or S): ', $)
        accept 212, pthtyp
        IF (batmod .eq. 'F') write(9,212) pthtyp
C
        IF (pthtyp .eq. 'H') go to 220
        IF (pthtyp .eq. 'V' .OR. pthtyp .eq. 'S') go to 240
        go to 215                                     ! error corr.
C
        Setup for horizontal path
C       -----
C
220  type 221
221  format(/t15,'The path is horizontal;'/
        *      t15,'we need the following information: '//
        *      t20,'Height of path (ft or m, <15 km): ', $)
        accept 222, pathht
222  format(f10.3)
        IF (batmod .eq. 'F') write(9,222) pathht
        phtkm = htconv * pathht                      ! path height in km
C
        type 223
223  format(/t20,'Length of path (ft or m): ', $)
        accept 222, pathln
        IF (batmod .eq. 'F') write(9,222) pathln
        plnkm = htconv * pathln                      ! path length in km
C
        call laynum (phtkm, layer, delht)            ! get atm. layer number
        iblayr = layer
        itlayr = layer                               !, (top and bottom same)
        go to 12

```

```

C      Setup for vertical or slant path
C      -----
C
240     type 241
241     format(/t15,'The path is vertical or slanted;'/
*         t15,'we need the following information: '//
*         t20,'Height of lower end of path (ft or m, <15 km): ', $)
        accept 222, phtbot
        IF (batmod .eq. 'F') write(9,222) phtbot
C
        type 242
242     format(/t20,'Height of upper end of path (ft or m, <15 km): ', $)
        accept 222, phttop
        IF (batmod .eq. 'F') write(9,222) phttop
C
        secant = 1.0                                ! set sec to 1 for vert. path
        IF (phttyp .eq. 'V') go to 245
C
        type 243                                ! define sec for slant path
243     format(/t20,'Zenith angle of path, in degrees: ', $)
        accept 244, zenang
244     format(f7.3)
        IF (batmod .eq. 'F') write(9,244) zenang
        zangrd = (3.141593/180.) * zenang          ! zenith angle in radians
        secant = 1./cos(zangrd)
C
245     phbkm = htconv * phtbot                  ! ht. of bottom of path (km)
        phtkm = htconv * phttop                 ! ht. of top   of path (km)
        call laynum (phbkm, layer, delht)       ! bottom layer no. and segment
        iblayr = layer
        botseg = 1. - delht
        IF (delht .eq. 0.) botseg = 0.
C
        call laynum (phtkm, layer, delht)      ! top layer no. and segment
        itlayr = layer
        topseg = delht
        IF (delht .eq. 0.) topseg = 1.
C
C      READ IN ALL ATMOSPHERIC MODELS, INCLUDING AEROSOL MODELS, FROM DATA SET 2
C      ////////////////////////////////////////////////////////////////////
C
12     call setup                                ! set up all atmosphere models
C
C      CHOOSE ATMOSPHERE AND AEROSOL MODELS
C      ////////////////////////////////////////////////////////////////////
C
        Atmosphere Model
C      -----
C
        type 20
20     format('0',t2,'which of the 3 atmosphere models do you want?'/
*         t10,'type: 1 for midlatitude summer model: '/
*         t17,'2 for midlatitude winter model: '/
*         t17,'3 for U.S. Standard model: ', $)
        accept 21, iatmod
21     format(i1)
        IF (batmod .eq. 'F') write(9,21) iatmod

```



```

C      Aerosol Model
C      -----
C
C      TYPE 4000
4000  FORMAT(//T2,'The sea-level meteorological range for the'/
*      t2,'aerosol models may be chosen from 5 km'/
•      t2,'(very hazy) to 50 km (very clear);'/
•      t2,'type the desired range in km: ', $)
      ACCEPT 4001, SLVIS
4001  FORMAT(F6.2)
      IF (BATMOD .EQ. 'F') WRITE(9,4001) SLVIS
C
      DO 4010 I=7,16                                ! construct interpolated
      FINT = (50.-SLVIS)/45.                          ! altitude scaling fac-
      HZINT(I) = HZ1(I) + FINT*(HZ2(I)-HZ1(I))        ! tors for aer. models
4010  CONTINUE                                       ! (Only below 9 km)
C
C                                     Note: HZ1 = clear (50 km) scaling factors
C                                     HZ2 = hazy ( 5 km) scaling factors
C                                     (stored in COMMON/BLK1/)
      type 22
22  format(//t2,'Does any part of the path lie between'/
*      t2,'0 km (sea-level) and 2 km altitude? (Y or N): ', $)
      accept 23, reply
23  format(a1)
      IF (batmod .eq. 'F') write(9,23) reply
      IF (reply .eq. 'Y') go to 24
      go to 27
C
24  type 25
25  format('0',t2,'which boundary-layer (0-2 km)',
•      t32,'aerosol model do you want?'/
*      t10,'type: 1 for rural model;'/
•      t16,      '2 for tropospheric model;'/
*      t16,      '3 for maritime model;'/
*      t16,      '4 for urban model: ', $)
      accept 26, iarmdg
26  format(i1)
      IF (batmod .eq. 'F') write(9,26) iarmdg
C
27  type 28
28  format(//t2,'Does any part of the path lie between'/
*      t2,'2 and 9 km altitude? (Y or N): ', $)
      accept 23, reply
      IF (batmod .eq. 'F') write(9,23) reply
      IF (reply .eq. 'Y') go to 29
      go to 31
C
29  IARMDT = 2
C
31  type 32
32  format(//t2,'Does any part of the path lie above 9 km altitude?',
•      t53,'(Y or N): ', $)
      accept 23, reply
      IF (batmod .eq. 'F') write(9,23) reply
      IF (reply .eq. 'Y') go to 33
      go to 39

```

```

33 type 34
34 format(//t2,'Which stratospheric aerosol model do you want?'/
*      t10,'type: 5 for clear (background stratospheric)'/
•      t16, '6 for hazy (aged volcanic): ', $)
accept 26, iarmds
IF (batmod .eq. 'F') write(9,26) iarmds
C
IF (IARMDS .EQ. 5) GO TO 4020
IF (IARMDS .EQ. 6) GO TO 4030
C
4020 DO 4025 I=1,6
HZINT(I) = HZ1(I)
4025 CONTINUE
GO TO 39
4030 DO 4035 I=1,6
HZINT(I) = HZ2(I)
4035 CONTINUE
C
C
C MODIFY ATMOSPHERIC MODEL (AT SEA LEVEL ONLY)
C ///////////////////////////////////////////////////
C
39 type 40
40 format(//t2,'The following are sea-level values of some parameters'//
*      t2,'in the atmosphere model chosen; any of them may be changed: '//)
C
C Modify Sea-level Temperature
C -----
C      tfahr = (9.0/5.0) * t(iatmod,16) - 459.67 ! convert t to Fahr.
C
C      type 41, tfahr
41 format(t2,'Temperature (deg. F) - ',t30,f7.2,
*      t45,'change? (Y or N): ', $)
accept 42, reply
42 format(a1)
IF (batmod .eq. 'F') write(9,42) reply
C
IF (reply .eq. 'N') go to 45
C
type 43
43 format(t15,'New value:',t29,' ', $)
accept 44, tfahr
44 format(f7.2)
IF (batmod .eq. 'F') write(9,44) tfahr
C
t(iatmod,16) = (5./9.)*(tfahr+459.67) ! convert back to deg. K
C
C Modify Sea-level Pressure
C -----
C
45 type 46, p(iatmod,16)
46 format(/t2,'Pressure (millibars) - ',t30,f8.3,
•      t45,'change? (Y or N): ', $)
C
IF (batmod .eq. 'C') write (8,47) p(iatmod,16)
accept 42, reply
IF (batmod .eq. 'F') write(9,42) reply
C
IF (reply .eq. 'N') go to 48
C
type 43
accept 47, p(iatmod,16)
47 format(f8.3)
IF (batmod .eq. 'F') write(9,47) p(iatmod,16)

```

```

C   Modify Sea-level Concentration of Water Vapor
C   -----
C
48  type 49, wg(iatmod,1)
49  format(/t2,'Water vapor conc. (molecules/cm**2/km) - ',
*      t46,e10.2,t61,'change? (Y or N): ',,$)
C      IF (batmod .eq. 'C') write (8,51) wg(iatmod,1)
      accept 42, reply
      IF (batmod .eq. 'F') write(9,42) reply
C
      IF (reply .eq. 'N') go to 52
C
      type 50
50  format(t33,'New value:',t45,' ',,$)
      accept 51, wg(iatmod,1)
51  format(e10.2)
      IF (batmod .eq. 'F') write(9,51) wg(iatmod,1)
C
C   Modify Sea-level Concentration of Carbon Dioxide
C   -----
C
52  type 53, wg(iatmod,2)
53  format(/t2,'Carbon dioxide conc. (molecules/cm**2/km) - ',
*      t46,e10.2,t61,'change? (Y or N): ',,$)
      accept 42, reply
      IF (batmod .eq. 'F') write(9,42) reply
C
      IF (reply .eq. 'N') go to 54
C
      type 50
      accept 51, wg(iatmod,2)
      IF (batmod .eq. 'F') write(9,51) wg(iatmod,2)
C
C   Modify Sea-level Concentration of Ozone
C   -----
C
54  type 55, wg(iatmod,3)
55  format(/t2,'Ozone conc. (molecules/cm**2/km) - ',
*      t46,e10.2,t61,'change? (Y or N): ',,$)
      accept 42, reply
      IF (batmod .eq. 'F') write(9,42) reply
C
      IF (reply .eq. 'N') go to 69
C
      type 50
      accept 51, wg(iatmod,3)
      IF (batmod .eq. 'F') write(9,51) wg(iatmod,3)
C
SET UP MASTER LINE FILE FOR THIS RUN (UNformatted, sequential access)
////////////////////////////////////
C
C   This master line file should contain ONLY H2O, CO2 and O3 lines (all
C   isotopes) in the region of the spectrum being investigated.
C
C   Each record in the file should contain the following quantities, in
C   UNformatted (i.e., binary) form:
C
C           Wavenumber           real*4
C           Line strength         "
C           Line half-width      "
C           Lower state energy   "
C           Isotope identification integer (coded)
C           Molecular species    integer (1, 2 or 3)

```

```

C   The exact definitions and units of these quantities are given in the
C   documentation for the Air Force Line Catalogue.
C
C   A line file in this format is called a "condensed" line file, to
C   distinguish it from the original segment of line data taken from the
C   AFGL tapes, which list the lines in ASCII (i.e., formatted) form.
C   A linefile editing routine is available to "condense" desired
C   segments of AFGL line data.
C
69  type 70
70  format('0',t2,'Type the name of the line file to be used: ',,$)
    accept 71, linfil
71  format(15A1)
    IF (batmod .eq. 'F') write(9,71) linfil
C
    OPEN (unit=4, name=linfil, type='OLD', form='UNFORMATTED',
*      disp='KEEP')
C
C   COUNT NUMBER OF LINES IN MASTER LINE FILE,
C   NOTE FIRST AND LAST WAVENUMBERS AND
C   DISPLAY THIS INFORMATION ON SCREEN
C   (in both frequency and wavelength units)
C   //////////////////////////////////////
C
    type 710
710 format(//t2,'Is this LINFIL the same as the one used for the'/
*      t2,'immediately preceding spectrum? (Y or N; type N'/
*      t2,'if this is the first spectrum for this run): ',,$)
    accept 711, reply
711 format(a1)
    IF (batmod .eq. 'F') write(9,711) reply
C
    IF (reply .eq. 'Y') go to 726
    read(4) wavnum, strnth, hwidth, energy, isotop, speces
    frstwn = wavnum
    rewind 4
C
    DO 722 linct=1,30000
722  read(4, end=725) wavnum,strnth,hwidth,energy,isotop,speces
C
725  lastwn = wavnum
    mlines = linct-1
    rewind 4
C
    end1hz = c*lastwn
    end2hz = c*frstwn
    end1a = 1.e8/lastwn
    end2a = 1.e8/frstwn
    end1mc = 1.e4/lastwn
    end2mc = 1.e4/frstwn
C
726  type 727, mlines, lastwn,frstwn, end1hz,end2hz,
*      end1a, end2a, end1mc,end2mc
727  format(//t2,'This line file contains',t34,I5,t40,'lines;'/
*      t2,'it covers the frequency-wavelength range: '//
*      t30,'FROM',t60,'TO'//
*      t5,'wavenumbers:',t25,f14.4,t53,f14.4//
*      t5,'freq. in hertz:',t24,e16.9,t52,e16.9//
*      t5,'angstroms:',t24,f15.4,t52,f15.4//
*      t5,'micrometers:',t26,f12.4,t54,f12.4//)

```

```

C   SPECIFY WAVENUMBER RANGE OVER WHICH TRANSMISSION IS TO BE CALCULATED
C   ////////////////////////////////////////////////////////////////////
C
      type 730
730  format(//t2,'Type the endpoints, in either order, of the'/
*      t2,'spectral range over which the transmission is'/
*      t2,'to be computed; the endpoints may be expressed'/
*      t2,'in wavenumbers, hertz, angstroms or micrometers;'/
*      t2,'use the E format (real). '//
*      t2,'Type: 1 for endpoints in wavenumbers'/
*      t2,'      2 "      "      " hertz'/
*      t2,'      3 "      "      " angstroms'/
*      t2,'      4 "      "      " micrometers: ', $)

C
      accept 731, iunit
731  format(i1)
      IF (batmod .eq. 'F') write(9,731) iunit

C
      type 732
732  format(//t10,'One endpoint: ', $)
      accept 733, highnu
733  FORMAT(E18.6)
      IF (batmod .eq. 'F') write(9,733) highnu
      IF (iunit .eq. 1) highnu = highnu      ! convert to wavenumbers
      IF (iunit .eq. 2) highnu = highnu/c
      IF (iunit .eq. 3) highnu = 1.E8/highnu
      IF (iunit .eq. 4) highnu = 1.E4/highnu

C
      type 734
734  format(//t10,'Other endpoint: ', $)
      accept 733, lownu
      IF (batmod .eq. 'F') write(9,733) lownu

C
      IF (iunit .eq. 1) lownu = lownu      ! convert to wavenumbers
      IF (iunit .eq. 2) lownu = lownu/c
      IF (iunit .eq. 3) lownu = 1.E8/lownu
      IF (iunit .eq. 4) lownu = 1.E4/lownu

C
      IF (lownu .lt. highnu) go to 1999      ! interchange endpoints
      hghnu = highnu                        ! (if necessary)
      highnu = lownu
      lownu = hghnu

C
C   CHOOSE EQUALLY SPACED OR LINE CENTER GRID OF WAVELENGTH PTS
C   ////////////////////////////////////////////////////////////////////
C
1999 TYPE 2000
2000 FORMAT(//T2,'Do you want to compute the spectrum: '//
*      t5,'o Over an equally spaced grid of wavenumbers (type E)'/
*      t5,'o Only at the absorption line centers (type C) '//
*      t2,'(If no absorption lines are to be used, you MUST type E): ', $)
      ACCEPT 2001, GRDMOD
2001 FORMAT(A1)
      IF (BATMOD .EQ. 'F') WRITE(9,2001) GRDMOD
      IF (GRDMOD .EQ. EQSPCE) GO TO 79
      IF (GRDMOD .EQ. CENTRS) GO TO 89

```

```

79 type 80
80 format('0',t2,'Type the wavelength increment (in wavenumbers)'/
*      t2,'to be used in computing the transmission'/
*      t2,'spectrum: ',,$)
      accept 81, delnu
81 format(f15.4)
      IF (batmod .eq. 'F') write(9,81) delnu
C
      xnmpts = (highnu-lownu)/delnu + 1.          ! corr. no. of points
      IF (xnmpts .lt. 32000.) numpts = xnmpts
      IF (xnmpts .gt. 32000.) numpts = 32000
      type 82, numpts
82 format(/t2,'This wavelength increment corresponds to'/
*      t2,'a total of ', t14, i5, t20, 'points;'/
•      t2,'is this number of points okay? (Y or N): ',,$)
      accept 83, reply
83 format(a1)
      IF (batmod .eq. 'F') write(9,83) reply
      IF (reply .eq. 'N') go to 79
C
C
C SPECIFY WIDTH OF WAVENUMBER WINDOW IN WHICH LINES ARE CONSIDERED
C TO CONTRIBUTE TO THE ABSORPTION AT THE WINDOW CENTER
C ////////////////////////////////////////////////////
C
89 type 90
90 format('0',t2,'Type the HALF-width, in wavenumbers, of the window'/
*      t2,'over which you want lines to be included in calcu'/
*      t2,'lating the absorption at the window center: ',,$)
C
      accept 91, bound
91 format(f12.4)
      IF (batmod .eq. 'F') write(9,91) bound
C
C EXAMINE WAVENUMBER RANGE AND WINDOW WIDTH;
C IF RANGE, EXTENDED ABOVE AND BELOW BY WINDOW HALF-WIDTH,
C IS NOT CONTAINED IN MASTER LINE FILE,
C GO BACK AND RE-DEFINE THEM (UNLESS this spectrum is to show
C ONLY aerosol and/or molecular continuum effects, with NO
C absorption lines included - query user for this information)
C ////////////////////////////////////////////////////
C
      ibotsw = 0
      itopsw = 0
C
      IF ((lownu-bound) .lt. frstwn)  ibotsw = -1
      IF ((highnu+bound) .gt. lastwn) itopsw = -1
      isumsw = ibotsw + itopsw
C
      IF (isumsw .eq. 0) go to 94
C
      IF(ibotsw .eq. -1) type 92
92 format(/t2,'lownu-bound falls below the first line'/
*      t2,'of the master line file - ')
      IF(itopsw .eq. -1) type 93
93 format(/t2,'highnu+bound falls above the last line'/
*      t2,'of the master line file - ')
      type 935
935 format(/t2,'Is this spectrum to show ONLY aerosol and/or'/
*      t2,'molecular continuum effects, with NO absorption'/
•      t2,'lines included? (Y or N): ',,$)
      accept 936, reply
936 format(a1)

```

```

IF (batmod .eq. 'F') write (9,936) reply
IF (reply .eq. 'Y') go to 94
type 931
931 format(/t2,'re-define lownu, highnu and/or bound:')
go to 726
94 IF (GRDMOD .EQ. CENTRS) PLTCOD = 1.
IF (GRDMOD .EQ. EQSPCE) PLTCOD = 0.

C
ANGMIC = 0. ! output in angstroms or microns
IF (1.E8/LOWNU .GE. 1.E4) ANGMIC = 1.

C
IF (BATMOD .EQ. 'C') WRITE(8) PLTCOD,ANGMIC
IF (BATMOD .EQ. 'C' .AND.
* ANGMIC .EQ. 0.) WRITE(8) 1.E8/HIGHNU,1.E8/LOWNU ! ang
IF (BATMOD .EQ. 'C' .AND.
* ANGMIC .EQ. 1.) WRITE(8) 1.E4/HIGHNU, 1.E4/LOWNU ! mic

C
C
C "NON-PHYSICAL" SELECTION OF LINE, CONTINUUM OR AEROSOL EXTINCTION
C ///////////////////////////////////////////////////
C

type 121
121 format(//t2,'The user may want to arbitrarily select or eliminate'//
* t2,'the molecular absorption lines of a given species,'//
* t2,'Rayleigh (molecular) scattering, any of the three'//
* t2,'molecular continua, aerosol absorption or aerosol'//
* t2,'scattering; to allow this, the following options'//
* t2,'are provided; provision is made later for more'//
* t2,'flexible selection of absorption lines.'//)
type 122
122 format(t2,'Eliminate or retain all water lines (E or R): ',)$)
accept 123, reply
123 format(a1)
IF (batmod .eq. 'F') write(9,123) reply
keeps(1) = 1
IF (reply .eq. 'E') keeps(1) = 0

C
type 124
124 format(t2,
* 'Eliminate or retain all carbon dioxide lines (E or R): ',)$)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(2) = 1
IF (reply .eq. 'E') keeps(2) = 0

C
type 125
125 format(t2,'Eliminate or retain all ozone lines (E or R): ',)$)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(3) = 1
IF (reply .eq. 'E') keeps(3) = 0

C
type 126
126 format(t2,'Eliminate or retain Rayleigh scattering (E or R): ',)$)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(4) = 1
IF (reply .eq. 'E') keeps(4) = 0

```

```

type 127
127 format(t2,
* 'Eliminate or retain 3.5-4.2 micron H2O continuum (E or R): ', $)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(5) = 1
IF (reply .eq. 'E') keeps(5) = 0
C
type 128
128 format(t2,
* 'Eliminate or retain 8-14 micron H2O continuum (E or R): ', $)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(6) = 1
IF (reply .eq. 'E') keeps(6) = 0
C
type 129
129 format(t2, 'Eliminate or retain nitrogen continuum (E or R): ', $)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(7) = 1
IF (reply .eq. 'E') keeps(7) = 0
C
type 130
130 format(t2, 'Eliminate or retain aerosol absorption (E or R): ', $)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(8) = 1
IF (reply .eq. 'E') keeps(8) = 0
C
type 131
131 format(t2, 'Eliminate or retain aerosol scattering (E or R): ', $)
accept 123, reply
IF (batmod .eq. 'F') write(9,123) reply
keeps(9) = 1
IF (reply .eq. 'E') keeps(9) = 0
C
SET UP THE LINE FILE FOR THIS RUN, BY TAKING THE APPROPRIATE LINES
FROM THE MASTER LINE FILE ENTERED ABOVE
////////////////////////////////////
C
Open the run line-file (unformatted, sequential access)
-----
C
OPEN (unit=3, name='SCR:RUNFIL.DAT', type='NEW', disp='DELETE',
* form='UNFORMATTED', recordsize=6, initialsize=974)
C
Set up interactive screening of lines
-----
type 961
961 format(//t2, 'Do you want to interactively screen the lines'/
* t2, 'before they are written to RUNFIL? (Y or N): ', $)
accept 962, reply
962 format(a1)
IF (batmod .eq. 'F') write(9,962) reply
C
iusedt = 0
IF(reply .eq. 'Y') iusedt = 1
C
Initialize line counts
-----
nlines = 0 ! init. count of lines read to RUNFIL
iread = 0 ! " " " " " from LINFIL

```



```

C      Read a line from LINFIL
C      -----
C
110  READ(4, end=200) wavnum, strnth, hwidth, energy, isotop, speces
C      *****
C
C      Automatic screening
C      -----
C      IF (wavnum .lt. (lownu-bound)) go to 110
C      IF (wavnum .gt. (highnu+bound)) go to 200
C      IF (keeps(w(speces) .eq. 0) go to 110
C
C      Interactive screening
C      -----
C      IF (iusedt .eq. 0) go to 990
C
C      iread = iread + 1
C      IF (iread .eq. 1) go to 970
C      go to 975
C
970  type 971
971  format(/t2,'Which mode do you want to use up to the next stop?')
*      t12,'R - reject all lines up to next stop'
*      t12,'I - include " " " " " " '
*      t12,'E - examine each line " " " " '
*      t2,'Type R, I or E: ',)
      accept 962, edmode
      IF (batmod .eq. 'F') write(9,962) edmode
C
      type 972
972  format(/t2,'Type next wavenumber stop: ',)
      accept 973, wnstop
973  format(f10.4)
      IF (batmod .eq. 'F') write(9,973) wnstop
C
      go to 980
C
975  IF (wavnum .le. wnstop) go to 980
C
      type 971
      accept 962, edmode
      IF (batmod .eq. 'F') write(9,962) edmode
      type 972
      accept 973, wnstop
      IF (batmod .eq. 'F') write(9,973) wnstop
C
980  IF (edmode .eq. 'R') go to 110
      IF (edmode .eq. 'I') go to 990
C
      type 385, molnam(speces), isotop, wavnum, strnth, hwidth, energy
      type 981
981  format(/t2,'Type R to reject this line'
*      t2,'Type I to include it: ',)
      accept 962, disp
      IF (batmod .eq. 'F') write(9,962) disp
C
      IF (disp .eq. 'R') go to 110

```

```

C      Write this line (which is retained if this point is reached)
C      into RUNFIL
C      -----
C 990  write(3) speces, isotop, wavnum, strnth, hwidth, energy
C
C      Keep a running count of the number of lines;
C      go back and read the next line
C
C      nlines = nlines + 1
C
C      go to 110
C      *****
C
C      If RUNFIL contains no lines, decide whether to continue the run
C      -----
C
C 200  IF (nlines .eq. 0) go to 201
C      go to 380
C
C 201  type 202
C 202  format(///t2,'No absorption lines in RUNFIL;'/
C      *      t2,'continue run? (Y or N): ', $)
C      accept 203, reply
C 203  format(a1)
C      IF (batmod .eq. 'F') write(9,203) reply
C
C      IF (reply .eq. 'N') STOP
C      go to 399
C
C      If RUNFIL contains any lines, do you want to display them
C      on the screen and/or print them?
C      -----
C 380  type 381, nlines
C 381  format(//t2,'There are',t12,I5,t18,'lines in RUNFIL;'/
C      *      t2,'do you want to display them on the screen? (Y or N): ', $)
C      accept 382, replys
C 382  format(a1)
C      IF (batmod .eq. 'F') write(9,382) replys
C      type 386
C 386  format(t2,'do you want to print them?'/
C      *      t2,'(Turn on the printer!) (Y or N): ', $)
C      accept 382, replyp
C      IF (batmod .eq. 'F') write(9,382) replyp
C
C      IF (replys .eq. 'N' .AND.
C      *      replyp .eq. 'N') go to 399
C
C      Display and/or print lines in RUNFIL
C      -----
C      rewind 3
C
C      DO 384 I=1,nlines
C      read(3) speces,isotop,wavnum,strnth,hwidth,energy
C      IF (replys .eq. 'Y' .and. I .eq. 1) type 388
C      IF (replyp .eq. 'Y' .and. I .eq. 1) print 388
C      IF (replys .eq. 'Y')
C      * type 385, molnam(speces),isotop,wavnum,strnth,hwidth,energy
C      IF (replyp .eq. 'Y')
C      * print 385, molnam(speces),isotop,wavnum,strnth,hwidth,energy
C 384  continue
C
C      rewind 3

```

```

388 format(//t2,'species',t10,'isotope',t18,'wavenumber',
*          t33,'strength',t48,'halfwidth',t60,'l.s. energy')
385 format(/t2,a4,t10,i4,t18,f10.4,t33,e10.3,t48,f8.4,t60,f10.3)
C
C      Compute number of blocks required to store RUNFIL
C      -----
C
399 irsize = IFIX((FLOAT(nlines)/512.)*24.) + 1 ! no. of blocks
C
C      If RUNFIL size is less than 164 blocks, put RUNFIL in VM
C      -----
C
      IF (irsize .GT. 164) go to 1507
C
      OPEN (unit=1, name='VM:RUNFIL.DAT', type='NEW', disp='KEEP',
*          form='UNFORMATTED', recordsize=6, initialsize=irsize)
C
      IF(nlines .eq. 0) go to 1509
      rewind 3
C
      DO 1508, II=1,nlines
      read (3) especes, isotop, wavnum, strnth, hwidth, energy
      write(1) especes, isotop, wavnum, strnth, hwidth, energy
1508 continue
C
1509 CLOSE(unit=3, disp='DELETE')
      CLOSE(unit=1, disp='KEEP')
      OPEN (unit=3, name='VM:RUNFIL.DAT', type='OLD', disp='DELETE',
*          form='UNFORMATTED', recordsize=6)
C
C      Close LINFIL
C      -----
1507 CLOSE(unit=4,disp='KEEP')
C
C      Comments
C      -----
C      The line file for this run is now set up; it contains only those
C      lines for which the concentration is non-zero; the total number of
C      lines in RUNFIL is nlines; the wavenumber range contained in the
C      file is lownu-bound to highnu+bound.
C
C*****
      IF (batmod .eq. 'F') go to 800
C
C
C          CALCULATION
C          *****
C
C      To compute the absorption coefficient at each of the net of points
C      of separation delnu spanning the wavenumber range (lownu, highnu),
C      we compute, at each point of wavenumber v, the sum of the contri-
C      butions from all lines which lie in the window (v-bound, v+bound);
C      for an overview spectrum the net of points consists of just the
C      line centers.
C
C
C      COMPUTE AEROSOL-ONLY TRANSMISSION AT HIGHNU AND LOWNU
C      (for use by PLOTSP in drawing the aerosol line
C      for a line-center plot)
C      //////////////////////////////////////
      IAERMD = 1
      ATL = -1.
      ATU = -1.

```

```

      V = HIGHNU
      GO TO 300
C
2900  IF (ATL .LT. 0.) GO TO 2910
      IF (ATU .LT. 0.) GO TO 2920
C
2910  ATL = TRANS
      V = LOWNU
      GO TO 300
C
2920  ATU = TRANS
      WRITE(8) ATL,ATU
      IAERMD = 0
C
C
C  LOOP OVER NET OF POINTS SPANNING THE SPECTRUM
C  ////////////////////////////////////////////////////////////////////
      IF (GRDMOD .EQ. EQSPCE) GO TO 3000
      IF (GRDMOD .EQ. CENTRS) GO TO 3100
C
3000  v = lownu                ! start at low end of spectrum
      i = 1                    ! index for net of points
      itopsw = 0               ! switch to jump out of loop
      GO TO 300
C
3100  REWIND 3
      LNINDX = 0
310  READ(3)  SPECES,ISOTOP,WAVNUM,STRNTH,HWIDTH,ENERGY
      LNINDX = LNINDX + 1
      IF (WAVNUM .LT. LOWNU) GO TO 310
      FRSTLN = LNINDX
      I = 1
      V = WAVNUM
C
300  call rdtape                ! loop over net of points
      *****                  *****
C
C  Go to section of code for given path type
C  -----
C
      IF (pthtyp .eq. 'H') go to 1000
      IF (pthtyp .eq. 'V' .OR. pthtyp .eq. 'S') go to 1100
C
      Horizontal path
      -----
C
1000  layer = iblayr
      extcof = cay(iatmod,layer) + aerosol(layer)      ! ext. coeff.
      trans = exp(-extcof*plnkm)                       ! transmittance
      waveno = v                                       ! wavenumber
      go to 1400
C
      Vertical or slant path
      -----
C
1100  depth = 0.
C

```

```

DO 1110 layer=itlayr,iblayr
C *****
C l = layer
C extcof = cay(iatmod,l) + aerosol(l)
C pathl = secant
C IF (l .eq. iblayr) pathl = botseg*secant
C IF (l .eq. itlayr) pathl = topseg*secant
C depth = depth + extcof*pathl
1110 continue
C *****
C
C trans = exp(-depth)
C waveno = v

C Store this point in spectrum file (formatted, seq. access)
C -----
1400 IF (IAERMD .EQ. 1) GO TO 2900
C
C IF (angmic .eq. 0.) write(8) 1.e8/waveno, trans ! conv. to ang.
C IF (angmic .eq. 1.) write(8) 1.e4/waveno, trans ! conv. to mic.
C
C Type this point on the screen to show the user
C the progress of the calculation
C -----
C type 1401, waveno,trans
1401 format(t2,e12.5,t18,f8.4)
C
C Test wavenumber to see if high end of range has been reached
C -----
C IF (GRDMOD .EQ. EQSPCE) GO TO 350
C
C REWIND 3
C LNINDEX = 0
351 READ(3,END=800) SPECES,ISOTOP,WAVNUM,STRNTH,HWIDTH,ENERGY
C LNINDEX = LNINDEX + 1
C IF (LNINDEX .EQ. FRSTLN+I) GO TO 355
C GO TO 351
355 I = I + 1
C V = WAVNUM
C WAVNOP = WAVENO
C TRANSP = TRANS
C IF (V .LE. HIGHNU) GO TO 300
C I = I - 1
C GO TO 800

C
C 350 rewind 3
C 356 read(3,end=358) speces,isotop,wavnum,strnth,hwidth,energy
C IF (wavnum .gt. v .and. wavnum .lt. v+delnu) go to 357
C go to 356
357 v = wavnum
C go to 359
358 v = v + delnu
359 i = i + 1

C
C wavnop = waveno ! store these as "previous" values
C transp = trans

C
C IF (v .gt. highnu) go to 3355
C go to 300
C *****
C

```

```

3355  IF (itopsw .eq. 1) go to 3356
      itopsw = 1
      v = highnu
      go to 300
C
C
3356  i = i-1                                ! correct the line count
C
C*****
      800  close(unit=3,disp='DELETE')
          IF (batmod .eq. 'C') close(unit=8,disp='KEEP')
C
      type 1560
1560  format(//t2,'Do you want to compute another spectrum? (Y or N): ', $)
      accept 1561, reply
1561  format(a1)
C
      IF (batmod .eq. 'F') write(9,1561) reply
      IF (reply .eq. 'Y') go to 9
C
      IF (batmod .eq. 'F') write(9,1562)
1562  format('$EOD'/'$EOJ')
      IF (batmod .eq. 'F') CLOSE (unit=9, disp='KEEP')
C
      STOP
      END

```



```

C   PATH      OPTICAL DEPTH AT LINE CENTER FOR 1 - KM SEA LEVEL PATH
C   PBAR      AVERAGE LAYER PRESSURE
C   PH2O      WATER VAPOR PRESSURE
C   PI        RATIO OF CIRCUMFERENCE OF A CIRCLE TO ITS DIAMETER
C   S         ABSORPTION LINE INTENSITY - INPUT FROM TAPE
C   SEC       SECANT ANGLE - INPUT DATA
C   ST        TEMPERATURE CORRECTED LINE INTENSITY
C   T         ATMOSPHERIC LAYER TEMPERATURE - INPUT DATA
C   TA1       CLEAR: AEROSOL ABSORPTION COEFFICIENT - OUTPUT DATA
C   TA2       HAZY: AEROSOL ABSORPTION COEFFICIENT - OUTPUT DATA
C   TBAR      AVERAGE LAYER TEMPERATURE
C   V         FREQUENCY AT WHICH THE EXTINCTION COEFFICIENT ARE BEING
C             CALCULATED - INPUT DATA
C   VBOT      LOWER FREQUENCY LIMIT FOR THE LINE CONTRIBUTIONS.
C             VBOT= V-BOUND
C   VTOP      UPPER FREQUENCY LIMIT FOR THE LINE CONTRIBUTIONS.
C             VTOP= V+BOUND
C   W         ABSORBER CONCENTRATION
C   WBAR      MEAN WATER VAPOR CONCENTRATION FOR A LAYER
C   WG        SEA LEVEL VALUES OF MOLECULAR ABUNDANCES - INPUT DATA
C   WH        WATER VAPOR CONCENTRATION AT A SPECIFIC LEVEL-INPUT DATA
C   WH1       SCALE HEIGHT ASSOCIATED WITH WATER VAPOR
C   WH3       SCALE HEIGHT ASSOCIATED WITH OZONE
C   WO        OZONE CONCENTRATION AT A SPECIFIC LEVEL - INPUT DATA
C   WO3       MEAN OZONE CONCENTRATION FOR A LAYER
C   WV        WAVELENGTH CORRESPONDING TO FREQUENCY, V
C   W1D       INTERMEDIATE QUANTITY ASSOCIATED WITH COMPUTING
C             INTEGRATED WATER AMOUNT
C   W3D       INTERMEDIATE QUANTITY ASSOCIATED WITH COMPUTING
C             INTEGRATED OZONE AMOUNT
C   X1        LORENTZ LINE MODIFICATION VARIABLE
C   Z         ATMOSPHERIC HEIGHT (KM) - INPUT DATA
C   ZDIST     ABSOLUTE DISTANCE FROM THE LINE CENTER FREQUENCY V
C*****

```

```

C             SPECIFICATION STATEMENTS
C             *****

```

```

C   IMPLICIT  INTEGER*2 (I-N)
C   REAL*4    LOWNU
C   DIMENSION ATMODL(3,5), ARSMOD(6,5), WH(3,16), WO(3,16), KEEPSW(9)

```

```

C   COMMON    Z(16),
C   *         P(3,16), T(3,16), W(3,3,16), CAY(3,16), WG(3,3),
C   *         TA2(16), FA(6,61), ASC(6,61), AAB(6,61),
C   *         HAS(16)

```

```

C   COMMON/BLK1/ DNU(16), CHI(16), HZ1(16), HZ2(16), ITP, JT, KSAM
C   COMMON/RSETUP/ BOUND, LOWNU, HIGHNU, DELNU, V,
C   *              IATMOD, IARMDG, IARMDT, IARMDS,
C   *              NUMPTS, NLINES, KEEPSW, IBLAYR, ITLAYR,
C   *              IRSIZE, HZINT(16), IAERMD

```

```

C   DATA     HCONV, O3CON/3.34E+21, 1.255E+21/

```

```

C*****
C

```



```

C READ ATMOSPHERIC MODELS FROM DATA SET 2 AND PROCESS THEM
C ///////////////////////////////////////////////////
C
C OPEN DATA SET 2 (formatted, sequential access)
C -----
C OPEN (UNIT=2, NAME='SD2:ATMMOD.DAT', TYPE='OLD',
* DISP='KEEP', FORM='FORMATTED')
C
C READ SEA LEVEL VALUES OF MOLECULAR ABUNDANCES FROM DATA SET 2;
C STORE THEM IN ARRAY WG
C -----
C DO 1 I=1,3 ! I = atm. model
C READ (2,19) (WG(I,M),M=1,3) ! M = molecule (H2O,CO2,O3)
1 CONTINUE
C
19 FORMAT(3E10.3)
C
C READ THE NAMES OF THE ATMOSPHERIC MODELS AND THE DATA
C -----
C DO 300 J=1,3 ! read names
C *****
C READ (2,21) (ATMODL(J,M),M=1,5)
C
C DO 200 L=1,ITP ! read data (L=1-16)
C *****
C K=ITP-L+1 ! reverse numbering of layers
C (K=16-1)
C READ (2,22) Z(K), P(J,K),T(J,K),WH(J,K),WO(J,K)
200 CONTINUE
C
300 CONTINUE
C *****
C
C Z = height in kilometers
C P = pressure
C T = temperature
C WH = water vapor concentration
C WO = ozone concentration
C
21 FORMAT (5A4)
22 FORMAT (F6.1, E10.3,F6.1,2E10.1)
C
C NOTE: The order of the atmospheric layers is reversed;
C i.e., K = 1-16 runs layers from top to ground.
C (K = 1 AT 15 KM,..., K = 16 at sea level.)
C
C COMPUTE THE AVERAGE LAYER CONCENTRATIONS OF H2O, CO2 AND O3
C -----
C NOTE: Molecular densities are assumed to decrease expo-
C nentially between consecutive levels; we compute
C the average layer value by integrating the expo-
C nential over the layer. If W1D or W3D approach
C zero, the log is poorly defined, so in this case
C we use just a simple average.
C K1=ITP-1 ! (= 16-1 = 15)
C

```

```

DO 7 I=1,3          ! atmospheric model loop (3 models)
DO 7 K=1,K1        ! layer loop - from level 15 at top (K=1)
*****          ! to level 1 at 1 km (K=15)
C
C
WBAR=(WH(I,K)+WH(I,K+1))/2.0
W03=(W0(I,K)+W0(I,K+1))/2.0
DELTAZ=Z(K)-Z(K+1)
C
W1D=WH(I,K)/WH(I,K+1)
IF (ABS(W1D-1.00).LT.0.1) GO TO 3
WH1=-DELTAZ/ALOG(W1D)
W(I,1,K)=WH1*HCONV*(WH(I,K+1)-WH(I,K))      ! water vapor conc.
GO TO 4
C
3 W(I,1,K)=WBAR*HCONV*DELTAZ                  ! " " "
C
4 W3D=W0(I,K)/W0(I,K+1)
IF (ABS(W3D-1.00).LT.0.1) GO TO 5
WH3=-DELTAZ/ALOG(W3D)
W(I,3,K)=WH3*O3CON*(W0(I,K+1)-W0(I,K))      ! ozone conc.
GO TO 6
C
5 W(I,3,K)=W03*O3CON*DELTAZ                  ! " "
C
C      NOTE: THE MOLECULAR DENSITY IN A GIVEN LAYER OF CO2,
C      A UNIFORMLY MIXED GAS, IS DIRECTLY RELATED TO
C      THE PRESSURE INCREMENT BETWEEN THE LAYER BOUNDARIES.
C      THE CONSTANT IN THE FOLLOWING EXPRESSION IS THE
C      UNIFORMLY MIXED GAS CONSTANT FOR CO2.
C
6 W(I,2,K)=((P(I,K+1)-P(I,K))/1013.0)*7.102E+21 ! CO2 conc.
C
7 CONTINUE
*****
C
C      READ AEROSOL MODELS FROM DATA SET 2
C      AND STORE THEM IN COMMON
C      -----
C
C      Aerosol models:  rural          - 1
C                      tropospheric - 2
C                      maritime     - 3
C                      urban         - 4
C                      background
C                      stratospheric - 5
C                      aged volcanic - 6
C
DO 10 I=1,6
READ (2,25) (ARSMOD (I,M),M=1,5)          ! read model names
READ (2,26) (FA(I,J),ASC(I,J),AAB(I,J),J=1,61) ! read data
10 CONTINUE
C
25 FORMAT(5A4)
26 FORMAT (F8.3,2F7.5)
C
DO 11 I=1,6          ! convert from microns
DO 12 J=1,61        ! to wavenumbers
FA(I,J) = 1.0E+04/FA(I,J)
12 CONTINUE
11 CONTINUE

```

```

C      CLOSE ATMOSPHERE MODEL FILE
C      -----
C 50   CLOSE (UNIT=2)
C
C      RETURN
C      END

SUBROUTINE RDTAPE
C
C      This routine selects, from RUNFIL, those lines in the range
C
C          (v-bound, v+bound)
C
C      and stores them in the (unformatted) file VM:WINDOW.DAT.
C
C      The number of retained lines, LINECT, is sent to subroutine ATMOS.
C*****
C
C          SPECIFICATION STATEMENTS
C          *****
C
C      IMPLICIT   INTEGER*2 (I-N)
C      INTEGER    SPECES, SWITCH
C      REAL*4     LOWNU
C      DIMENSION  KEEPSW(9)
C
C      COMMON     Z(16),
C      *          P(3,16),T(3,16),W(3,3,16),CAY(3,16),WG(3,3),
C      *          TA2(16),FA(6,61),ASC(6,61),AAB(6,61),
C      *          HAS(16)
C
C      COMMON/BLK1/  DNU(16),CHI(16),HZ1(16),HZ2(16),ITP,JT,KSAM
C      COMMON/RSETUP/ BOUND,LOWNU,HIGHNU,DELNU,V,
C      *              IATMOD,IARMDG,IARMDT,IARMDS,
C      *              NUMPTS,NLINES,KEEPSW,IBLAYR,ITLAYR,
C      *              IRSIZE, HZINT(16),IAERMD
C      DATA       SWITCH/0/
C*****
C
C      OPEN THE VIRTUAL MEMORY FILE FOR THE WINDOW OF LINES
C      (unformatted, sequential access)
C
C      IF(IRSIZE .LE. 164)
C      *   OPEN (UNIT=10,NAME='VM:WINDOW.DAT',TYPE='NEW',DISP='DELETE',
C      *       FORM='UNFORMATTED',RECORDSIZE=6,INITIALSIZE=IRSIZE)
C
C      IF(IRSIZE .GT. 164)
C      *   OPEN (UNIT=10,NAME='SD3:WINDOW.DAT',TYPE='NEW',DISP='DELETE',
C      *       FORM='UNFORMATTED',RECORDSIZE=6,INITIALSIZE=650)
C
C      READ, FROM RUNFIL, ALL LINES IN THE RANGE (V-BOUND, V+BOUND)
C      AND STORE THEM IN VM:WINDOW.DAT
C      //////////////////////////////////////
C
C      LINECT = 0           ! initialize line count
C      IF (NLINES .EQ. 0) GO TO 20 ! no lines, so no window file
C      REWIND 3           ! position pointer to beginning of 3

```

```

1 READ(3,END=20) SPECES, ISOTOP, WAVNUM, STRNTH, HWIDTH, ENERGY
*****
C
C
IF (WAVNUM .LT. V-BOUND) GO TO 1
IF (WAVNUM .GT. V+BOUND) GO TO 20
C
LINECT = LINECT + 1
C
WRITE(10) SPECES, ISOTOP, WAVNUM, STRNTH, HWIDTH, ENERGY
C
GO TO 1
*****
C
C*****
C
SEND THIS 'WINDOW' OF LINES TO SUBROUTINE ATMOS
////////////////////////////////////
C
20 CALL ATMOS (LINECT)
C
RETURN
END

SUBROUTINE ATMOS (LINECT)
C
C
C SPECIFICATION STATEMENTS
C *****
C
IMPLICIT INTEGER*2 (I-N)
REAL*4 LOWNU
DIMENSION CS2(3),KEEPSW(9)
C
C NOTE: ST = temp. corrected line intensity
C CS2 = partition func. temp. correction
C
COMMON Z(16),
* P(3,16),T(3,16),W(3,3,16),CAY(3,16),WG(3,3),
* TA2(16),FA(6,61),ASC(6,61),AAB(6,61),
* HAS(16)
C
COMMON/BLK1/ DNU(16),CHI(16),HZ1(16),HZ2(16),ITP,JT,KSAM
COMMON/RSETUP/ BOUND,LOWNU,HIGHNU,DELNU,V,
* IATMOD,IARMDG,IARMDT,IARMDS,
* NUMPTS,NLINES,KEEPSW,IBLAYR,ITLAYR,
* IRSIZE, HZINT(16),IAERMD
C
DATA CONST,PI/1.380258E-24,3.141593/
C
C NOTE: CONST= (R*1.0E-03)/(A*1.0E+05)
C = (8.3144E+07*1.0E-03)/(6.0238E+23*1.0E+05)
C (R = GAS CONST; A = AVOGADRO'S NUMBER)
C*****

```

```

C      DO 18 IS THE MAJOR COMPUTATIONAL LOOP ON ATMOSPHERIC LAYERS
C      WITHIN WHICH MONOCHROMATIC MOLECULAR ABSORPTION COEFFICIENTS
C      ARE COMPUTED.
C
C      IK1=IATMOD                ! (use only this atm. model)
C
C      DO 18 K=ITLAYR,IBLAYR    ! Loop over atm. layers
C      *****                  *****
C      IF (K.EQ.ITP) GO TO 2
C
C      PBAR=(P(IK1,K)+P(IK1,K+1))/2.0      ! av. layer pressure
C      TBAR=(T(IK1,K)+T(IK1,K+1))/2.0      ! av. layer temp.
C      PH2O=CONST*TBAR*W(IK1,1,K)/ABS(Z(K)-Z(K+1)) ! water vapor pressure
C      GO TO 3
C
C      2  PBAR=P(IK1,ITP)                ! av. pressure AT GROUND
C      TBAR=T(IK1,ITP)                  ! av. temp. at ground
C      PH2O=CONST*TBAR*WG(IK1,1)        ! water vapor press. at ground
C
C      3  CONTINUE
C
C      DETERMINE CORRECT TEMPERATURE DEPENDENCE OF ROTATIONAL PARTITION
C      FUNCTION
C
C      CS2(2) = 296.0/TBAR
C      CS2(1) = CS2(2)**1.5
C      CS2(3) = CS2(1)
C
C      DO 17 LOOP CYCLES THRU ALL ABSORPTION LINES
C      AND ADDS THEIR CONTRIBUTIONS TO THE ABSORPTION
C      COEFFICIENT AT THE WAVENUMBER V.
C
C      CAY(IK1,K) = 0.0                ! initialize for loop over LINES
C      IF (LINECT .EQ. 0) GO TO 18
C      IF (IAERMD .EQ. 1) GO TO 18
C      REWIND 10                        ! position pointer to beginning of 10
C
C      DO 17 I=1,LINECT                ! Loop over absorption lines in window
C      *****                          *****
C
C      READ(10) MOL, ISOTOP, GNU, S, ALFA0, EPP
C
C      M = MOL                          ! Molecular species:
C                                     1 - H2O
C                                     2 - CO2
C                                     3 - O3
C
C      ST = S*CS2(M)*EXP(-EPP*          ! temp. corrected line int.
*      (296.-TBAR)/(296.*TBAR*.6951)) ! (Boltzmann temp.
C                                     correction factor)
C
C      INSERT THE ISOTOPIC RELATIVE ABUNDANCE FACTOR
C
C      IF (M .EQ. 1) GO TO 100
C      IF (M .EQ. 2) GO TO 200
C      IF (M .EQ. 3) GO TO 300
C

```

```

C      Water Vapor Isotopes
C
100    RELBUN = 0.
        IF (ISOTOP .EQ. 161) RELBUN = 0.9976
        IF (ISOTOP .EQ. 171) RELBUN = 0.0004
        IF (ISOTOP .EQ. 181) RELBUN = 0.002
        GO TO 400

C
C      Carbon Dioxide Isotopes
C
200    RELBUN = 0.
        IF (ISOTOP .EQ. 626) RELBUN = 0.98418
        IF (ISOTOP .EQ. 636) RELBUN = 0.01103
        IF (ISOTOP .EQ. 627) RELBUN = 0.00079
        IF (ISOTOP .EQ. 628) RELBUN = 0.00394
        GO TO 400

C
C      Ozone Isotopes
C
300    RELBUN = 0.
        IF (ISOTOP .EQ. 666) RELBUN = 1.

C
C      Modify Molecular Abundance
C
400    IF (K .EQ. ITP) CONC = RELBUN*WG(IK1,M)
        IF (K .NE. ITP) CONC = RELBUN*W(IK1,M,K)

C
C      INSERT IN PBAR THE WATER VAPOR BROADENING FACTOR
C
        PEFF=PBAR
        IF (M. EQ. 1) PEFF=PBAR + 4.0*PH20

C
C      SQRT(296./TBAR) IS THE HALF-WIDTH TEMPERATURE CORRECTION FACTOR
C
        ALPHAL=ALFAD*PEFF*SQRT(296./TBAR)/1013.0
        ZDIST=ABS(V-GNU)
        X1=1.0

C
        IF (M .NE. 2)          GO TO 15

C
        X1=0.0
        JT1=JT-1

C
        DO 9 L=1,JT1
C      *****
        IF (ZDIST.GE.DNU(L).AND.ZDIST.LE.DNU(L+1)) GO TO 8
        GO TO 9

C
8      X1=((CHI(L+1)-CHI(L))/(DNU(L+1)-DNU(L)))*(ZDIST-DNU(L))
      *
        GO TO 15

C
9      CONTINUE
C      *****
C

```

```

C      SINCE PBAR .GT. 10 MB, LORENTZ SHAPE APPLIES.
C      NOTE ALSO THAT SINCE WE NEVER CONSIDER PRESSURE LESS THAN 10 MB,
C      NEITHER VOIGT SHAPE NOR DOPPLER BROADENING APPLY.
C
15      IF (K .EQ. ITP) GO TO 16          ! (ITP=16 AT ground)
C
      CAY(IK1,K) = CAY(IK1,K)
      *          + ST*ALPHAL*CONC*X1/(PI*(ZDIST**2+ALPHAL**2))
      GO TO 17
C
16      CAY(IK1,K) = CAY(IK1,K)
      *          + ST*ALPHAL*CONC*X1/(PI*(ZDIST**2+ALPHAL**2))
C
C          IK1 = atmosphere model index
C          K   = atmosphere layer index
C          M   = molecule (1=H2O, 2=CO2, 3=O3)
C
17      CONTINUE                          ! End of absorption line loop
      *****                             *****
C          NOTE: CAY now is the total abs. coeff.
C          (ALL lines) for this atm model
C          and this layer.
C
18      CONTINUE
      *****                             ! End of layer loop
      *****
C      CLOSE WINDOW FILE AND DELETE IT
C      //////////////////////////////////////
C      CLOSE (UNIT=10, DISP='DELETE')
C
C      COMPUTE CONTINUUM ABSORPTION AND AEROSOL EXTINCTION COEFFICIENTS
C      //////////////////////////////////////
C      CALL CONT (V)
C
      RETURN
      END

```

SUBROUTINE CONT

```

C      SPECIFICATION STATEMENTS
C      *****
C
      IMPLICIT INTEGER*2 (I-N)
      REAL*4   LOWNU, CON(20), FN2(91), CN2(91)
      DIMENSION KEEPSW(9)
C
      COMMON Z(16),
      * P(3,16), T(3,16), W(3,3,16), CAY(3,16), WG(3,3), TA2(16),
      * FA(6,61), ASC(6,61), AAB(6,61), HAS(16)
C
      COMMON/BLK1/  DNU(16), CHI(16), HZ1(16), HZ2(16), ITP, JT, KSAM
      COMMON/RSETUP/ BOUND, LOWNU, HIGHNU, DELNU, V,
      * IATMOD, IARMDG, IARMDT, IARMDS,
      * NUMPTS, NLINES, KEEPSW, IBLAYR, ITLAYR,
      * IRSIZE, HZINT(16), IAERMD
C

```

```

DATA CON/0.230, 0.187, 0.147, 0.117, 0.097, 0.087,
* 0.100, 0.120, 0.147, 1.174, 0.200, 0.240,
* 0.280, 0.330, 6*0.000/

```

C

```

DATA FN2/2000., 2050., 2075., 2100., 2125., 2150., 2155., 2160.,
* 2165., 2170., 2175., 2180., 2185., 2190., 2195., 2200.,
* 2205., 2210., 2215., 2220., 2225., 2230., 2235., 2240.,
* 2245., 2250., 2255., 2260., 2265., 2270., 2275., 2280.,
* 2285., 2290., 2295., 2300., 2305., 2310., 2315., 2320.,
* 2325., 2330., 2335., 2340., 2345., 2350., 2355., 2360.,
* 2365., 2370., 2375., 2380., 2385., 2390., 2395., 2400.,
* 2405., 2410., 2415., 2420., 2425., 2430., 2435., 2440.,
* 2445., 2450., 2455., 2460., 2465., 2470., 2475., 2480.,
* 2485., 2490., 2495., 2500., 2505., 2510., 2515., 2520.,
* 2525., 2530., 2535., 2540., 2545., 2550., 2575., 2600.,
* 2625., 2650., 2800./

```

C

```

DATA CN2/1.00E-21, 1.20E-07, 1.80E-07, 6.30E-07, 2.00E-06,
* 9.00E-06, 1.13E-05, 1.36E-05, 1.66E-05, 1.96E-05,
* 2.16E-05, 2.36E-05, 2.63E-05, 2.90E-05, 3.15E-05,
* 3.40E-05, 3.66E-05, 3.92E-05, 4.26E-05, 4.60E-05,
* 4.95E-05, 5.30E-05, 5.65E-05, 6.00E-05, 6.30E-05,
* 6.60E-05, 6.89E-05, 7.18E-05, 7.39E-05, 7.60E-05,
* 7.84E-05, 8.08E-05, 8.39E-05, 8.70E-05, 9.13E-05,
* 9.56E-05, 1.08E-04, 1.20E-04, 1.36E-04, 1.52E-04,
* 1.60E-04, 1.69E-04, 1.60E-04, 1.51E-04, 1.37E-04,
* 1.23E-04, 1.19E-04, 1.16E-04, 1.14E-04, 1.12E-04,
* 1.12E-04, 1.11E-04, 1.11E-04, 1.12E-04, 1.14E-04,
* 1.14E-04, 1.12E-04, 1.10E-04, 1.07E-04, 1.02E-04,
* 9.90E-05, 9.50E-05, 9.00E-05, 8.65E-05, 8.20E-05,
* 7.65E-05, 7.05E-05, 6.50E-05, 6.10E-05, 5.50E-05,
* 4.95E-05, 4.50E-05, 4.00E-05, 3.75E-05, 3.50E-05,
* 3.10E-05, 2.65E-05, 2.50E-05, 2.20E-05, 1.95E-05,
* 1.75E-05, 1.60E-05, 1.40E-05, 1.20E-05, 1.05E-05,
* 9.50E-06, 6.00E-06, 3.50E-06, 2.00E-06, 1.50E-06,
* 1.00E-20/

```

C*****

C

```

INITIALIZE AEROSOL ABSORPTION AND SCATTERING COEFFICIENTS
////////////////////////////////////

```

C

```

DO2 K=1,ITP ! loop over 16 atm. layers

```

C

```

*****
TA2(K) = 0.0
HAS(K) = 0.0

```

2

```

CONTINUE
*****

```

C

C

```

COMPUTE MOLECULAR CONTINUUM ABSORPTION
AND RALEIGH (MOLECULAR) SCATTERING,
AND ADD TO MOLECULAR LINE ABSORPTION

```

C

```

////////////////////////////////////

```

C

```

J = IATMOD ! (use only this atm. model)

```

C

C

```

DO 17 K=ITLAYR,IBLAYR ! Loop over atm. layers
*****

```

C

C

```

TBAR = T(J,ITP) ! Temp. at ground
PBAR = P(J,ITP) ! Press. at ground
IF (K .EQ. ITP) GO TO 100

```



```

PBAR = (P(J,K) + P(J,K+1))/2.0 ! av. layer pressure
TBAR = (T(J,K) + T(J,K+1))/2.0 ! av. layer temperature

```

```

C
C
C
C
C
RAYLEIGH SCATTERING
-----

```

```

100 IF (KEEPSW(4) .EQ. 0) GO TO 200
    IF (V .GE. 2740.) GO TO 110 ! waveno. range for this effect
    GO TO 200

```

```

110 IF (K .EQ. ITP) GO TO 120
    IF (K .NE. ITP) GO TO 130

```

```

120 EVM = (273./1013.) * P(J,ITP) / T(J,ITP) ! ground layer
    GO TO 140

```

```

130 HM = 1./ALOG((P(J,K+1)*T(J,K))/(P(J,K)*T(J,K+1))) ! other layers
    EVM = (273./1013.) * HM * ((P(J,K+1)/T(J,K+1))-(P(J,K)/T(J,K)))

```

```

140 TM = 9.807E-20 * EVM * V**4.0117
    CAY(J,K) = CAY(J,K) + TM

```

```

C
C
C
C
C
3.5-4.2 MICRON H2O CONTINUUM
-----

```

```

200 IF (KEEPSW(5) .EQ. 0) GO TO 300
    IF (V .GE. 2350. .AND. V .LE. 2800.) GO TO 210 ! range of effect
    GO TO 300

```

```

210 XI = (V-2350.)/50. + 1.
    MH = XI + 1.001
    XH = XI - FLOAT(MH)
    TX5 = CON(MH)
    TX5 = TX5 + XH*(CON(MH)-CON(MH-1))
    CCONT = TX5/3.34E22
    TDEP = EXP (4.56*(296./TBAR-1.))
    CNCS = 0.12 * TDEP

```

```

    IF (K .EQ. ITP) GO TO 220
    IF (K .NE. ITP) GO TO 230

```

```

220 H2OLAY = WG(J,1) ! ground layer
    PH20 = 1.38E-24 * H2OLAY * T(J,ITP)
    GO TO 240

```

```

230 H2OLAY = W(J,1,K) ! other layers
    PH20 = 4.712E-23 * H2OLAY / ALOG(P(J,K+1)/P(J,K))

```

```

240 ABSOR = (CCONT * (PH20*TDEP + CNCS*(PBAR-PH20))/1013.) * H2OLAY
    CAY(J,K) = CAY(J,K) + ABSOR

```

```

C      8-14 MICRON H2O CONTINUUM
C      -----
C
C 300  IF (KEEPSW(6) .EQ. 0) GO TO 400
      IF (V .GE. 700. .AND. V .LE. 1250.) GO TO 310      ! range of effect
      GO TO 400                                          ! (8-14 microns)
C
C 310  TDEP = EXP (6.08*(296.0/TBAR-1.0))
      CNCS = 0.002
      CCONT = (4.18 + 5578.*EXP(-7.87E-3*V))/3.34E+22
C
      IF (K .EQ. ITP) GO TO 320
      IF (K .NE. ITP) GO TO 330
C
C 320  H2OLAY = WG(J,1)                                  ! ground layer
      PH2O = 1.38E-24 * H2OLAY * T(J,ITP)
      GO TO 340
C
C 330  H2OLAY = W(J,1,K)                                ! other layers
      PH2O = 4.712E-23 * H2OLAY / ALOG(P(J,K+1)/P(J,K))
C
C 340  ABSOR = (CCONT * (PH2O*TDEP + CNCS*(PBAR-PH2O))/1013.) * H2OLAY
      CAY(J,K) = CAY(J,K) + ABSOR
C
C
C      NITROGEN CONTINUUM
C      -----
C
C 400  IF (KEEPSW(7) .EQ. 0) GO TO 17
      IF (V .GE. 2000. .AND. V .LE. 2800.) GO TO 410      ! range of effect
      GO TO 17
C
C 410  DO 413 I=1,90
      *****
      IF (V .GE. FN2(I) .AND. V .LE. FN2(I+1)) GO TO 412
      GO TO 413
C
C 412  DELN2 = (CN2(I+1)-CN2(I))/(FN2(I+1)-FN2(I))
      CCONT = DELN2*(V-FN2(I)) + CN2(I)
      GO TO 414
C
C 413  CONTINUE
      *****
C
C 414  IF (K .EQ. ITP) GO TO 415
C
      DELP = P(J,K+1) - P(J,K)
      TN2 = 0.781 * CCONT * (PBAR/1013.) * 29.24 * TBAR * (DELP/1013.)
      GO TO 416
C
C 415  TN2 = 0.781 * CCONT * (PBAR/1013.)**2 * 1.000*296./T(J,ITP)
C
C 416  CAY(J,K) = CAY(J,K) + TN2
C
C 17   CONTINUE                                          ! end of loop over atm. layers
      *****
C
C
C
C

```

```

C      COMPUTE AEROSOL ABSORPTION AND SCATTERING COEFFICIENTS
C      ///////////////////////////////////////////////////
C
C      IF (KEEPSW(8) .EQ. 0 .AND. KEEPSW(9) .EQ. 0) RETURN
C
C      TEST WHETHER V LIES WITHIN RANGE OF AEROSOL MODEL
C      -----
C      IF (V .LT. FA(1,1) .OR.
*      V .GT. FA(1,61)) RETURN ! v out of range of table
C
C      DO 19 J=1,60
C      IF (V .GE. FA(1,J) .AND. V .LE. FA(1,J+1)) GO TO 20
19     CONTINUE
C
C      RETURN ! v above range of table
C
C      20 FAC=(V-FA(1,J))/(FA(1,J+1)-FA(1,J)) ! interpolating factor
C      (for v not in table)
C
C      IF IT DOES, COMPUTE THE COEFFICIENTS
C      -----
C
C      DO 32 K=ITLAYR,IBLAYR ! Loop over atm. layers
C      *****
C
C      Compute Scaling Factors for Clear and Hazy Aerosol Models
C      .....
C
C      HZ1 = vertical scaling factor for clear aerosol model
C      HZ2 = " " " " hazy " "
C
C      HZ1 and HZ2 are given in the block data routine,
C      for each of the 16 atmospheric layers;
C      EVH1 is just a local variable for this loop.
C
C      IF (K .EQ. ITP) GO TO 24
C      IF (HZINT(K) .EQ. HZINT(K+1)) GO TO 21
C
C      HA1=1.0/ALOG(HZINT(K+1)/HZINT(K))
C      EVH1=HA1*(HZINT(K+1)-HZINT(K)) ! effective value
C      GO TO 25
C
C      21 EVH1=HZINT(K)
C      GO TO 25
C
C      24 EVH1=HZINT(ITP)
C
C      Aerosol model to be used depends on the height Z
C      .....
C      L = aerosol model index
C
C      25 IF (Z(K) .LT. 9.0 .AND. Z(K) .GE. 2.0) GO TO 28
C      IF (Z(K) .LE. 2.0) GO TO 29
C
C      Above 9 km: Use background stratospheric (clear)
C      or aged volcanic (hazy)
C
C      IF (IARMDS .EQ. 5) L=5
C      IF (IARMDS .EQ. 6) L=6
C      GO TO 30

```

```

C      Between 2 and 9 km, use tropospheric model
C
C 28   L = 2
      GO TO 30
C
C      Below 2 km, use rural, tropospheric, maritime or urban model
C
C 29   L = IARMDG
C
C      Compute aerosol extinction coefficients for this layer
C      .....
C
C          AK = aerosol absorption coeff. at v
C          AS = aerosol scattering coeff. at v
C
C          (These are local variables, for this loop.)
C
C 30   AK = AAB(L,J) + FAC*(AAB(L,J+1)-AAB(L,J))
      AS = ASC(L,J) + FAC*(ASC(L,J+1)-ASC(L,J))
C
C      Multiply these coefficients by the proper vertical scaling
C      factors and store them in the proper arrays in COMMON
C      .....
C
C          TA2(atm. layer) - aer. abs. coeff.
C          HAS( " " ) - aer. scatt. coeff.
C
C 31   TA2(K) = EVH1*AK
      HAS(K) = EVH1*AS
C
C      IF (KEEPSW(8) .EQ. 0) TA2(K) = 0.
C      IF (KEEPSW(9) .EQ. 0) HAS(K) = 0.
C
C 32   CONTINUE
      *****
C
      RETURN
      END

BLOCK DATA
C
C IMPLICIT INTEGER*2 (I-N)
C
C COMMON/BLK1/ DNU(16),CHI(16),HZ1(16),HZ2(16),ITP,JT,KSAM
C
C DATA DNU(1),DNU(2),DNU(3),DNU(4),DNU(5),DNU(6),DNU(7),DNU(8),
.      DNU(9),DNU(10),DNU(11),DNU(12),DNU(13),DNU(14),DNU(15),
.      DNU(16)/0.0,.5,.6,.7,.8,.9,1.0,1.2,1.5,2.0,2.5,3.0,5.0,8.0,
.      10.0,15.0/
C
C DATA CHI(1),CHI(2),CHI(3),CHI(4),CHI(5),CHI(6),CHI(7),CHI(8),
.      CHI(9),CHI(10),CHI(11),CHI(12),CHI(13),CHI(14),CHI(15),
.      CHI(16)/1.00,1.00,.96,.89,.82,.77,.70,.60,.50,.41,.34,.31,
.      .29,.23,.19,0.00/

```

```

DATA  HZ1(1),HZ1(2),HZ1(3),HZ1(4),HZ1(5),HZ1(6),HZ1(7),HZ1(8),
      HZ1(9),HZ1(10),HZ1(11),HZ1(12),HZ1(13),HZ1(14),HZ1(15),HZ1(16)
      /6.43E-04,6.45E-04,6.22E-04,6.63E-04,7.14E-04,
      7.87E-04,9.80E-04,1.41E-03,2.30E-03,3.54E-03,
      4.85E-03,6.43E-03,8.19E-03,9.70E-03,2.85E-02,6.95E-02/

```

```

DATA  HZ2(1),HZ2(2),HZ2(3),HZ2(4),HZ2(5),HZ2(6),HZ2(7),HZ2(8),HZ2(9),
      HZ2(10),HZ2(11),HZ2(12),HZ2(13),HZ2(14),HZ2(15),HZ2(16)
      /2.92E-03,2.89E-03,2.80E-03,2.45E-03,2.11E-03,
      1.85E-03,1.81E-03,3.36E-03,6.22E-03,7.71E-03,
      9.30E-03,1.85E-02,3.46E-02,6.21E-02,7.57E-01,7.57E-01/

```

```

DATA  ITP,JT,KSAM /16,16,3/

```

```

END

```

```

SUBROUTINE LAYNUM (Z, LAYNO, DELHT)

```

```

C This subroutine accepts the altitude Z (above sea-level) in km;
C it returns the atmospheric layer number (layno) and the height
C (in km) above the bottom of the layer (delht).

```

```

C The layers are numbered as follows:

```

HEIGHT Z *****	LAYER NUMBER *****
14 < Z <= 15 km	1
13 < Z <= 14	2
.	.
.	.
.	.
0 < Z <= 1	15
Z = 0	16

```

C *****
C IF (Z .GT. 15.) Z = 15.

```

```

C IZ = INT(Z) ! height of bottom of layer
C DELHT = Z - FLOAT(IZ) ! height above bottom of layer

```

```

C LEVELZ = IZ + 1 ! layer number, numbering from ground up
C IF (DELHT .EQ. 0.) LEVELZ = IZ

```

```

C LAYNO = 16 - LEVELZ ! layer number, numbering from top down

```

```

C RETURN
C END

```

FUNCTION AERSOL(LAYNUM)

C
C This subroutine accepts the atmosphere layer number and returns
C the total aerosol extinction coefficient for that layer; the
C extinction coefficient is the sum of the aerosol absorption and
C scattering coefficients. The layers are defined as follows:

Height (above sea-level)	Layer Number	Region
*****	*****	*****
14 < height <= 15 km	1	:
13 < height <= 14	2	:
.	.	: Stratosphere
.	.	:
.	.	:
9 < height <= 10	6	:
8 < height <= 9	7	:
.	.	:
.	.	: Troposphere
.	.	:
2 < height <= 3	13	:
1 < height <= 2	14	:
0 < height <= 1	15	: Boundary layer
height = 0	16	:

C The aerosol models to be used for the various regions are specified
C by the integer variables iarmdg, iarmdt and iarmds; these variables
C are defined interactively by the user and are stored in the common
C block RSETUP.

C*****

```

C      REAL*4    lownu
C      DIMENSION keepsw(9)
C
C      COMMON    z(16),p(3,16),t(3,16),w(3,3,16),cay(3,16),wg(3,3),
*              ta2(16),fa(6,61),asc(6,61),aab(6,61),
C              has(16)
C
C      COMMON/RSETUP/ bound,lownu,highnu,delnu,v,
*                   iatmod,iarmdg,iarmdt,iarmds,
*                   numpts,nlines,keeps,iblayr,itlayr,
*                   irsize, HZINT(16),iaermd
C
C      IF (V .LT. FA(1,1) .OR.
*         V .GT. FA(1,61)) go to 400 ! out of aerosol model range
C      aerosol = ta2(laynum) + has(laynum)
C      return
C
C 400 aerosol = 0.
C      return
C
C      end
    
```

AUXILIARY ROUTINES AND COMMAND FILES

PROGRAM LNEDIT

C
C This program reads lines one-by-one from an 'original' line file
C (AFGL format), retains only water, carbon dioxide and ozone lines (all
C isotopes), strips away the quantum identification numbers and writes
C the processed lines out to a 'master' line file for use by the
C MICTRA routine.

C
C The disk containing this program may be run from any drive;
C the drives and filenames for the input and output files will be
C typed in by the user upon prompting by the program, which will halt
C to allow mounting of the input and output disks.

C *****

SPECIFICATION STATEMENTS *****

C
C BYTE qntmid, outfil, infil, reply, go
C DIMENSION qntmid(35), outfil(14), infil(14)

C
C *****

C SPECIFY WAVENUMBER RANGE TO BE INCLUDED IN OUTPUT FILE
C ////////////////////////////////////

C
C type 1
C 1 format(//t2,'Specify wavenumber range to be included',
C * t42,'in output file:/'
C * t10,'type minimum wavenumber: ', \$)
C accept 2, wnmin
C 2 format(f10.4)
C type 3
C 3 format(t10,'type maximum wavenumber: ', \$)
C accept 2, wnmax

C
C OPEN THE OUTPUT LINE FILE (for MICTRA; UNformatted, seq. access)
C ////////////////////////////////////

C
C type 10
C 10 format(//t2,'Type name of output file in DL0:/'
C * t2,'(i.e., the master file, unformatted,/'
C * t2,' to be used by the MICTRA routine): ', \$)
C accept 11, outfil
C 11 format(14A1)

C
C type 12
C 12 format(//t2,'Mount disk containing output file,/'
C * t2,'then type 'G' for 'GO': ', \$)
C accept 13, go
C 13 format(A1)

C
C OPEN (unit=8, name=outfil, type='NEW',
C * disp='KEEP', form='UNFORMATTED', initialsize=974,
C * recordsize=6)

C

```

C      OPEN THE INPUT LINE FILE (formatted, sequential access)
C      ////////////////////////////////////////////////////
C
C      This file should be one of the AFGL files;
C      specify the file name:
C
100    type 120
C ***
120    format(//t2,'Type name of input file'/
*      t2,'(one of the AFGL files): ',,$)
      accept 11, infil
C
      type 126
126    format(//t2,'Mount disk containing input file, '/
*      t2,'then type 'G' for 'GO': ',,$)
      accept 121, go
121    format(A1)
C
      OPEN (unit=4, name=infil, type='OLD',
*        disp='KEEP', form='FORMATTED')
C
C      OPEN TEMPORARY VM FILE
C      ////////////////////////////////////////////////////
C
199    OPEN (unit=3, name='VM:TEMP.DAT', type='NEW',
*        disp='DELETE', form='UNFORMATTED',
*        initialsize=300, recordsize=6)
C
C      LOOP - READ LINES ONE AT A TIME, PROCESS AND WRITE TO OUTPUT FILE
C      (via temporary VM file, in groups of 6000 lines)
C      ////////////////////////////////////////////////////
C
C      Read a line
C      -----
C
C      linect = 0
C
200    READ(4,230,end=300) wavnum, strnth, hwidth, energy,
*      qntmid, idate, ! strip these
*      isotop, molcul
C
230    FORMAT(f10.4, e10.3, f5.4, f10.3, 35a1, i3, i4, i3)
C
C      Reject this line if it is below the minimum wavenumber
C      or if it is other than H2O, CO2 or O3
C      -----
C
C      IF (wavnum .LT. wnmin) go to 200
C      IF (molcul .GT. 3) go to 200
C      linect = linect + 1
C
C      Write stripped-down line into temporary VM file (unformatted)
C      -----
C      WRITE(3) wavnum, strnth, hwidth, energy, isotop, molcul
C
C      Branch to appropriate code for given 'end' condition
C      -----
C
C      IF (wavnum .GT. wnmax) go to 400
C      IF (linect .GT. 6000) go to 500
C      go to 200 ! read another line

```



```

C      End of input file reached
C      -----
C
300      wnlast = wavnum
        IF (linect .EQ. 0) go to 320
C
        rewind 3
        DO 310 I=1,linect
        READ (3) wavnum, strnth, hwidth, energy, isotop, molcul
        WRITE(8) wavnum, strnth, hwidth, energy, isotop, molcul
310      continue
C
320      CLOSE (unit=3, disp='DELETE')
        CLOSE (unit=4, disp='KEEP')
C
        type 330, wnlast,wnmax
330      format(//t2,'Last wavenumber in this input file is:      ',F10.4/
*          t2,'Max. wavenumber desired for output file is: ',F10.4/
*          t2,'Is the output file complete? (Y or N): ',,$)
        accept 331, reply
331      format(A1)
        IF (reply .EQ. 'N') go to 100
        CLOSE (unit=8, disp='KEEP')
        STOP
C
C      Maximum wavenumber reached
C      -----
C
400      REWIND 3
        DO 410 I=1,linect
        READ (3) wavnum, strnth, hwidth, energy, isotop, molcul
        WRITE(8) wavnum, strnth, hwidth, energy, isotop, molcul
410      continue
C
        CLOSE (unit=3, disp='DELETE')
        CLOSE (unit=4, disp='KEEP')
        CLOSE (unit=8, disp='KEEP')
        STOP
C
C      Temporary (VM) file is filled
C      -----
C
500      REWIND 3
        DO 510 I=1,linect
        READ (3) wavnum, strnth, hwidth, energy, isotop, molcul
        WRITE(8) wavnum, strnth, hwidth, energy, isotop, molcul
510      continue
C
        CLOSE (unit=3, disp='DELETE')
        GO TO 199
C
        END

```



```

* TWENTY/' ',' ','2','0',//,THIRTY/' ',' ','3','0',//,
* FORTY/' ',' ','4','0',//,FIFTY/' ',' ','5','0',//,
* SIXTY/' ',' ','6','0',//,SEVNTY/' ',' ','7','0',//,
* EIGHTY/' ',' ','8','0',//,NINETY/' ',' ','9','0',//,
* HUNDRD/' ',' ','1','0','0',//,USER/'U'//,LCPLT/'L'//,
* CONPLT/'C'//,GODARD/'G'//,MTHOPK/'H'//,CLEAR/'C'//,
* HAZY/'H'//,HPARAM/1,6,8*0//,
* LFTNUM/9*' '//, RGTNUM/9*' '//,
* BEGNUM/12*' '//,ENDNUM/12*' '//,
* UNTDES/'L','a','r','g','e',' ','d','i','v','i','s','i','
* 'o','n',' ','='','12*' '//,11*' '//,
* ANGST/'a','n','g','s','t','r','o','m','s',' ',' '//,
* MICRON/'m','i','c','r','o','m','e','t','e','r','s'//,
* HERTZ/'H','E','R','T','Z',6*' '//,
* WAVENM/'W','A','V','E','N','U','M','B','E','R','S'//,
* A/'A'//, M/'M'//, R/'R'//, F/'F'//, E/'E'//

```

C

DATA C/2.997925E10/

C

C

SET UP PLOT SYSTEM

C

CALL BEGPLT

C

INSTRUCT USER HOW TO USE THIS ROUTINE

C

CALL TYP100

TYPE 10

10 FORMAT(//T5,'The following options are provided: '//

* t15,'Type P to plot a spectrum; '//

* t15,'Type S to stop this program.'////

* t5,'FIRST type G,RETURN to remove this message.'/

* t5,'*****'//)

CALL ECHO(YES)

ACCEPT 11, GO

11 FORMAT(A1)

CALL CLR100

CALL ECHO(NO)

CALL ALPMOD(1)

C

USE ITTINR TO RECEIVE USER'S INSTRUCTIONS

C

1000 ICHAR = ITTINR()

C

IF (ICCHAR .EQ. PLOT) GO TO 2000

IF (ICCHAR .EQ. STOPP) GO TO 4000

GO TO 1000

C

PLOT MODE

C

2000 CALL TYP100

CALL ECHO(YES)

TYPE 2010

2010 FORMAT(//T5,'Spectrum file? ', \$)

ACCEPT 2011, SPCFL1

2011 FORMAT(15A1)

TYPE 2020

2020 FORMAT(//T5,'Mount file; when ready, type G: ', \$)

ACCEPT 2021, GO

2021 FORMAT(A1)

```

C      OPEN THE SPECTRUM FILE; READ THE FIRST THREE RECORDS
C      *****
C
C      OPEN (UNIT=1, NAME=SPCFL1, TYPE='OLD', DISP='KEEP',
*        FORM='UNFORMATTED', RECORDSIZE=2)
C
C      READ(1)  PLTCOD, ANGMIC           ! Line-center or continuous plot,
C                                          ! in angstroms or microns
C      READ(1)  LOWERW, UPPERW         ! bounding wavelengths
C                                          ! (in angstroms or microns)
C      READ(1)  ATL, ATU               ! bounding aerosol+continuum
C                                          ! transmission
C
C      MORE USER'S INSTRUCTIONS
C      *****
C
C      TYPE 2022
2022    FORMAT(///T2,'The spectrum may be plotted on an angstrom, '/
*        t2,'micrometer, hertz or wavenumber scale, '/
*        t2,'increasing either to the right or left;'/)
C
C      IF (ANGMIC .EQ. 0.) TYPE 2023
C      IF (ANGMIC .EQ. 1.) TYPE 2024
C
C      2023    FORMAT(/T2,'this spectrum is set up to be plotted on an '/
*        t2,'angstrom scale, increasing toward the right;'/
*        t2,'do you want to change this? (Y or N): ', $)
2024    FORMAT(/T2,'this spectrum is set up to be plotted on a '/
*        t2,'micrometer scale, increasing toward the right;'/
*        t2,'do you want to change this? (Y or N): ', $)
C
C      ACCEPT 2025, REPLY
2025    FORMAT(A1)
C
C      IF (REPLY .EQ. 'Y') GO TO 2026
C      CALL CLR100
C      CALL ECHO(NO)
C      CALL ALPMOD(1)
C      IF (ANGMIC .EQ. 0.) SCALE = A
C      IF (ANGMIC .EQ. 1.) SCALE = M
C      SENSE = R
C      GO TO 2032
C
C      2026    TYPE 2027
2027    FORMAT(//T2,'Angstrom, micrometer, hertz or wavenumber scale?'/
*        t2,'(A,M,H or W): ', $)
C      ACCEPT 2025, SCALE
C
C      TYPE 2028
2028    FORMAT(/T2,'Increasing toward right or left?'/
*        t2,'(R or L): ', $)
C      ACCEPT 2025, SENSE
C
C      CALL CLR100
C      CALL ECHO(NO)
C      CALL ALPMOD(1)
C

```



```

SCRWIN(1) = 150.                                ! define screen window
SCRWIN(2) = 1000.
IF (SENSE .EQ. 'L') SCRWIN(1) = 1000.         ! scale increasing to left
IF (SENSE .EQ. 'L') SCRWIN(2) = 150.
SCRWIN(3) = 170.
SCRWIN(4) = 770.

C
CALL SETMAP (MAPTYP, BPARAM,IPARAM,RPARAM, USRWIN,SCRWIN)

C
C Draw a box around the screen window
C -----
CALL BOXWIN

C
C Draw tick marks and grid lines
C -----
BTICK = 0.
ANGSTR = LOWVAL- 1.                            ! draw long vertical ticks
2040 ANGSTR = ANGSTR + 1.                        ! and grid lines every
IF (ANGSTR .GT. UPPERW) GO TO 2045            ! wunit
IF (ANGSTR .LT. LOWERW) GO TO 2040
CALL TICK (BOTTOM,LONG,ANGSTR,BRIGHT)
IF (BTICK .EQ. 0.) BTICK = ANGSTR
ETICK = ANGSTR
IF (ANGSTR .NE. LOWERW .AND. ANGSTR .NE. UPPERW)
* CALL DRWLIN (ANGSTR,0.,ANGSTR,100.,ARCLIN,BRIGHT,LINTYP,PARAM)
GO TO 2040

C
2045 ISWETK = 1
IF (ETICK .EQ. BTICK) ISWETK = 0             ! eliminate unneeded end tick

C
2050 ANGSTR = LOWVAL - 0.1                    ! draw short vertical ticks
2051 ANGSTR = ANGSTR + 0.1                    ! every 1/10 th wunit
IF (ANGSTR .GT. UPPERW) GO TO 2060
IF (AMOD(ANGSTR,1.) .EQ. 0.) GO TO 2051
IF (ANGSTR .GE. LOWERW) CALL TICK (BOTTOM,SHORT,ANGSTR,BRIGHT)
GO TO 2051

C
2060 TRANS = -10.                            ! draw long hor. ticks and
2061 TRANS = TRANS + 10.                     ! grid lines every 10%
IF (TRANS .GT. 100.) GO TO 2070
CALL TICK (LEFT,LONG,TRANS,BRIGHT)
IF (TRANS .GT. 0. .AND.
* TRANS .LT. 100.)
* CALL DRWLIN (USRWIN(1),TRANS,USRWIN(2),
• TRANS,ARCLIN,BRIGHT,LINTYP,PARAM)
GO TO 2061

C
2070 TRANS = TRANS - 1.                      ! draw short hor. ticks
2071 TRANS = TRANS + 1.                      ! every 1%
IF (TRANS .GT. 100.) GO TO 2200
IF (AMOD(TRANS,10.) .EQ. 0.) GO TO 2071
CALL TICK (LEFT,SHORT,TRANS,BRIGHT)
GO TO 2071

C

```

```

C      Plot the spectrum
C      -----
2200  IF (PLTCOD .EQ. 1.) GO TO 2300
      IF (PLTCOD .EQ. 0.) GO TO 2400
C
C      The aerosol+continuum extinction is given precisely enough
C      (over a small wavelength band) by a straight line,
C      having the general form:
C
C       $AT = AT1 + (AT2-AT1)/(W2-W1) * (W-W1)$  ,
C
C      where AT = aerosol+continuum transmission in %
C      and W = wavelength in angstroms or microns
C
C      (W1,AT1) and (W2,AT2) are the two endpoints of the range.
C
2300  W1 = LOWERW                                ! LINE-CENTER PLOT
      W2 = UPPERW                                ! (overview)
C
      AT1 = ATL*100.
      AT2 = ATU*100.
C
      LOWERT = AT1 + (AT2-AT1)/(W2-W1) * (LOWERW-W1) ! draw aer/con line
      UPPERT = AT1 + (AT2-AT1)/(W2-W1) * (UPPERW-W1)
      CALL DRWLIN (LOWERW,LOWERT, UPPERW,UPPERT,
      *           ARCLIN,BRIGHT,NORMAL,PARAM)
C
      WIDTH = UPPERW - LOWERW                    ! draw hatched area above line
      HATCHS = WIDTH/50.
C
      W = LOWERW - WIDTH
      DO 2340 I=1,100
      W = W + HATCHS
      AT = AT1 + (AT2-AT1)/(W2-W1) * (W-W1)
      CALL DRWLIN (W,AT, W+WIDTH,AT+100.,
      *           ARCLIN,BRIGHT,LINTYP,HPARAM)
2340  CONTINUE
C
      W = UPPERW + WIDTH
      DO 2350 I=1,100
      W = W - HATCHS
      AT = AT1 + (AT2-AT1)/(W2-W1) * (W-W1)
      CALL DRWLIN (W,AT, W-WIDTH,AT+100.,
      *           ARCLIN,BRIGHT,LINTYP,HPARAM)
2350  CONTINUE
C
2360  READ (1,END=2100) WAVLEN, TRANSM           ! draw absorption lines
      IF (CNVFAC .GT. 0.) WAVLEN = CNVFAC * WAVLEN
      IF (CNVFAC .LT. 0.) WAVLEN = -CNVFAC/WAVLEN
      WAVLEN = WAVLEN/WUNIT
      IF (WAVLEN .GT. UPPERW) GO TO 2360
      IF (WAVLEN .LT. LOWERW) GO TO 2100
      AT = AT1 + (AT2-AT1)/(W2-W1) * (WAVLEN-W1)
      CALL DRWLIN (WAVLEN,AT, WAVLEN,TRANSM*100.,
      *           ARCLIN,BRIGHT,NORMAL,PARAM)
      GO TO 2360
C

```

```

2400 IPTCT = 1 ! CONTINUOUS PLOT
2410 READ (1,END=2100) WAVLEN, TRANSM
IF (CNVFAC .GT. 0.) WAVLEN = CNVFAC * WAVLEN
IF (CNVFAC .LT. 0.) WAVLEN = -CNVFAC/WAVLEN
WAVLEN = WAVLEN/WUNIT
IF (IPTCT .EQ. 1) GO TO 2420
CALL DRWLIN (WAVLNP,TRANSP*100., WAVLEN,TRANSM*100.,
* ARCLIN,BRIGHT,NORMAL,PARAM)
2420 WAVLNP = WAVLEN
TRANSP = TRANSM
IPTCT = IPTCT + 1
GO TO 2410

C
C Label the plot
C -----
C2100 CALL MOVSEQ (LHEDNG,21,2, 1,700,-90.,20, ! left heading
C * FNX,FNY,FNANGL,FNSPCE)
C PRINT 2110, FNX,FNY,FNSPCE
C2110 FORMAT(/T10,'Left heading: '//
C * t15,'X: ',t20,I4//
C * t15,'Y: ',t20,I4//
C * t15,'spacing:',t25,I3//)
2100 CALL DISPSQ (LHEDNG,21,BRIGHT,2,
* 42,742,-90.,33)

C
C IF (ANGMIC .EQ. 0.) ! bottom heading
C * CALL MOVSEQ (BHDNGA,23,2, 1,10, 0.,20,
C * FNX,FNY,FNANGL,FNSPCE)
C IF (ANGMIC .EQ. 1.)
C * CALL MOVSEQ (BHDNGM,25,2, 1,10, 0.,20,
C * FNX,FNY,FNANGL,FNSPCE)
C PRINT 2120, FNX,FNY,FNSPCE
C2120 FORMAT(/T10,'Bottom heading: '//
C * t15,'X: ',t20,I4//
C * t15,'Y: ',t20,I4//
C * t15,'spacing:',t25,I3//)
IF (SCALE .EQ. 'A') CALL DISPSQ (BHDNGA,23,BRIGHT,2,
* 239,35,0.,28)
IF (SCALE .EQ. 'M') CALL DISPSQ (BHDNGM,25,BRIGHT,2,
* 239,35,0.,28)
IF (SCALE .EQ. 'H') CALL DISPSQ (BHDNGH,18,BRIGHT,2,
* 239,35,0.,28)
IF (SCALE .EQ. 'W') CALL DISPSQ (BHDNGW,24,BRIGHT,2,
* 239,35,0.,28)

```

```

C
C Note: MOVSEQ and DISPSQ are two routines from the author's
C plot package, which he wrote to provide acceptable
C plotting capability for the DEC 11/23 microcomputer,
C with a DEC VT100 terminal enhanced with another
C manufacturer's board to give it better plotting
C capability. MOVSEQ allows the user to display any
C string of characters on the screen and move the entire
C string about at will, including changing the angle at
C which the string is displayed, by using the arrow keys
C and other keys on the terminal keyboard; when the
C desired position and orientation of the string have
C been found, MOVSEQ can be commanded to return those
C values, which may then be permanently coded into the
C plot routine. This gives the user a very easy way of
C labelling the plot.

```



```

C
C Comment out after get x-value:
C .....
C      DSPLFT = (UPPERW-LOWERW)/20.          ! left numbers
C      UX = LOWERW - DSPLFT
C      UY = 0.
C      CALL MAP(USER)
C      ILNX = INT(SX)
C      ILNY = INT(SY)
C
C      CALL MOVSEQ (ZERO,4,1, ILNX,ILNY, 0.,5,
C *          FNX,FNY,FNANGL,FNSPCE)
C
C      PRINT 2130, FNX,FNY,FNSPCE
C2130     FORMAT(/T10,'Left numbers: '//
C *          t15,'X: ',t20,I4//
C *          t15,'Y: ',t20,I4//
C *          t15,'spacing:',t25,I3//)
C .....
C
C      IILNX = printed x-value for 'left numbers'
C      IILNX = 87
C      ISPNG = printed spacing for 'left numbers'
C      ISPNG = 15
C
C      DO 2150 I=1,11
C      UX = LOWERW
C      UY = 0. + FLOAT((I-1)*10)
C      CALL MAP(USER)
C      IIX = INT(SX)
C      IILNY(I) = INT(SY)
C2150     CONTINUE
C
C      CALL DISPSQ (ZERO,4,BRIGHT,1,IILNX,IILNY(1)-6,0.,ISPNG)
C      CALL DISPSQ (TEN,4,BRIGHT,1,IILNX,IILNY(2)-6,0.,ISPNG)
C      CALL DISPSQ (TWENTY,4,BRIGHT,1,IILNX,IILNY(3)-6,0.,ISPNG)
C      CALL DISPSQ (THIRTY,4,BRIGHT,1,IILNX,IILNY(4)-6,0.,ISPNG)
C      CALL DISPSQ (FORTY,4,BRIGHT,1,IILNX,IILNY(5)-6,0.,ISPNG)
C      CALL DISPSQ (FIFTY,4,BRIGHT,1,IILNX,IILNY(6)-6,0.,ISPNG)
C      CALL DISPSQ (SIXTY,4,BRIGHT,1,IILNX,IILNY(7)-6,0.,ISPNG)
C      CALL DISPSQ (SEVENTY,4,BRIGHT,1,IILNX,IILNY(8)-6,0.,ISPNG)
C      CALL DISPSQ (EIGHTY,4,BRIGHT,1,IILNX,IILNY(9)-6,0.,ISPNG)
C      CALL DISPSQ (NINETY,4,BRIGHT,1,IILNX,IILNY(10)-6,0.,ISPNG)
C      CALL DISPSQ (HUNDRD,4,BRIGHT,1,IILNX,IILNY(11)-6,0.,ISPNG)
C
C
C      UX = BTICK          ! map bottom numbers
C      UY = 0.            ! to screen
C      CALL MAP(USER)
C      IXBTCK = INT(SX)
C      IYBTCK = INT(SY)
C
C      UX = ETICK
C      UY = 0.
C      CALL MAP(USER)
C      IXETCK = INT(SX)
C      IYETCK = INT(SY)
C

```

```

BTICK = BTICK * WUNIT          ! encode bottom numbers
ETICK = ETICK * WUNIT

C
FTBTCK = E
IF (BTICK .GE. 1.E-4 .AND. BTICK .LE. 99999.) FTBTCK = F
FTETCK = E
IF (ETICK .GE. 1.E-4 .AND. ETICK .LE. 99999.) FTETCK = F

C
IF (FTBTCK .EQ. F) ENCODE (12,2162,BEGNUM) BTICK
IF (FTBTCK .EQ. E) ENCODE (12,2161,BEGNUM) BTICK
IF (FTETCK .EQ. F) ENCODE (12,2162,ENDNUM) ETICK
IF (FTETCK .EQ. E) ENCODE (12,2161,ENDNUM) ETICK
2161 FORMAT(E12.5)
2162 FORMAT(F12.4)

C
IF (SENSE .EQ. 'R') GO TO 2500          ! adjust positions
IF (SENSE .EQ. 'L') GO TO 2600          ! of bottom no's

C
2500 IF (FTBTCK .EQ. F) IXBTCK = MAX0(IXBTCK-7*15,85)
IF (FTBTCK .EQ. E) IXBTCK = MAX0(IXBTCK-6*15,85)
IF (FTETCK .EQ. F) IXETCK = MIN0(IXETCK-7*15,1024-12*15)
IF (FTETCK .EQ. E) IXETCK = MIN0(IXETCK-6*15,1024-12*15)
GO TO 2610

C
2600 IF (FTETCK .EQ. F) IXETCK = MAX0(IXETCK-7*15,85)
IF (FTETCK .EQ. E) IXETCK = MAX0(IXETCK-6*15,85)
IF (FTBTCK .EQ. F) IXBTCK = MIN0(IXBTCK-7*15,1024-12*15)
IF (FTBTCK .EQ. E) IXBTCK = MIN0(IXBTCK-6*15,1024-12*15)

C
2610 CALL DISPSQ (BEGNUM,12,BRIGHT,1,IXBTCK,100,0.,15) ! display bottom
IF (ISWETK .EQ. 1) ! numbers
* CALL DISPSQ (ENDNUM,12,BRIGHT,1,IXETCK,100,0.,15)

C
IF (WUNIT .LT. 1.E-4 .OR. WUNIT .GT. 99999.) ! encode scale
* ENCODE (12,2161,WLUNIT) WUNIT ! unit
IF (WUNIT .GE. 1.E-4 .AND. WUNIT .LE. 99999.)
* ENCODE (12,2162,WLUNIT) WUNIT

C
DO 2165 I=1,12
UNTDES(I+17) = WLUNIT(I)
2165 CONTINUE

C
DO 2166 I=1,11
IF (SCALE .EQ. 'A') UNTDES(I+30) = ANGST(I)
IF (SCALE .EQ. 'M') UNTDES(I+30) = MICRON(I)
IF (SCALE .EQ. 'H') UNTDES(I+30) = HERTZ(I)
IF (SCALE .EQ. 'W') UNTDES(I+30) = WAVENM(I)
2166 CONTINUE

C
IF (SCALE .EQ. 'A') IXUNT = 250          ! display scale
IF (SCALE .EQ. 'M') IXUNT = 265
IF (SCALE .EQ. 'H') IXUNT = 230
IF (SCALE .EQ. 'W') IXUNT = 260
CALL DISPSQ (UNTDES,41,BRIGHT,1,IXUNT,1,0.,15)

```

```
C      CLOSE THE SPECTRUM FILE, CLEAR THE SCREEN
C      AND TAKE NEXT INSTRUCTION AT USER'S CUE
C      *****
      CLOSE (UNIT=1,DISP='KEEP')
      CALL ECHO(YES)
2180  ACCEPT 2180, GO
      FORMAT(A1)
      CALL ECHO(NO)
      CALL CLR640
      GO TO 1000

C
C
C      STOP THE PROGRAM
C      *****
4000  CALL ENDPLT
      STOP
      END
```

```

!EDMICB.COM
!
!           COMMAND FILE TO EDIT, DEVELOP AND DEBUG MICTRA
!           *****
!
! This command file directs the editing, compilation and linking of
! the MICTRA routines. Any or all of the routines may be edited;
! the edited routines are then compiled and linked with the unedited
! routines to form an updated load module. The user is prompted to
! mount any needed floppy disks, and housekeeping chores (deleting
! unneeded files, squeezing disks, etc) are done automatically.
!
!           LOAD FLOPPIES: OPERATING   DISK IN DRIVE 0
!                       DEVELOPMENT  DISK IN DRIVE 1
!           *****
!                       (Scratch disk is SD3)
!
! RUN COM:EDMICB.SV1           ! Tell user to save files
! INIT/NOQ SD3:
! RUN COM:EDMICB.SV2           ! Mount floppies in DY0,
!                               DY1
!
!-----
!           HOUSEKEEPING
!           *****
! COPY/NOQ DY1:*.LST SD3:
! DELETE/NOQ DY1:*.LST
! SQUEEZE/NOQUERY DY1:
!-----
!           EDIT AND RECOMPILE MICDRB
!           *****
! EDIT DY1:MICDRB.FOR
! COPY DY1:MICDRB.BAK SD3:MICDRB.BAK
! DELETE/NOQUERY DY1:MICDRB.BAK
! SQUEEZE/NOQUERY DY1:
! FORT/OBJ:DY1:MICDRB.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:MICDRB      DY1:MICDRB.FOR
!-----
!           EDIT AND RECOMPILE SETUP
!           *****
! EDIT DY1:SETUP.FOR
! COPY DY1:SETUP.BAK SD3:SETUP.BAK
! DELETE/NOQUERY DY1:SETUP.BAK
! SQUEEZE/NOQUERY DY1:
! FORT/OBJ:DY1:SETUP.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:SETUP      DY1:SETUP.FOR
!-----
!           EDIT AND RECOMPILE RDTAPE
!           *****
! EDIT DY1:RDTAPE.FOR
! COPY DY1:RDTAPE.BAK SD3:RDTAPE.BAK
! DELETE/NOQUERY DY1:RDTAPE.BAK
! SQUEEZE/NOQUERY DY1:
! FORT/OBJ:DY1:RDTAPE.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:RDTAPE    DY1:RDTAPE.FOR
!-----
!           EDIT AND RECOMPILE ATMOS
!           *****
! EDIT DY1:ATMOS.FOR
! COPY DY1:ATMOS.BAK SD3:ATMOS.BAK
! DELETE/NOQUERY DY1:ATMOS.BAK
! SQUEEZE/NOQUERY DY1:
! FORT/OBJ:DY1:ATMOS.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:ATMOS      DY1:ATMOS.FOR
!-----

```

```

!           EDIT AND RECOMPILE CONT
!           *****
EDIT DY1:CONT.FOR
COPY DY1:CONT.BAK SD3:CONT.BAK
DELETE/NOQUERY DY1:CONT.BAK
SQUEEZE/NOQUERY DY1:
FORT/OBJ:DY1:CONT.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:CONT      DY1:CONT.FOR
!-----
!           EDIT AND RECOMPILE BLOCK
!           *****
EDIT DY1:BLOCK.FOR
COPY DY1:BLOCK.BAK SD3:BLOCK.BAK
DELETE/NOQUERY DY1:BLOCK.BAK
SQUEEZE/NOQUERY DY1:
FORT/OBJ:DY1:BLOCK.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:BLOCK    DY1:BLOCK.FOR
!-----
!           EDIT AND RECOMPILE LAYNUM
!           *****
EDIT DY1:LAYNUM.FOR
COPY DY1:LAYNUM.BAK SD3:LAYNUM.BAK
DELETE/NOQUERY DY1:LAYNUM.BAK
SQUEEZE/NOQUERY DY1:
FORT/OBJ:DY1:LAYNUM.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:LAYNUM  DY1:LAYNUM.FOR
!-----
!           EDIT AND RECOMPILE AERSOL
!           *****
EDIT DY1:AERSOL.FOR
COPY DY1:AERSOL.BAK SD3:AERSOL.BAK
DELETE/NOQUERY DY1:AERSOL.BAK
SQUEEZE/NOQUERY DY1:
FORT/OBJ:DY1:AERSOL.OBJ/CODE:THR/EXT/U:10/WA/LIST:SD3:AERSOL  DY1:AERSOL.FOR
!-----
!           HOUSEKEEPING
!           *****
SQUEEZE/NOQUERY DY1:
!-----
!           LINK ENTIRE PROGRAM, THEN DELETE OBJECT MODULES
!           (to save space on disk)
!           *****
LINK/EXECUTE:DY1:MICTRA.SAV/PROMPT DY1:MICDRB,DY1:SETUP
DY1:RDTAPE,DY1:ATMOS,DY1:CONT,DY1:BLOCK,DY1:LAYNUM
DY1:AERSOL,SY:FORLIB//
!
DELETE/NOQUERY DY1:(MICDRB.OBJ,SETUP.OBJ,RDTAPE.OBJ)
DELETE/NOQUERY DY1:(ATMOS.OBJ,CONT.OBJ,BLOCK.OBJ)
DELETE/NOQUERY DY1:(LAYNUM.OBJ,AERSOL.OBJ)
SQUEEZE/NOQUERY DY1:
!-----
!           COPY MICTRA TO OPERATING DISK (DYO)
!           AND SHOW BOTH DIRECTORIES
!           *****
COPY DYO:MICTRA.SAV SD3:MICBBK.SAV
DELETE/NOQ DYO:MICTRA.SAV
SQUEEZE/NOQ DYO:
COPY DY1:MICTRA.SAV DYO:MICTRA.SAV
DELETE/NOQ DY1:MICTRA.SAV
COPY/NOQ SD3:*.LST DY1:
DIR/ORDER/FULL DYO:
DIR/ORDER/FULL DY1:

```

```

!RNMICB.COM
!   This command file, to run the batch version of MICTRA,
!   does nothing but run a FORTRAN program which types on the screen
!   instructions to the user on which further command files to run
!   in order to run the MICTRA batch routine in its various modes.
!
RUN COM:RNMICB.SV1

!RMICB1.COM
!   COMMAND FILE TO RUN MICTRA FROM THE RLO2
!   IN THE MODE WHICH GENERATES A BATCH FILE
!   *****
!
! Initialize the RLO2 simulated disks
! *****
!
RUN COM:RNMICB.LF      ! Make sure user has a condensed linefile on hand.
RUN COM:RNMICB.SVO    ! Tell user to save files before initializing disks.
INIT/NOQ SD0:
INIT/NOQ SD1:
INIT/NOQ SD2:
INIT/NOQ SD3:
INIT/NOQ SCR:
!
! Instruct user to load operating MICTRA floppy on DY0:
! and a floppy containing condensed line files on DY1:
! *****
RUN COM:RNMICB.SV2
!
! Copy appropriate files from floppy to simulated disk on RLO2
! *****
COPY DY0:MICTRA.SAV SD2:MICTRA.SAV
COPY DY0:ATMMOD.DAT SD2:ATMMOD.DAT
COPY DY1:*. * SD1:
!
! Instruct user how to run MICTRA
! *****
RUN COM:RNMICB.SV3
!
! Run the batch MICTRA routine from the RLO2
! *****
RUN SD2:MICTRA.SAV

!RMICB2.COM
!   COMMAND FILE TO RUN MICTRA FROM THE RLO2
!   IN THE MODE WHICH EXECUTES A BATCH FILE
!   *****
!
! Initialize the RLO2 simulated disks
! *****
!
RUN COM:RNMICB.LF      ! Make sure the user has a condensed linefile on hand.
RUN COM:RNMICB.SOD    ! Tell user to save files on disks to be initialized.
INIT/NOQ SD1:
INIT/NOQ SD2:
INIT/NOQ SD3:
INIT/NOQ SCR:
!

```

```

! Instruct user to load operating MICTRA floppy on DY0:
! and a floppy containing condensed line files on DY1:
! *****
RUN COM:RNMICB.SV2
!
! Copy appropriate files from floppy to simulated disk on RL02
! *****
COPY DY0:MICTRA.SAV SD2:MICTRA.SAV
COPY DY0:ATMMOD.DAT SD2:ATMMOD.DAT
COPY DY1:*. * SD1:
!
! See if batch file is on SD0:
! *****
RUN COM:RNMICB.SV4
DIR/ORDER/FULL/VOLUMEID SD0:
!
! If not, load a floppy with the batch file into DY0:
! *****
RUN COM:RNMICB.SV5
!
! Run the BATCH compiler
! *****
@COM: BATCH.COM

```

```

!RMICB3.COM
!
! COMMAND FILE TO RUN MICTRA FROM THE RL02
! IN THE MODE WHICH GENERATES A BATCH FILE
! AND THEN EXECUTES IT
! *****
!
! Initialize the RL02 simulated disks
! *****
RUN COM:RNMICB.LF ! Make sure the user has a condensed linefile on hand.
RUN COM:RNMICB.SV0 ! Tell user to save files on disks to be initialized.
INIT/NOQ SD0:
INIT/NOQ SD1:
INIT/NOQ SD2:
INIT/NOQ SD3:
INIT/NOQ SCR:
!
! Instruct user to load operating MICTRA floppy on DY0:
! and a floppy containing condensed line files on DY1:
! *****
RUN COM:RNMICB.SV2
!
! Copy appropriate files from floppy to simulated disk on RL02
! *****
COPY DY0:MICTRA.SAV SD2:MICTRA.SAV
COPY DY0:ATMMOD.DAT SD2:ATMMOD.DAT
COPY DY1:*. * SD1:
!
! Instruct user how to run MICTRA
! *****
RUN COM:RNMICB.SV3
!

```

```

! Run the batch MICTRA routine from the RLO2,
! to create a batch file to be executed
! *****
RUN SD2:MICTRA.SAV
!
! Instruct the user how to respond to BATCH's prompt
! for a file to be executed (the file is SDO:BATFIL.BAT)
! *****
RUN COM:RNMICB.SV6
!
! Run the BATCH compiler
! *****
@COM:BATC.COM

```

```

! BATCH.COM
! This command file runs the BATCH compiler
LOAD TT
!LOAD LP
LOAD BA
ASSIGN TT LOG ! (or LP)
ASSIGN TT LST
R BATCH

```

```

!EDLPLT.COM
! This command file edits, compiles and links the plot routine PLOTSP,
! which displays spectra computed by MICTRA
!
RUN COM:EDLPLT.SAV ! tell user to save any important files on SDO
! and to mount the MICTRA operating disk in DY0
! and the development disk in drive DY1.
INIT/NOQ SDO:
EDIT DY1:PLOTSP.FOR
FORT/OBJ:SDO:PLOTSP.OBJ/CODE:THR/EXT/WA/LIST:SDO:PLOTSP.LST DY1:PLOTSP.FOR
SQUEEZE/NOQ DY1:
LINK/EXE:SDO:PLOTSP.SAV SDO:PLOTSP.OBJ,PLT:PLTLIB,SY:FORLIB
DELETE/NOQ DY0:PLOTSP.SAV
SQUEEZE/NOQ DY0:
COPY SDO:PLOTSP.SAV DY0:PLOTSP.SAV
DIR/ORDER/FULL/VOLUMEID DY0:
DIR/ORDER/FULL/VOLUMEID DY1:

```

```

!RNLPLT.COM
! This command file runs the auxiliary plotting routine PLOTSP,
! which displays spectra computed by MICTRA.
!
RUN COM:RNLPLT.SAV ! Tell user to mount MICTRA operating disk
! in drive DY0.
RUN DY0:PLOTSP.SAV

```



```

!EDLLIN.COM
! This command file edits, compiles and links the line editing routine
! LNEDIT, which processes an Air Force line file in ASCII characters
! into the condensed binary form used by MICTRA.
!
RUN COM:EDLLIN.SAV          ! tell user to save any important files on SDO
!                          and to mount the MICTRA operating disk in DY0
!                          and the development disk in drive DY1.
INIT/NOQ SDO:
EDIT DY1:LNEDIT.FOR
FORT/OBJ:SDO:LNEDIT.OBJ/CODE:THR/EXT/WA/LIST:SDO:LNEDIT.LST DY1:LNEDIT.FOR
SQUEEZE/NOQ DY1:
LINK/EXE:SDO:LNEDIT.SAV SDO:LNEDIT.OBJ,PLT:PLTLIB,SY:FORLIB
DELETE/NOQ DY0:LNEDIT.SAV
SQUEEZE/NOQ DY0:
COPY SDO:LNEDIT.SAV DY0:LNEDIT.SAV
DIR/ORDER/FULL/VOLUMEID DY0:
DIR/ORDER/FULL/VOLUMEID DY1:

```

```

!RNLLIN.COM
! This command file runs the auxiliary line editing routine LNEDIT,
! which processes Air Force line files in ASCII characters into
! the condensed binary form used by MICTRA.
!
RUN COM:RNLLIN.SV1        ! Make sure user has an Air Force ASCII
!                          line file at hand
RUN COM:RNLLIN.SAV        ! Tell user to mount MICTRA operating disk
!                          in drive DY0.
RUN DY0:LNEDIT.SAV

```

ATMOSPHERE AND AEROSOL MODELS

(These models are on a disk file, in exactly the form shown below; the first three lines are the sea-level molecular abundances of H2O, CO2 and O3, in molecules/sq cm/km, for the three atmosphere models.)

4.69E+22 8.24E+20 7.53E+16
 1.17E+22 8.95E+20 7.53E+16
 1.97E+22 8.41E+20 6.78E+16

MIDLATITUDE SUMMER

0.0	1.013E+03	294.0	1.4E+01	6.0E-05
1.0	9.020E+02	290.0	9.3E+00	6.0E-05
2.0	8.020E+02	285.0	5.9E+00	6.0E-05
3.0	7.100E+02	279.0	3.3E+00	6.2E-05
4.0	6.280E+02	273.0	1.9E+00	6.4E-05
5.0	5.540E+02	267.0	1.0E+00	6.6E-05
6.0	4.870E+02	261.0	6.1E-01	6.9E-05
7.0	4.260E+02	255.0	3.7E-01	7.5E-05
8.0	3.720E+02	248.0	2.1E-01	7.9E-05
9.0	3.240E+02	242.0	1.2E-01	8.6E-05
10.0	2.810E+02	235.0	6.4E-02	9.0E-05
11.0	2.430E+02	229.0	2.2E-02	1.1E-04
12.0	2.090E+02	222.0	6.0E-03	1.2E-04
13.0	1.790E+02	216.0	1.8E-03	1.5E-04
14.0	1.530E+02	216.0	1.0E-03	1.8E-04
15.0	1.300E+02	216.0	7.6E-04	1.9E-04

MIDLATITUDE WINTER

0.0	1.018E+03	272.2	3.5E+00	6.0E-05
1.0	8.973E+02	268.7	2.5E+00	5.4E-05
2.0	7.897E+02	265.2	1.8E+00	4.9E-05
3.0	6.938E+02	261.7	1.2E+00	4.9E-05
4.0	6.081E+02	255.7	6.6E-01	4.9E-05
5.0	5.313E+02	249.7	3.8E-01	5.8E-05
6.0	4.627E+02	243.7	2.1E-01	6.4E-05
7.0	4.016E+02	237.7	8.5E-02	7.7E-05
8.0	3.473E+02	231.7	3.5E-02	9.0E-05
9.0	2.992E+02	225.7	1.6E-02	1.2E-04
10.0	2.568E+02	219.7	7.5E-03	1.6E-04
11.0	2.199E+02	219.2	6.9E-03	2.1E-04
12.0	1.882E+02	218.7	6.0E-03	2.6E-04
13.0	1.610E+02	218.2	1.8E-03	3.0E-04
14.0	1.378E+02	217.7	1.0E-03	3.2E-04
15.0	1.178E+02	217.2	7.6E-04	3.4E-04

U.S. STANDARD

0.0	1.013E+03	288.1	5.9E+00	5.4E-05
1.0	8.986E+02	281.6	4.2E+00	5.4E-05
2.0	7.950E+02	275.1	2.9E+00	5.4E-05
3.0	7.012E+02	268.7	1.8E+00	5.0E-05
4.0	6.166E+02	262.2	1.1E+00	4.6E-05
5.0	5.405E+02	255.7	6.4E-01	4.6E-05
6.0	4.722E+02	249.2	3.8E-01	4.5E-05
7.0	4.111E+02	242.7	2.1E-01	4.9E-05
8.0	3.565E+02	236.2	1.2E-01	5.2E-05
9.0	3.080E+02	229.7	4.6E-02	7.1E-05
10.0	2.650E+02	223.2	1.8E-02	9.0E-05
11.0	2.270E+02	216.8	8.2E-03	1.3E-04
12.0	1.940E+02	216.6	3.7E-03	1.6E-04
13.0	1.658E+02	216.6	1.8E-03	1.7E-04
14.0	1.417E+02	216.6	8.4E-04	1.9E-04
15.0	1.211E+02	216.6	7.2E-04	2.1E-04

Explanation:

Atmosphere models

Column	Quantity
-----	-----
1	height (km)
2	pressure (mb)
3	temp. (K)
4	density of water vapor (g/cubic m)
5	density of ozone (g/cubic m)

(The density of CO2 is calculated in the program.)

Aerosol models

1	wavelength (micrometers)
2	aerosol scattering coeff.
3	aerosol absorption coefficient

The aerosol coefficients are normalized to 1.00 per km at 0.55 micrometers.

RURAL

40.000	.02311	.04972
35.000	.02601	.04885
30.000	.02807	.04817
27.900	.03095	.04713
25.000	.03505	.04737
22.500	.03883	.04856
21.300	.03988	.04959
20.000	.04175	.04857
18.500	.04276	.04286
18.000	.04495	.04387
17.200	.04512	.04857
16.400	.04353	.04407
15.000	.03376	.05278
14.800	.03926	.03855
14.000	.04733	.03407
13.000	.05164	.03310
12.500	.05370	.03269
11.500	.05823	.03384
11.000	.05972	.03607
10.591	.05749	.04301
10.000	.05722	.05123
9.800	.05774	.05266
9.500	.05477	.06358
9.200	.05289	.08719
9.000	.05719	.07649
8.700	.05497	.07125
8.500	.03319	.07937
8.200	.01340	.05110
7.900	.03074	.03904
7.200	.05216	.04713
6.500	.05488	.03354
6.200	.05696	.03113
6.000	.05870	.02921
5.500	.06904	.02455
5.000	.07672	.02033
4.500	.08011	.02219
4.000	.09001	.01465
3.750	.09394	.01357
3.500	.09598	.01583
3.392	.09251	.01979
3.200	.09339	.02046
3.000	.08330	.03696
2.700	.07835	.06559
2.500	.1016	.02932
2.250	.1140	.02894
2.000	.1312	.02763
1.800	.1535	.04432
1.536	.2057	.05907
1.300	.2697	.06247
1.060	.3802	.06775
0.860	.5122	.06694
0.694	.7023	.06129
0.633	.7900	.06159
0.550	.9307	.06928
0.515	1.008	.06748
0.488	1.068	.07050
0.400	1.311	.08337
0.337	1.537	.09620
0.300	1.673	.1335
0.250	1.856	.2317
0.200	1.916	.5029

TROPOSPHERIC

40.000	.00006	.01807
35.000	.00009	.01730
30.000	.00012	.01732
27.900	.00017	.01608
25.000	.00026	.01555
22.500	.00040	.01553
21.300	.00049	.01616
20.000	.00061	.01530
18.500	.00056	.01345
18.000	.00072	.01324
17.200	.00092	.01578
16.400	.00068	.01488
15.000	.00046	.02374
14.800	.00049	.01389
14.000	.00075	.01076
13.000	.00104	.01036
12.500	.00122	.00997
11.500	.00185	.01027
11.000	.00235	.01108
10.591	.00275	.01445
10.000	.00375	.01895
9.800	.00423	.01966
9.500	.00484	.02700
9.200	.00620	.04812
9.000	.00847	.03750
8.700	.00785	.03419
8.500	.00121	.04847
8.200	.00017	.02929
7.900	.00055	.01801
7.200	.00294	.02036
6.500	.00383	.01276
6.200	.00393	.01169
6.000	.00401	.01080
5.500	.00611	.00805
5.000	.00849	.00623
4.500	.01159	.00701
4.000	.01564	.00434
3.750	.01838	.00402
3.500	.02174	.00483
3.392	.02133	.00648
3.200	.02350	.00677
3.000	.02394	.01589
2.700	.02650	.03891
2.500	.03797	.01157
2.250	.05005	.01146
2.000	.06761	.01092
1.800	.09904	.02200
1.536	.1603	.03451
1.300	.2313	.03796
1.060	.3532	.04336
0.860	.4971	.04326
0.694	.7028	.03916
0.633	.7983	.03985
0.550	.9528	.04716
0.515	1.037	.04589
0.488	1.102	.04884
0.400	1.368	.06144
0.337	1.616	.07429
0.300	1.766	.1119
0.250	1.966	.2184
0.200	2.036	.5090

MARITIME

40.000	.02533	.1336
35.000	.03006	.1119
30.000	.04085	.1123
27.900	.04611	.1198
25.000	.05367	.1299
22.500	.06128	.1433
21.300	.06514	.1509
20.000	.07033	.1606
18.500	.07484	.1752
18.000	.07611	.1795
17.200	.07504	.1858
16.400	.06689	.1876
15.000	.05950	.1866
14.800	.05801	.1832
14.000	.05448	.1789
13.000	.04841	.1629
12.500	.04623	.1490
11.500	.05652	.1052
11.000	.07216	.08415
10.591	.09447	.06893
10.000	.13066	.06245
9.800	.14392	.06222
9.500	.16085	.06397
9.200	.17855	.07218
9.000	.19546	.07085
8.700	.1996	.07051
8.500	.1848	.07119
8.200	.1869	.06368
7.900	.1988	.05840
7.200	.2415	.05965
6.500	.2879	.07387
6.200	.28860	.14480
6.000	.19840	.14880
5.500	.34310	.03207
5.000	.40720	.03660
4.500	.45200	.04247
4.000	.52890	.01985
3.750	.56450	.01718
3.500	.5849	.04164
3.392	.5670	.08004
3.200	.4594	.2208
3.000	.3271	.3328
2.700	.4399	.09134
2.500	.6093	.02125
2.250	.6772	.01062
2.000	.7159	.01380
1.800	.7534	.00918
1.536	.7951	.01207
1.300	.8288	.01178
1.060	.8682	.01210
0.860	.9013	.01084
0.694	.9396	.00979
0.633	.9556	.00996
0.550	.9882	.01179
0.515	.9975	.01147
0.488	1.014	.01221
0.400	1.063	.01536
0.337	1.115	.01858
0.300	1.142	.02800
0.250	1.181	.05471
0.200	1.188	.13000

URBAN

40.000	.02037	.04028
35.000	.02248	.04112
30.000	.02422	.04226
27.900	.02601	.04246
25.000	.02861	.04374
22.500	.03099	.04552
21.300	.03174	.04672
20.000	.03298	.04691
18.500	.03376	.04505
18.000	.03499	.04595
17.200	.03524	.04911
16.400	.03457	.04751
15.000	.02985	.05336
14.800	.03275	.04621
14.000	.03713	.04476
13.000	.03961	.04562
12.500	.04083	.04604
11.500	.04350	.04842
11.000	.04448	.05040
10.591	.04350	.05493
10.000	.04365	.06037
9.800	.04400	.06161
9.500	.04262	.06802
9.200	.04181	.08090
9.000	.04414	.07610
8.700	.04318	.07416
8.500	.03205	.07916
8.200	.02199	.06560
7.900	.03118	.06030
7.200	.04271	.06784
6.500	.04496	.06393
6.200	.04634	.06492
6.000	.04767	.06472
5.500	.05386	.06613
5.000	.05900	.06824
4.500	.06220	.07474
4.000	.06939	.07739
3.750	.07258	.08063
3.500	.07500	.08610
3.392	.07383	.08966
3.200	.07572	.09346
3.000	.07209	.1075
2.700	.07286	.1288
2.500	.08758	.1157
2.250	.09833	.1240
2.000	.1129	.1339
1.800	.1302	.1524
1.536	.1670	.1753
1.300	.2131	.1964
1.060	.2896	.2256
0.860	.3828	.2545
0.694	.5110	.2866
0.633	.5706	.3027
0.550	.6657	.3343
0.515	.7167	.3482
0.488	.7567	.3591
0.400	.9171	.4036
0.337	1.065	.4417
0.300	1.154	.4792
0.250	1.251	.5474
0.200	1.287	.6767

BACKGR	STRATOSPHERIC	
40.000	0.	.00592
35.000	0.	.00580
30.000	.00001	.00632
27.900	.00001	.00680
25.000	.00001	.00836
22.500	.00002	.01530
21.300	.00002	.01617
20.000	.00002	.01343
18.500	.00004	.01710
18.000	.00005	.02321
17.200	.00005	.04147
16.400	.00004	.03661
15.000	.00005	.01948
14.800	.00005	.01890
14.000	.00007	.01860
13.000	.0001	.01930
12.500	.00013	.01962
11.500	.00026	.03549
11.000	.00026	.05710
10.591	.00027	.04041
10.000	.00049	.04971
9.800	.00067	.07256
9.500	.00071	.09987
9.200	.00058	.08722
9.000	.00072	.09217
8.700	.00096	.1264
8.500	.00095	.1448
8.200	.00077	.1463
7.900	.00050	.1182
7.200	.00019	.04437
6.500	.00055	.03283
6.200	.00088	.04304
6.000	.00105	.05263
5.500	.00103	.05704
5.000	.00145	.04132
4.500	.00242	.04522
4.000	.00403	.05391
3.750	.00510	.06019
3.500	.00623	.07917
3.392	.00627	.08300
3.200	.00600	.07671
3.000	.00632	.05878
2.700	.00931	.00403
2.500	.01560	.00289
2.250	.02570	.00157
2.000	.04055	.00128
1.800	.05817	.00064
1.536	.09972	.00020
1.300	.1642	.00002
1.060	.2886	.00000
0.860	.4685	.00000
0.694	.7063	.00000
0.633	.8224	.00000
0.550	1.000	.00000
0.515	1.087	.00000
0.488	1.150	.00000
0.400	1.376	.00000
0.337	1.515	.00000
0.300	1.555	.00000
0.250	1.553	.00000
0.200	1.487	.00000

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40.000	.00002	.01328
35.000	.00003	.01192
30.000	.00006	.01124
27.900	.00008	.01072
25.000	.00014	.01038
22.500	.00020	.01104
21.300	.00023	.01054
20.000	.00026	.01390
18.500	.00031	.01741
18.000	.00031	.01826
17.200	.00029	.01706
16.400	.00028	.01620
15.000	.00032	.01635
14.800	.00033	.01532
14.000	.00046	.01447
13.000	.00071	.01530
12.500	.00090	.01623
11.500	.00191	.02282
11.000	.00260	.02511
10.591	.00322	.02859
10.000	.00449	.03009
9.800	.00475	.02979
9.500	.00518	.02871
9.200	.00573	.02658
9.000	.00602	.02506
8.700	.00481	.02438
8.500	.00257	.02218
8.200	.00099	.02180
7.900	.00136	.01652
7.200	.00250	.01112
6.500	.00378	.00665
6.200	.00467	.00525
6.000	.00570	.00451
5.500	.00869	.00396
5.000	.01204	.00335
4.500	.01775	.00319
4.000	.02394	.00349
3.750	.02913	.00488
3.500	.03435	.00654
3.392	.03761	.00743
3.200	.04262	.00936
3.000	.05126	.00949
2.700	.06343	.00842
2.500	.07792	.00843
2.250	.1021	.00867
2.000	.1353	.01019
1.800	.1849	.01191
1.536	.2648	.01490
1.300	.3693	.01881
1.060	.5057	.02452
0.860	.6570	.03178
0.694	.8079	.04084
0.633	.8676	.04532
0.550	.9473	.05271
0.515	.9799	.05651
0.488	1.0030	.05962
0.400	1.0680	.07244
0.337	1.0950	.06497
0.300	1.0790	.1126
0.250	.9009	.2818
0.200	.7006	.4482

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16. Abstract A computer program is described which calculates line by line atmospheric transmission spectra on a microcomputer. Radiance calculations are not included. The program is written in Fortran and could be modified to run on a microcomputer other than the one on which it was implemented, except that the plotting routine would have to be replaced. The program is based on the 1978 Air Force Geophysics Laboratory LASER routine, and uses absorption line data from the 1982 update of the AFGL Atmospheric Absorption Line Parameters Compilation; segments of needed line data are taken from the tape, preprocessed into the form used by the program and stored on floppy disks. The program calculates transmission spectra over a variety of paths and spanning an arbitrarily chosen wavelength or frequency range; the path may be chosen to be horizontal, vertical or slanted at any zenith angle, and may extend between any two altitudes between 0 and 15 km. Aerosol effects are included, presently based on the 1976 Shettle and Fenn models, plus some molecular continuum effects and Rayleigh scattering.			
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