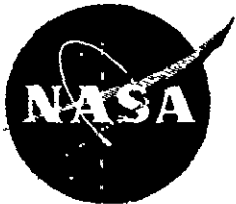


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# DELTA PAC ATTITUDE DETERMINATION AND ERROR PREDICTION

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AND ERROR PREDICTION

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## CONTENTS

	<u>Page</u>
INTRODUCTION .....	1
COORDINATE AXES DEFINITION .....	2
THE REFERENCE AXES TRANSFORMATION EQUATIONS .....	4
THE SUN SENSOR EQUATIONS .....	7
THE EARTH SENSOR EQUATIONS .....	14
THE ALTITUDE DETERMINATION EQUATIONS .....	22
EULER ANGLE TRANSFORMATIONS .....	27
PRESENTATION OF SPECIAL CASES .....	30
ERROR PROPAGATION .....	39
CONCLUSIONS .....	54
REFERENCES .....	55
APPENDIX .....	56

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## DELTA PAC ATTITUDE DETERMINATION AND ERROR PREDICTION

### INTRODUCTION

The Delta Packaged control spacecraft (Delta PAC) was an earth satellite designed, built, and tested in house by the Technology Directorate under the leadership of the System Division. The Delta PAC attitude determination and error prediction project was assigned to the author in Jan. 1969, and was completed in six months. It consists of both an analysis and a FORTRAN IV program. The computer program was utilized to evaluate the spacecraft attitude control system after it was launched into a circular orbit.

The Delta PAC control system contains one gimballed fly-wheel scanner with electronic damping that actively controls one axis and passively controls the other two axes. The spacecraft is therefore stabilized about all three axes. The mission objectives of this spacecraft is to test the earth stabilized control system for future spacecraft applications. The evaluation of the behavior of this spacecraft and the flight performance of its control system depends upon a knowledge of the PAC attitude. Therefore, the problem of attitude determination is of great importance.

This report presents the analysis of that project with the FORTRAN IV as its appendix. The analysis includes:

- (1) Formulation of attitude determination equations: An approach is developed in which the complete sensor data are used to determine two independent vectors, such as the local vertical and the sun line. These vectors are then applied to obtain the attitude of the spacecraft.
- (2) Presentation of special cases: Various cases arising from incomplete sensor output have been considered.
- (3) Examination of error propagation: Since measurement errors could exist in the sensor data, each calculation is subjected to errors. The effect of the errors on the determination of orientation of Delta PAC caused by errors of the measurement quantities is also explored.

### Coordinate Axes Definitions

For convenience, it is necessary to make the following definitions.

- (1) First Reference Axes.-The first coordinate system  $X_s, Y_s, Z_s$  has center at the spacecraft's noon position where

$X_s$  axis is along the spacecraft's velocity vector.

$Z_s$  axis is along the negative sun line vector.

$Y_s$  axis is determined by

$$\vec{Y}_s = \vec{Z}_s \times \vec{X}_s$$

- (2) Second Reference Axes.-The second coordinate system  $X_0, Y_0, Z_0$  has the same center as the first coordinate system with

$$\vec{X}_0 = \vec{X}_s$$

$\vec{Y}_0$  = normal to the orbit plane

$\vec{Z}_0$  = the vector pointing downward toward the local vertical.

- (3) Orbital Reference Axes.-The third coordinate system X, Y, Z has its center at the instantaneous location of the spacecraft with

$\vec{X}$  = spacecraft's velocity vector

$\vec{Y} = \vec{Y}_0$

$\vec{Z}$  = the downward local vertical

- (4) The Spacecraft Body Axes.-The spacecraft body coordinate system x, y, z is a set of right handed mutually orthogonal axes fixed in the spacecraft with their origin at the center of mass. It is required that the attitude control system maintain the alignment of the spacecraft body axes with the orbital axes system. That is, the yaw axis (z) is pointed toward the downward vertical and the pitch axis (y) is aligned with the normal to the orbit plane, thus causing the roll axis (x) to align with the spacecraft's velocity vector. Any attitude error is detected by a control subsystem which consists of an earth sensor and three sun sensors.

- (5) The Sun Sensor Axes System.-There are three sun sensors mounted at 120 degree intervals on the top surface of the Delta PAC. Each sun sensor head i has its own right handed system  $e_i, f_i, g_i$  which is oriented such that

(a) The axis  $g_i$  is in the x-y plane

(b) The angle between  $g_i$  and y is  $90^\circ + \xi_i$

(c) The angle between  $g_1$  and  $x$  is  $180^\circ + \xi_1$

(d) The angle between  $e_1$  and  $z$  is  $90^\circ + \eta_1$

The nominal constant  $\xi_i$  and  $\eta_i$  are:

$$\xi_1 = 0, \xi_2 = 120^\circ, \xi_3 = 240^\circ$$

$$\eta_1 = \eta_2 = \eta_3 = 26^\circ$$

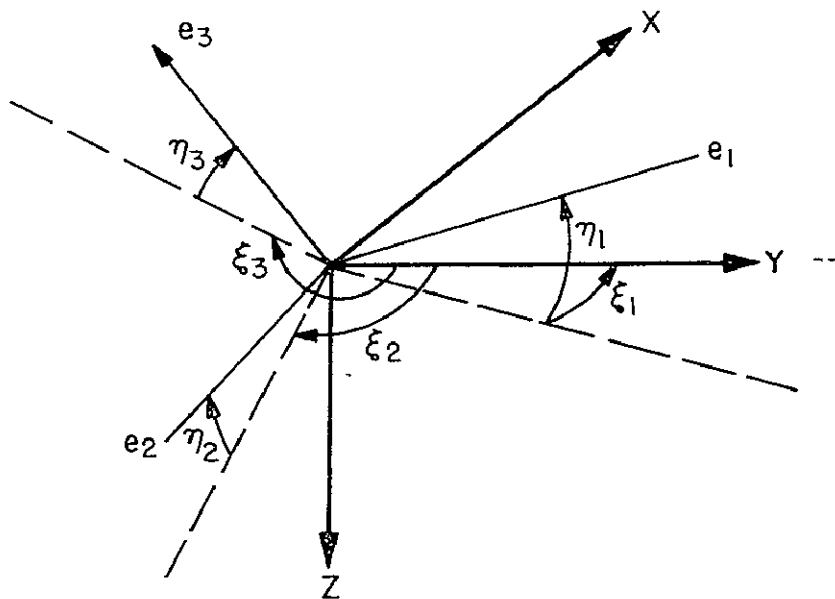


Figure 1 Sun Sensor Zenith Axes  $e_1, e_2, e_3$

### The Reference Axes Transformation Equations

If

$\alpha$  = the angle between  $\vec{Z}_0$  and  $\vec{Z}$

$\beta$  = the angle between orbit plane and the earth-sun line with positive sign when the sun is above the  $\vec{X}_0 - \vec{Z}_0$  plane (see Figure 2), the relationship between the first three reference axes can be expressed as

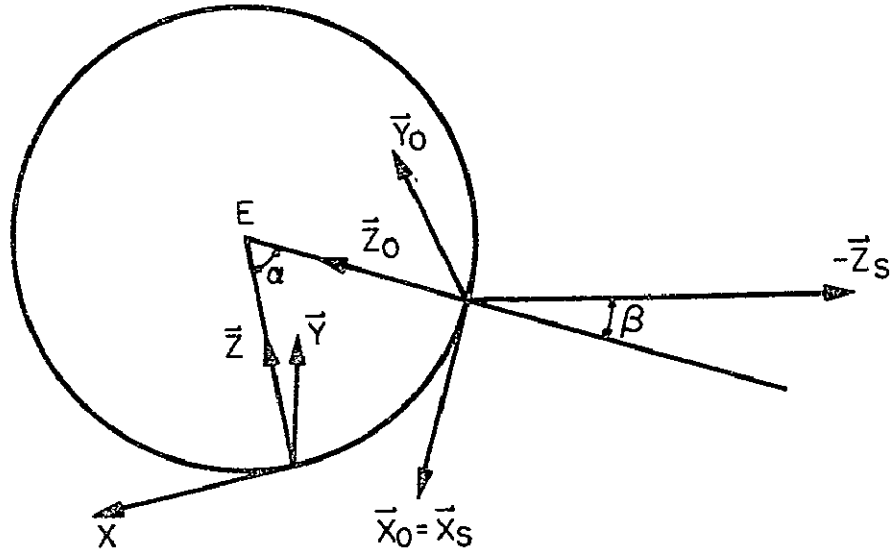


Figure 2. The orbital Reference Axes

$$\begin{aligned}
 \begin{bmatrix} \vec{X} \\ \vec{Y} \\ \vec{Z} \end{bmatrix} &= \begin{bmatrix} \cos \alpha & 0 & \sin \alpha \\ 0 & 1 & 0 \\ -\sin \alpha & 0 & \cos \alpha \end{bmatrix} \begin{bmatrix} \vec{X}_0 \\ \vec{Y}_0 \\ \vec{Z}_0 \end{bmatrix} \\
 &= \begin{bmatrix} \cos \alpha & 0 & \sin \alpha \\ 0 & 1 & 0 \\ -\sin \alpha & 0 & \cos \alpha \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \beta & -\sin \beta \\ 0 & \sin \beta & \cos \beta \end{bmatrix} \begin{bmatrix} \vec{X}_s \\ \vec{Y}_s \\ \vec{Z}_s \end{bmatrix} \quad (T1) \\
 &= \begin{bmatrix} \cos \alpha & \sin \alpha \sin \beta & \sin \alpha \cos \beta \\ 0 & \cos \beta & -\sin \beta \\ -\sin \alpha & \sin \beta \cos \alpha & \cos \alpha \cos \beta \end{bmatrix} \begin{bmatrix} \vec{X}_s \\ \vec{Y}_s \\ \vec{Z}_s \end{bmatrix}
 \end{aligned}$$

The range of  $\beta$  is  $-55^\circ \leq \beta \leq 55^\circ$  for Delta PAC calculated from orbit determination.



Let  $\vec{S}$  be the sun line vector, the components of  $\vec{S}$  in each of three reference coordinate systems can therefore be determined. It is obvious that

$$\vec{S} = 0 \cdot \vec{X}_s + 0 \cdot \vec{Y}_s - \vec{Z}_s$$

in the  $X_s, Y_s, Z_s$  system, and

$$\vec{S} = \sin \beta \vec{Y}_0 - \cos \beta \vec{Z}_0$$

in the  $X_0, Y_0, Z_0$  system.

Since the unit sun line vector has the coordinate (0,0,-1) in the  $X_s, Y_s, Z_s$  system, by using equation (T1) the components of sun line in the X, Y, Z, system are:

$$[S]_0 = \begin{bmatrix} S_1 \\ S_2 \\ S_3 \end{bmatrix} = \begin{bmatrix} -\sin a \cos \beta \\ \sin \beta \\ -\cos a \cos \beta \end{bmatrix} \quad (T2)$$

Therefore

$$\vec{S} = -\sin a \cos \beta \cdot \vec{X} + \sin \beta \cdot \vec{Y} - \cos a \cos \beta \cdot \vec{Z} \quad (T3)$$

and

$$\vec{S} \cdot \vec{Z} = -\cos a \cos \beta = S_3 \quad (T4)$$

If there is no attitude error, the body axes x, y, z should be coincident with the orbital axes X, Y, Z respectively and equation (T2) are the direction cosines of sun line in both axes systems. However, this is no longer true if a

misorientation takes place. For that case, the components of sun line in the body axes need to be determined and the sun sensors are used for this purpose.

### The Sun Sensor Equations

Let  $S_x, S_y, S_z$  be the components of the sun line vector  $\vec{S}$  in the body axes system. Then

$$\vec{S} = S_x \vec{x} + S_y \vec{y} + S_z \vec{z} \quad (S1)$$

The method of determining the values of  $S_x, S_y,$  and  $S_z$  is described as follows.

The body axes  $x, y, z$  can be aligned with the sun sensor axes  $e_1, f_1, g_1$  by the sequence of three rotations: a first rotation  $R_z (90^\circ + \xi_1)$  about axis  $z$  through an angle  $90^\circ + \xi_1$ , a second rotation  $R_y (\eta_1)$  about displaced  $y$  axis through an angle  $\eta_1$  and a third rotation  $R_x (-90^\circ)$  about displaced  $x$  axis through an angle  $-90^\circ$  (See Figure 3). The transformation may be written in the following matrix form.

$$\begin{bmatrix} \vec{e}_1 \\ \vec{f}_1 \\ \vec{g}_1 \end{bmatrix} = R_x (-90^\circ) R_y (\eta_1) R_z (90^\circ + \xi_1) \begin{bmatrix} \vec{x} \\ \vec{y} \\ \vec{z} \end{bmatrix} \quad (S2)$$

$$= \begin{bmatrix} 1 & 0 & 0 \\ 0 & 0 & -1 \\ 0 & 1 & 0 \end{bmatrix} \begin{bmatrix} \cos \eta_1 & 0 & -\sin \eta_1 \\ 0 & 1 & 0 \\ \sin \eta_1 & 0 & \cos \eta_1 \end{bmatrix} \begin{bmatrix} -\sin \xi_1 & \cos \xi_1 & 0 \\ -\cos \xi_1 & -\sin \xi_1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} \vec{x} \\ \vec{y} \\ \vec{z} \end{bmatrix}$$

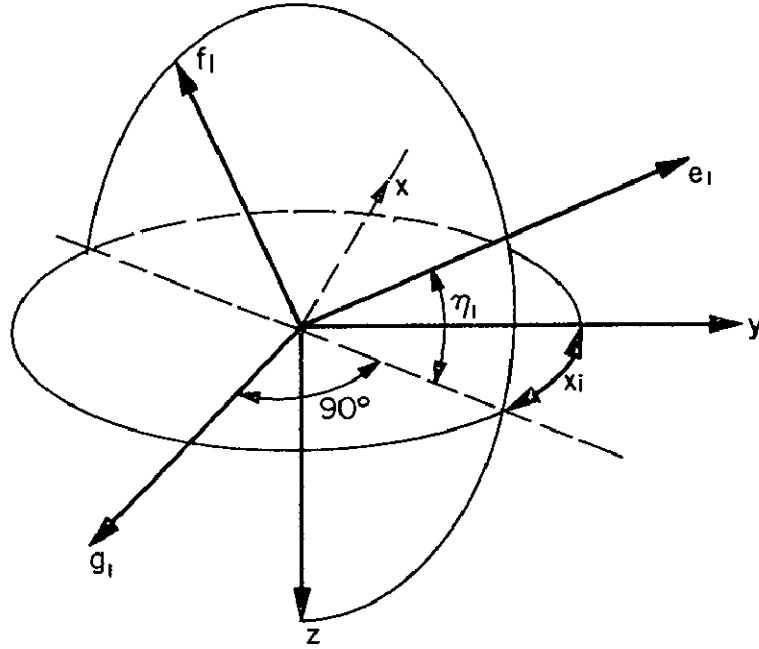


Figure 3 Sensor Frame System  $e_1, f_1, g_1$

$$= \begin{bmatrix} -\cos \eta_1 \sin \xi_1 & \cos \eta_1 \cos \xi_1 & -\sin \eta_1 \\ \sin \eta_1 \sin \xi_1 & -\sin \eta_1 \cos \xi_1 & -\cos \eta_1 \\ -\cos \xi_1 & -\sin \xi_1 & 0 \end{bmatrix} \begin{bmatrix} \vec{x} \\ \vec{y} \\ \vec{z} \end{bmatrix} \quad (\text{S3})$$

In simple notation, that is

$$\begin{bmatrix} \vec{e}_1 \\ \vec{f}_1 \\ \vec{g}_1 \end{bmatrix} = [D] \begin{bmatrix} \vec{x} \\ \vec{y} \\ \vec{z} \end{bmatrix} \quad (\text{S4})$$

Where

$$[D] = R_x(-90) R_y(\eta_1) R_z(90^\circ + \xi_1)$$

Since  $[D]$  is an orthogonal transformation matrix, therefore

$$[D]^{-1} = [D]^T = \begin{bmatrix} -\cos \eta_1 \sin \xi_1 & \sin \eta_1 \sin \xi_1 & -\cos \xi_1 \\ \cos \eta_1 \cos \xi_1 & -\sin \eta_1 \cos \xi_1 & -\sin \xi_1 \\ -\sin \eta_1 & -\cos \eta_1 & 0 \end{bmatrix} \quad (S5)$$

and

$$[D]^{-1} \begin{bmatrix} \vec{e}_1 \\ \vec{f}_1 \\ \vec{g}_1 \end{bmatrix} = [D]^{-1} [D] \begin{bmatrix} \vec{x} \\ \vec{y} \\ \vec{z} \end{bmatrix}$$

or

$$\begin{bmatrix} \vec{x} \\ \vec{y} \\ \vec{z} \end{bmatrix} = [D]^T \begin{bmatrix} \vec{e}_1 \\ \vec{f}_1 \\ \vec{g}_1 \end{bmatrix} = \begin{bmatrix} -\cos \eta_1 \sin \xi_1 & \sin \eta_1 \sin \xi_1 & -\cos \xi_1 \\ \cos \eta_1 \cos \xi_1 & -\sin \eta_1 \cos \xi_1 & -\sin \xi_1 \\ -\sin \eta_1 & -\cos \eta_1 & 0 \end{bmatrix} \begin{bmatrix} \vec{e}_1 \\ \vec{f}_1 \\ \vec{g}_1 \end{bmatrix} \quad (S6)$$

The sensor telemetry yields  $i$  ( $i = 1, 2, \text{ or } 3$ ) which indicates that the  $i$  th sensor receives the greatest sun intensity and is selected to produce the output. The output of the selected sensor is two angles  $\alpha_1$  and  $\beta_1$  which define two planes, each containing the sun (Figure 4). The intersection of these two planes determines the sun-line.

Let  $S_{11}$ ,  $S_{21}$  and  $S_{31}$  be the direction cosines of sun-line OS with respect to the  $i$  th sun sensor system (see Figure 4). Then:

$$\vec{S} = S_{11} \vec{e}_1 + S_{21} \vec{f}_1 + S_{31} \vec{g}_1 \quad (S7)$$

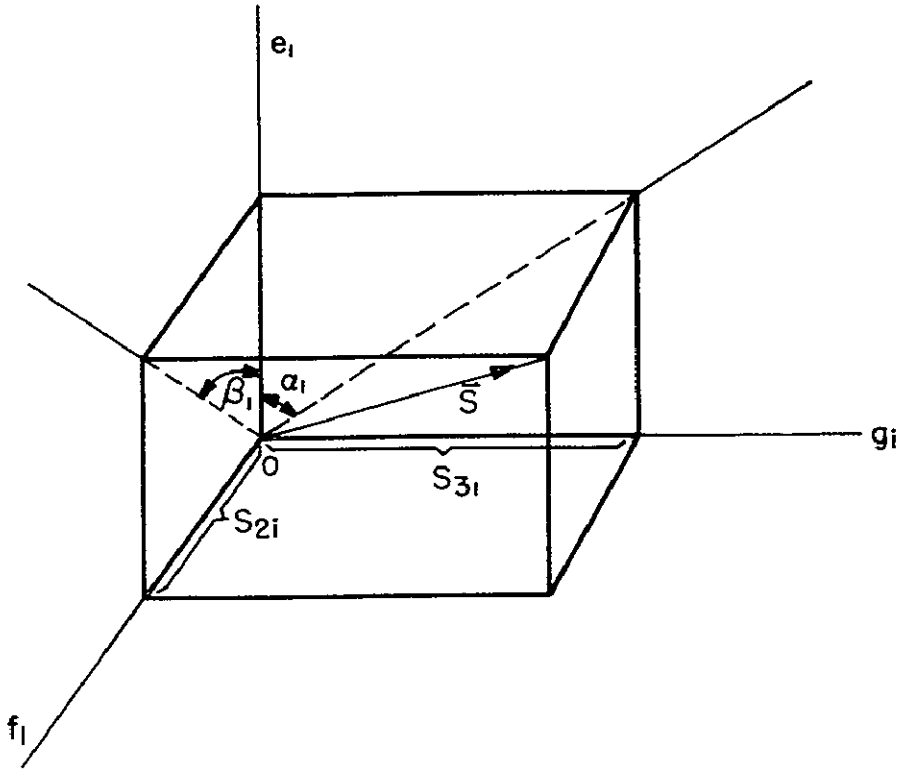


Figure 4 Sun line in  $e_1, f_1, g_1$  system

Since:

$$\left. \begin{aligned} \tan \alpha_1 &= S_{31}/S_{11} \\ \tan \beta_1 &= S_{21}/S_{11} \end{aligned} \right\} \quad (S8)$$

and

$$S_{11}^2 + S_{21}^2 + S_{31}^2 = 1$$

giving

$$\left. \begin{aligned} S_{11} &= 1 / \sqrt{1 + \tan^2 \alpha_1 + \tan^2 \beta_1} \\ S_{21} &= \tan \beta_1 / \sqrt{1 + \tan^2 \alpha_1 + \tan^2 \beta_1} \\ S_{31} &= \tan \alpha_1 / \sqrt{1 + \tan^2 \alpha_1 + \tan^2 \beta_1} \end{aligned} \right\} \quad (S9)$$

Using equation (S7) and equation (S3) or equation (S6) and equation (S7), the components of the sun vector  $\vec{S}$  in the body axes system can be found. That is

$$\begin{bmatrix} S_x \\ S_y \\ S_z \end{bmatrix} = \begin{bmatrix} -\cos \eta_1 \sin \xi_1 & \sin \eta_1 \sin \xi_1 & -\cos \xi_1 \\ \cos \eta_1 \cos \xi_1 & -\sin \eta_1 \cos \xi_1 & -\sin \xi_1 \\ -\sin \eta_1 & -\cos \eta_1 & 0 \end{bmatrix} \begin{bmatrix} S_{11} \\ S_{21} \\ S_{31} \end{bmatrix} \quad (\text{S10})$$

In terms of known quantities, equation (S10) becomes

$$[S]_b = \begin{bmatrix} S_x \\ S_y \\ S_z \end{bmatrix} = \frac{1}{\sqrt{1 + \tan^2 \alpha_1 + \tan^2 \beta_1}} \begin{bmatrix} -\cos \eta_1 \sin \xi_1 & \sin \eta_1 \sin \xi_1 & -\sin \xi_1 \\ \cos \eta_1 \cos \xi_1 & -\sin \eta_1 \cos \xi_1 & -\sin \xi_1 \\ -\sin \eta_1 & -\cos \eta_1 & 0 \end{bmatrix} \begin{bmatrix} 1 \\ \tan \beta_1 \\ \tan \alpha_1 \end{bmatrix} \quad (\text{S11})$$

This final form and equation (T2) are used to obtain the attitude determination equations.

### The Earth Sensor Equations

Since the local vertical  $\vec{Z}$  in the orbital axes system has components

$$[Z]_o = \begin{bmatrix} 0 \\ 0 \\ 1 \end{bmatrix} \quad (\text{E1})$$

the same vector  $\vec{Z}$  in the body axes system can be written as

$$\vec{Z} = a_{13} \vec{x} + a_{23} \vec{y} + a_{33} \vec{z} \quad (\text{E2})$$

Here the components

$$[Z]_b = \begin{bmatrix} a_{13} \\ a_{23} \\ a_{33} \end{bmatrix} \quad (E3)$$

are to be determined by the measured quantities obtained from the earth sensor.

For Delta PAC there is one earth sensor, which is a horizon scanner. The scanner is mounted on a gimballed platform which rotates about the spacecraft roll axes (x). Eddy currents induced by the motion of the gimbal with respect to the spacecraft provide the means for damping the spacecraft roll-yaw motion. The scanner cone axis ( $Y_w$ ) coincides with body pitch axis (y) when the gimbal angle  $\gamma$  is zero (see Figure 5). The range of  $\gamma$  is:  $-30^\circ \leq \gamma \leq 30^\circ$ .

In general, the scan pattern of a single scanner is a cone. As the scan cone cuts the earth, a voltage earth pulse of width  $2\rho$  is emitted by a sensing element called a bolometer. The width of this pulse represents the angular portion of the cone which intersects the earth (see Figure 6).

Let an intermediate parameter  $C_{23}$  be calculated first, where

$$C_{23} = \text{the scalar product of two unit vectors which represent the scan cone axis } \vec{Y}_w \text{ and the local vertical } \vec{Z}.$$

The equation expressing the relationship between the variable  $C_{23}$  and the earth pulse half width  $\rho$  are presented as follows:

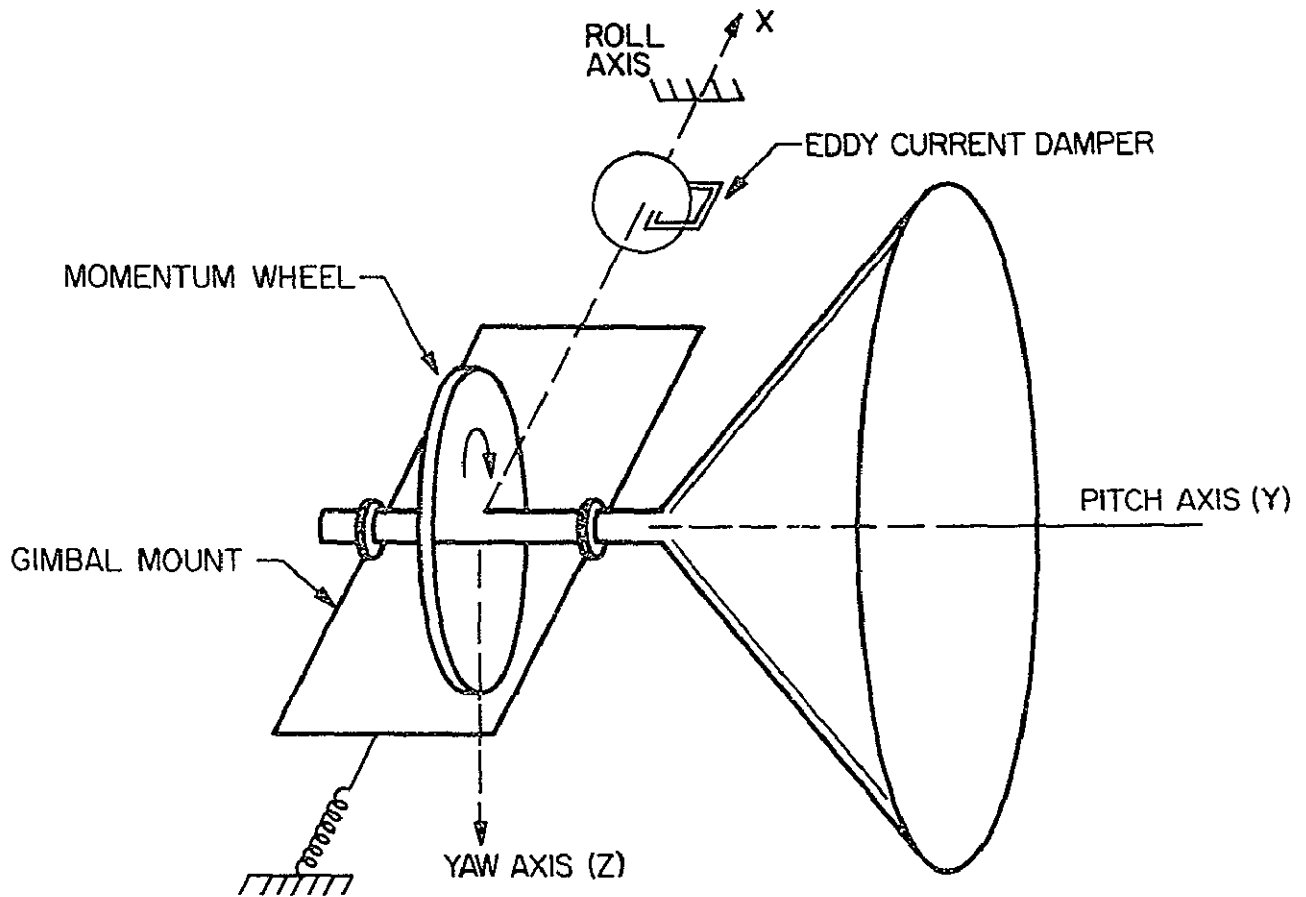


Figure 5 The Horizon Scanner

$$\cos \rho = \frac{\cos \alpha_e - C_{23} \cos \sigma}{\sqrt{1 - C_{23}^2 \sin^2 \sigma}} \quad (E4)$$

with

$\sigma$  = the apex half-angle of the cone

$\alpha_e$  = one-half the angle subtended by the earth.

The value of  $\alpha_e$  is calculated from:

$$\sin \alpha_e = \frac{R_e}{R_e + h}$$



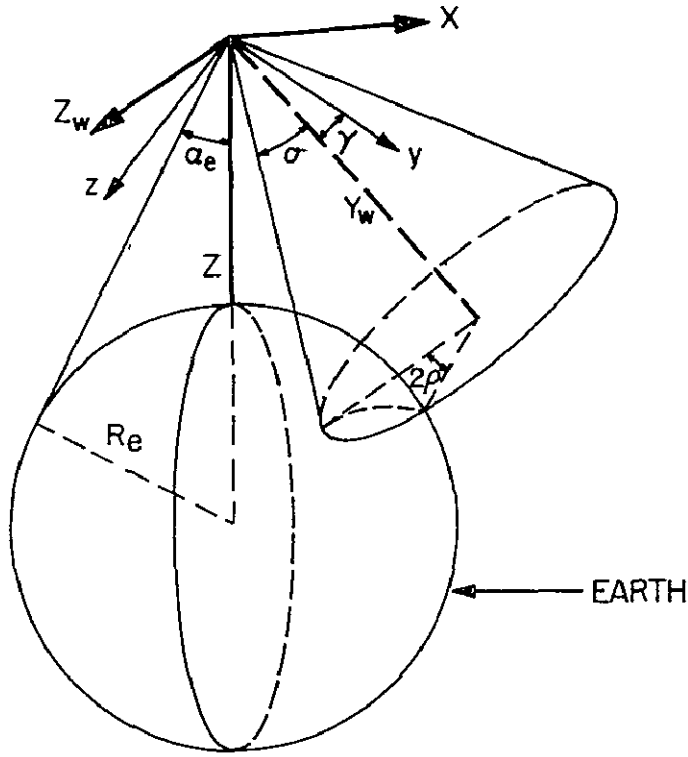


Figure 6. The Scanner Geometry

where

$R_e$  = the earth's radius

$h$  = the orbital altitude

The earth pulse half width  $\rho$  is in the range  $(0, 180)$ . If the scanner does not see the earth,  $\rho$  is then zero and the variable  $C_{23}$  is less than  $\cos(\alpha_e + \sigma)$ . If the scan cone is completely on the earth,  $\rho$  is  $180^\circ$  and  $C_{23}$  is larger than  $\cos(\alpha_e - \sigma)$ . The variable  $C_{23}$  is computed from equation (E4) only when  $\rho$  lies between  $0^\circ$  and  $180^\circ$ . In this case, the result of derivation gives:

$$C_{23} = \frac{\cos \alpha_e \cos \sigma \pm \cos \rho \sin \sigma \sqrt{\sin^2 \alpha_e - \sin^2 \sigma \sin^2 \rho}}{1 - \sin^2 \sigma \sin^2 \rho} \quad (\text{E5})$$

Based on the fact that  $C_{23}$  and  $\cos \rho$  vary inversely, it can be shown that the positive sign in the numerator should be discarded, i.e.,

$$C_{23} = \frac{\cos a_e \cos \sigma - \cos \rho \sin \sigma \sqrt{\sin^2 a_e - \sin^2 \sigma \sin^2 \rho}}{1 - \sin^2 \sigma \sin^2 \rho} \quad (\text{E6})$$

for  $0 < \rho < 180$

The unit vector representing the scan cone axis can be written as

$$\vec{Y}_w = \cos \gamma \vec{y} + \sin \gamma \vec{z}$$

where  $\gamma$  is the gimbal angle. Therefore

$$C_{23} = \vec{Z} \cdot \vec{Y}_w = a_{23} \cos \gamma + a_{33} \sin \gamma \quad (\text{E7})$$

or

$$a_{33} \sin \gamma = C_{23} - a_{23} \cos \gamma \quad (\text{E8})$$

Suppose that a unit vector  $\vec{Z}_w$  is defined such that in conjunction with  $\vec{X}_w = \vec{x}$  (spacecraft roll axis) and the scanner cone axis  $\vec{Y}_w$ , a right handed orthogonal coordinate system is formed. Then

$$\vec{Z}_w = \vec{x} \times \vec{Y}_w = -\sin \gamma \vec{y} + \cos \gamma \vec{z}$$

giving

$$\vec{Z} \cdot \vec{Z}_w = a_{33} \cos \gamma - a_{23} \sin \gamma$$

and

$$\vec{Z} \cdot \vec{X}_w = \vec{Z} \cdot \vec{x} = a_{13}$$

Define the pitch computed angle  $\theta_s$  measured about the scanner axis  $Y_w$  as follows

$$\tan \theta_s = \frac{-\vec{Z} \cdot \vec{X}_w}{\vec{Z} \cdot \vec{Z}_w} = \frac{-a_{13}}{a_{33} \cos \gamma - a_{23} \sin \gamma} \quad (\text{E9})$$

This implies

$$\sin \theta_s = -a_{13} \cdot V$$

and

$$\cos \theta_s = (a_{33} \cos \gamma - a_{23} \sin \gamma) \cdot V$$

The variable  $V$ , after use of:

$$\sin^2 \theta_s + \cos^2 \theta_s = 1$$

and equation (E7) gives:

$$V = \pm 1/\sqrt{1 - C_{23}^2}$$

Hence,

$$a_{13} = \pm \sqrt{1 - C_{23}^2} \sin \theta_s$$

and

$$a_{33} \cos \gamma - a_{23} \sin \gamma = \pm \sqrt{1 - C_{23}^2} \cos \theta_s \quad (\text{E10})$$

Solving equation (E7) and equation (E10), gives:

$$\left. \begin{aligned} a_{23} &= C_{23} \cos \gamma \pm \sqrt{1 - C_{23}^2} \sin \gamma \cos \theta_s \\ a_{33} &= C_{23} \sin \gamma \mp \sqrt{1 - C_{23}^2} \cos \gamma \cos \theta_s \end{aligned} \right\} \quad (\text{E11})$$

In addition to providing a measure of the angle  $\rho$ , the earth pulse is used to obtain  $\theta_s$  through the use of signal processing which is described in the following paragraph:

When the scanner passes a reference mark in the yaw-pitch plane of the spacecraft, a reference pulse is generated by a magnetic wedge attached to the pitch wheel. By comparing the reference pulse with the center of the earth pulse, a measure of the angle  $\theta_s$  is obtained. For  $180^\circ$  of rotation after the reference pulse, the signal processor causes any portion of the earth pulse to contribute to  $e$  in a positive sense. For the remaining  $180^\circ$ , of rotation, the earth pulse contributes in a negative sense. If the earth pulse contains the reference mark but not the  $180$  mark, then it does indeed measure  $\theta_s$ . If these conditions are not met, the exact relation is summarized mathematically by the following equation:

$$e = (\text{sign } \theta_s) \text{Min} \{ |\theta_s|, |180^\circ - \theta_s|, \rho, 180^\circ - \rho, E \} \quad (\text{E12})$$

where

$E = 45$  is the electronic saturation angle

$e$  = telementered scanner pitch signal.

The plot of  $e$  vs.  $\theta_s$  is shown in Figure 7.

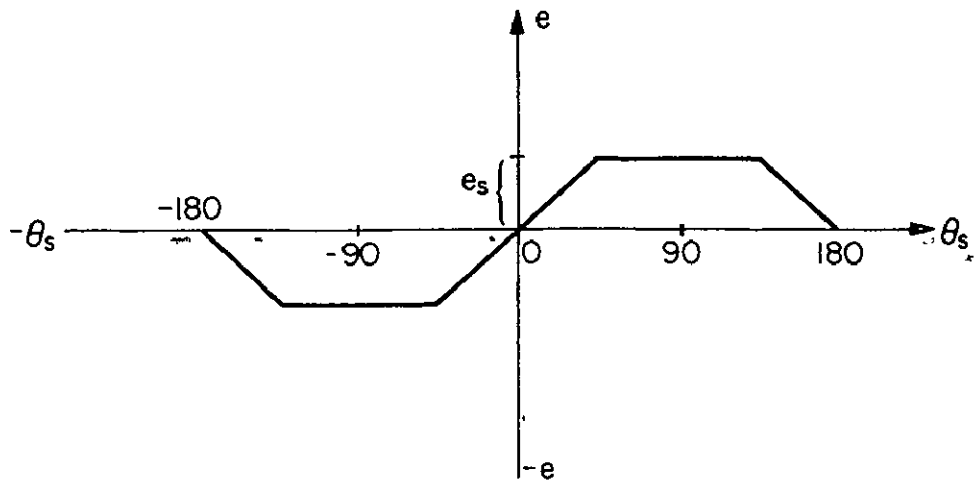


Figure 7 The Error Signal  $e$

For convenience, a new variable,  $e_s$ , is defined and equal to:

$$e_s = \text{Min} (\rho, 180^\circ - \rho, E) \quad (\text{E13})$$

Accordingly, the case:

$$|e| < e_s$$

and

$$|e| = e_s$$

are to be discussed.

(1) When the absolute value of  $e$  is less than  $e_s$ , there are possibilities that:

$$\theta_s = e$$

or

$$\theta_s = \pm (180^\circ - |e|) \quad (\text{E14})$$

Care must be taken to consider all these possible cases. Four sets of equations are thus obtained to represent  $[Z]_b$ . They are

$$\left. \begin{aligned}
 \begin{bmatrix} a_{13} \\ a_{23} \\ a_{33} \end{bmatrix} &= \begin{bmatrix} \sqrt{1 - C_{23}^2} \sin e \\ C_{23} \cos \gamma + \sqrt{1 - C_{23}^2} \sin \gamma \cos e \\ C_{23} \sin \gamma - \sqrt{1 - C_{23}^2} \cos \gamma \cos e \end{bmatrix} \\
 \begin{bmatrix} a_{13} \\ a_{23} \\ a_{33} \end{bmatrix} &= \begin{bmatrix} -\sqrt{1 - C_{23}^2} \sin e \\ C_{23} \cos \gamma + \sqrt{1 - C_{23}^2} \sin \gamma \cos e \\ C_{23} \sin \gamma - \sqrt{1 - C_{23}^2} \cos \gamma \cos e \end{bmatrix} \\
 \begin{bmatrix} a_{13} \\ a_{23} \\ a_{33} \end{bmatrix} &= \begin{bmatrix} \sqrt{1 - C_{23}^2} \sin e \\ C_{23} \cos \gamma - \sqrt{1 - C_{23}^2} \sin \gamma \cos e \\ C_{23} \sin \gamma + \sqrt{1 - C_{23}^2} \cos \gamma \cos e \end{bmatrix} \\
 \begin{bmatrix} a_{13} \\ a_{23} \\ a_{33} \end{bmatrix} &= \begin{bmatrix} -\sqrt{1 - C_{23}^2} \sin e \\ C_{23} \cos \gamma - \sqrt{1 - C_{23}^2} \sin \gamma \cos e \\ C_{23} \sin \gamma + \sqrt{1 - C_{23}^2} \cos \gamma \cos e \end{bmatrix}
 \end{aligned} \right\} \quad (E15)$$

In order to determine which of these four sets is the correct local vertical, note that

$$[\vec{S} \cdot \vec{Z}]_0 = -\cos \alpha \cos \beta$$

Since the scalar product of two vectors has the same value regardless of what coordinate axis system is used, therefore

$$[\vec{S} \cdot \vec{Z}]_b = S_x a_{13} + S_y a_{23} + S_z a_{33} = -\cos \alpha \cos \beta \quad (E16)$$

The representation in equation (E15) which best satisfies this criterion is the proper expression.

(2) When the absolute value of  $e$  is equal to  $e_s$ ,  $e$  is saturated. The telemetered error signal  $e$  fails to have a definite value and equation (E14) is no longer useful in determining  $\theta_s$ . However it is still possible to find  $[Z]_b$  by making use of  $\rho$  and  $[S]_b$ . The parameter  $C_{23}$  determined from  $\rho$  in equation (E6) is the cosine of the angle between the sensor cone axis  $\vec{Y}_w$  and the local vertical  $\vec{Z}$ . Thus the local vertical must be on a cone with axis  $\vec{Y}_w$ . The cosine of the angle between  $\vec{Z}$  and  $\vec{S}$  is  $-\cos \alpha \cos \beta$  and thus  $\vec{Z}$  must lie on a cone with axis  $\vec{S}$ . The determination of the intersection of these two cones is accomplished by solving equation (E16) with equation (E7) to obtain  $a_{23}$  and  $a_{33}$  in terms of  $a_{13}$ . This gives

$$\left. \begin{aligned} a_{23} &= \frac{-\sin \gamma (\cos \alpha \cos \beta + S_x a_{13}) - C_{23} S_z}{S_y \sin \gamma - S_z \cos \gamma} \\ a_{33} &= \frac{\cos \gamma (\cos \alpha \cos \beta + S_x a_{13}) + C_{23} S_y}{S_y \sin \gamma - S_z \cos \gamma} \end{aligned} \right\} \quad (E17)$$

Assuming, of course, that:

$$S_y \sin \gamma - S_z \cos \gamma \neq 0$$

If both equations in (E17) are squared and are substituted in the following:

$$a_{23}^2 + a_{33}^2 = 1 - a_{13}^2 \quad (E18)$$

a quadratic equation of  $a_{13}$  is obtained. That is,

$$E a_{13}^2 + F a_{13} + G = 0 \quad (\text{E19})$$

where

$$E = S_x^2 + (S_y \sin \gamma - S_z \cos \gamma)^2$$

$$F = 2 S_x (\cos \alpha \cos \beta + C_{23} (S_z \sin \gamma + S_y \cos \gamma))$$

and

$$G = - (S_y \sin \gamma - S_z \cos \gamma)^2 + 2 C_{23} \cos \alpha \cos \beta (S_z \sin \gamma + S_y \cos \gamma) \\ + \cos^2 \alpha \cos^2 \beta + C_{23}^2 (S_y^2 + S_z^2)$$

Solutions of the resulting equation for  $a_{13}$  are then substituted in equation (E17)

to obtain  $a_{23}$  and  $a_{33}$ .

In case when

$$\gamma \neq 0$$

$$S_y \sin \gamma - S_z \cos \gamma = 0$$

that is, when the sun is in the  $\vec{x}, \vec{Y}_w$  plane, then the above expression (E17) can

not be used. Since

$$\frac{S_y}{\cos \gamma} = \frac{S_z}{\sin \gamma} = \frac{S_y a_{23} + S_z a_{33}}{a_{23} \cos \gamma + a_{33} \sin \gamma}$$

or

$$\frac{S_y}{\cos \gamma} = \frac{-\cos \alpha \cos \beta - S_x a_{13}}{C_{23}} \quad (\text{E20})$$



Consequently the following equations apply

$$a'_{13} = \frac{-\cos \alpha \cos \beta \cos \gamma - S_y C_{23}}{S_x \cos \gamma}$$

$$\left. \begin{aligned} a_{23} &= C_{23} \cos \gamma \pm \tan \gamma \sqrt{(1 - C_{23}^2) S_x^2 \cos^2 \gamma - (\cos \alpha \cos \beta \cos \gamma + S_y C_{23})^2 / S_x} \\ a_{33} &= C_{23} \sin \gamma \mp \sqrt{(1 - C_{23}^2) S_x^2 \cos^2 \gamma - (\cos \alpha \cos \beta \cos \gamma + S_y C_{23})^2 / S_x} \end{aligned} \right\} \text{(E21)}$$

which apply except where  $S_x = 0$  that is when the sun lies on the sensor cone axis. Note that if  $\theta_s$  can not be determined, then in general two solutions are obtained, from which selection of the correct solution can not be made without additional information.

### The Attitude Determination Equations

At this point the direction cosines of the local vertical and the sun line with respect to the body axes system are known. They can be utilized to determine the spacecraft's attitude.

Suppose that the transformation matrix taking the orbital axes system into the body axes system is:

$$\begin{aligned} \vec{x} &= A_{11} \vec{X} + A_{12} \vec{Y} + A_{13} \vec{Z} \\ \vec{y} &= A_{21} \vec{X} + A_{22} \vec{Y} + A_{23} \vec{Z} \\ \vec{z} &= A_{31} \vec{X} + A_{32} \vec{Y} + A_{33} \vec{Z} \end{aligned} \quad \text{(A1)}$$

The coefficient  $A_{ij}$  ( $i, j=1, 2, 3$ ) in equation (A1) defines a matrix

$$[A] = \begin{bmatrix} A_{11} & A_{12} & A_{13} \\ A_{21} & A_{22} & A_{23} \\ A_{31} & A_{32} & A_{33} \end{bmatrix} \quad (A2)$$

Equation (A1) implies that any vector  $V$  in the  $X, Y, Z$  system, written in column form, may be transformed into  $x, y, z$  system by the application of the matrix  $[A]$ . That is

$$\begin{bmatrix} v_x \\ v_y \\ v_z \end{bmatrix} = [A] \begin{bmatrix} V_1 \\ V_2 \\ V_3 \end{bmatrix} \quad (A3)$$

where

$$[V]_b = \begin{bmatrix} v_x \\ v_y \\ v_z \end{bmatrix} \quad \text{and} \quad [V]_0 = \begin{bmatrix} V_1 \\ V_2 \\ V_3 \end{bmatrix}$$

are the components of  $\vec{V}$  in the body axes system and in the orbital axes system respectively.

In order to determine the direction cosine matrix  $[A]$  it is assumed that  $[S]_0$  and  $[Z]_0$  are known and that  $[S]_b$  and  $[Z]_b$  have been determined. By defining

$$\vec{F} = \vec{Z} \times \vec{S} \quad (A4)$$

the column matrix  $[F]_0$  and  $[F]_b$  can be obtained

$$[F]_0 = \begin{bmatrix} F_1 \\ F_2 \\ F_3 \end{bmatrix} = \begin{bmatrix} -S_2 \\ S_1 \\ 0 \end{bmatrix} \quad (A5)$$

$$[F]_b = \begin{bmatrix} F_x \\ F_y \\ F_z \end{bmatrix} = \begin{bmatrix} S_z a_{23} - S_y a_{33} \\ S_x a_{33} - S_z a_{13} \\ S_y a_{13} - S_x a_{23} \end{bmatrix} \quad (A6)$$

If equation (A3) is applied to the vectors  $\vec{Z}$ ,  $\vec{S}$  and  $\vec{F}$  respectively, it produces

$$\begin{bmatrix} a_{13} \\ a_{23} \\ a_{33} \end{bmatrix} = [A] \begin{bmatrix} 0 \\ 0 \\ 1 \end{bmatrix} \quad (A7)$$

$$\begin{bmatrix} S_x \\ S_y \\ S_z \end{bmatrix} = [A] \begin{bmatrix} S_1 \\ S_2 \\ S_3 \end{bmatrix} \quad (A8)$$

and

$$\begin{bmatrix} F_x \\ F_y \\ F_z \end{bmatrix} = [A] \begin{bmatrix} F_1 \\ F_2 \\ F_3 \end{bmatrix} \quad (A9)$$

Combining these three column matrix equations into a single square matrix equation, yields

$$\begin{bmatrix} a_{13} & S_x & F_x \\ a_{23} & S_y & F_y \\ a_{33} & S_z & F_z \end{bmatrix} = [A] \begin{bmatrix} 0 & S_1 & F_1 \\ 0 & S_2 & F_2 \\ 1 & S_3 & F_3 \end{bmatrix} \quad (\text{A10})$$

It may be written in this simple notation

$$[B] = [A] [Q] \quad (\text{A11})$$

Where

$$[Q] = \begin{bmatrix} 0 & S_1 & F_1 \\ 0 & S_2 & F_2 \\ 1 & S_3 & F_3 \end{bmatrix} = \begin{bmatrix} 0 & S_1 & -S_2 \\ 0 & S_2 & S_1 \\ 1 & S_3 & 0 \end{bmatrix} \quad (\text{A12})$$

and

$$[B] = \begin{bmatrix} a_{13} & S_x & F_x \\ a_{23} & S_y & F_y \\ a_{33} & S_z & F_z \end{bmatrix} = \begin{bmatrix} a_{33} & S_x & S_z a_{23} - S_y a_{33} \\ a_{13} & S_y & S_x a_{33} - S_z a_{13} \\ a_{23} & S_z & S_y a_{13} - S_x a_{23} \end{bmatrix} \quad (\text{A13})$$

The determinant of [Q] is

$$\begin{aligned} \det [Q] &= |Q| = S_1^2 + S_2^2 \\ &= 1 - S_3^2 \end{aligned} \quad (\text{A14})$$

If  $S_3 \neq \pm 1$ , then

$$|Q| \neq 0$$

The matrix [Q] is non-singular and possesses an inverse  $[Q]^{-1}$  where

$$[Q]^{-1} = \frac{1}{1 - S_3^2} \begin{bmatrix} -S_1 S_3 & -S_2 S_3 & 1 - S_3^2 \\ S_1 & S_2 & 0 \\ -S_2 & S_1 & 0 \end{bmatrix} \quad (A15)$$

By multiplying the matrix equation (A12) with  $[Q]^{-1}$ , gives

$$[B] [Q]^{-1} = [A] [Q] [Q]^{-1}$$

$$[A] = [B] [Q]^{-1} \quad (A16)$$

Therefore

$$\begin{bmatrix} A_{11} & A_{12} & A_{13} \\ A_{21} & A_{22} & A_{23} \\ A_{31} & A_{32} & A_{33} \end{bmatrix} = \frac{1}{1 - S_3^2} \begin{bmatrix} S_1 (S_x - a_{13} S_3) - S_2 F_x & S_2 (S_x - a_{13} S_3) + S_1 F_x & a_{13} (1 - S_3^2) \\ S_1 (S_y - a_{23} S_3) - S_2 F_y & S_2 (S_y - a_{23} S_3) + S_1 F_y & a_{23} (1 - S_3^2) \\ S_1 (S_z - a_{33} S_3) - S_2 F_z & S_2 (S_z - a_{33} S_3) + S_1 F_z & a_{33} (1 - S_3^2) \end{bmatrix} \quad (A17)$$

This directly indicates:

$$A_{11} = \frac{1}{1 - S_3^2} (S_1 (S_x - a_{13} S_3) - S_2 (S_z a_{23} - S_y a_{33})) \quad (A18)$$

$$A_{12} = \frac{1}{1 - S_3^2} (S_2 (S_x - a_{13} S_3) + S_1 (S_z a_{23} - S_y a_{33})) \quad (A19)$$

$$A_{13} = a_{13} \quad (A20)$$

$$A_{21} = \frac{1}{1 - S_3^2} (S_1 (S_y - a_{23} S_3) - S_2 (S_x a_{33} - S_z a_{13})) \quad (A21)$$

$$A_{22} = \frac{1}{1 - S_3^2} (S_2 (S_y - a_{23} S_3) + S_1 (S_x a_{33} - S_z a_{13})) \quad (A22)$$

$$A_{23} = a_{23} \quad (A23)$$

$$A_{31} = \frac{1}{1 - S_3^2} (S_1 (S_z - a_{33} S_3) - S_2 (S_y a_{13} - S_x a_{23})) \quad (A24)$$

$$A_{32} = \frac{1}{1 - S_3^2} (S_2 (S_z - a_{33} S_3) + S_1 (S_y a_{13} - S_x a_{23})) \quad (A25)$$

$$A_{33} = a_{33} \quad (A26)$$

### Euler Angle Transformations

Any attitude errors of the body axes relative to the desired orientation (that is, any deviation of the x, y, z, axes from the X, Y, Z axes) can be represented by Euler angle rotations: yaw angle  $\psi$  around the Z axis, roll angle  $\phi$  about the displaced X axis, and pitch angle  $\theta$  about the doubly displaced Y axis (see Figure 8). The notation to be used for the matrix which performs each of these three Euler rotations is

$$R_z(\psi) = \begin{bmatrix} \cos \psi & \sin \psi & 0 \\ -\sin \psi & \cos \psi & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$R_x(\phi) = \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \phi & \sin \phi \\ 0 & -\sin \phi & \cos \phi \end{bmatrix}$$

$$R_y(\theta) = \begin{bmatrix} \cos \theta & 0 & -\sin \theta \\ 0 & 1 & 0 \\ \sin \theta & 0 & \cos \theta \end{bmatrix}$$

Given any vector  $V$  with components  $V_1, V_2, V_3$ , in the Orbital axes system, its components  $v_x, v_y, v_z$  in the body axes system are

$$\begin{bmatrix} v_x \\ v_y \\ v_z \end{bmatrix} = R_y(\theta) R_x(\phi) R_z(\psi) \begin{bmatrix} V_1 \\ V_2 \\ V_3 \end{bmatrix}$$

$$= \begin{bmatrix} \cos \theta & 0 & -\sin \theta \\ 0 & 1 & 0 \\ \sin \theta & 0 & \cos \theta \end{bmatrix} \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cos \phi & \sin \phi \\ 0 & -\sin \phi & \cos \phi \end{bmatrix} \begin{bmatrix} \cos \psi & \sin \psi & 0 \\ -\sin \psi & \cos \psi & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} V_1 \\ V_2 \\ V_3 \end{bmatrix}$$

$$= \begin{bmatrix} \cos \theta \cos \psi - \sin \theta \sin \phi \sin \psi & \cos \theta \sin \psi + \sin \theta \sin \phi \cos \psi & -\sin \theta \cos \phi \\ -\cos \phi \sin \psi & \cos \phi \cos \psi & \sin \phi \\ \sin \theta \cos \psi - \cos \theta \sin \phi \sin \psi & \sin \theta \sin \psi - \cos \theta \sin \phi \cos \psi & \cos \theta \cos \phi \end{bmatrix} \begin{bmatrix} V_1 \\ V_2 \\ V_3 \end{bmatrix}$$

Comparing this equation with matrix equation (A3) gives

$$[A] = \begin{bmatrix} \cos \theta \cos \psi - \sin \theta \sin \phi \sin \psi & \cos \theta \sin \psi + \sin \theta \sin \phi \cos \psi & -\sin \theta \cos \phi \\ -\cos \phi \sin \psi & \cos \phi \cos \psi & \sin \phi \\ \sin \theta \cos \psi - \cos \theta \sin \phi \sin \psi & \sin \theta \sin \psi - \cos \theta \sin \phi \cos \psi & \cos \theta \cos \phi \end{bmatrix}$$

The roll angle  $\phi$ , pitch angle  $\theta$ , and yaw angle  $\psi$  can thus be derived:

$$a_{13} = -\sin \theta \cos \phi$$

$$a_{23} = \sin \phi$$

$$a_{33} = \cos \theta \cos \phi$$

and

$$A_{21} = -\sin \psi \cos \phi$$

Thus

$$\phi = \sin^{-1} (a_{23}^i)$$

$$\theta = \tan^{-1} \left( -\frac{a_{13}}{a_{33}} \right)$$

If the angle of  $\phi$  is limited to  $-90 \leq \phi \leq 90$ , then  $\cos \phi$  is always positive and

$$\sin \psi = -\frac{A_{21}}{\cos \phi} = -\frac{A_{21}}{\sqrt{1 - A_{23}^2}}$$

Therefore

$$\psi = \sin^{-1} -\frac{A_{21}}{\sqrt{1 - A_{23}^2}}$$

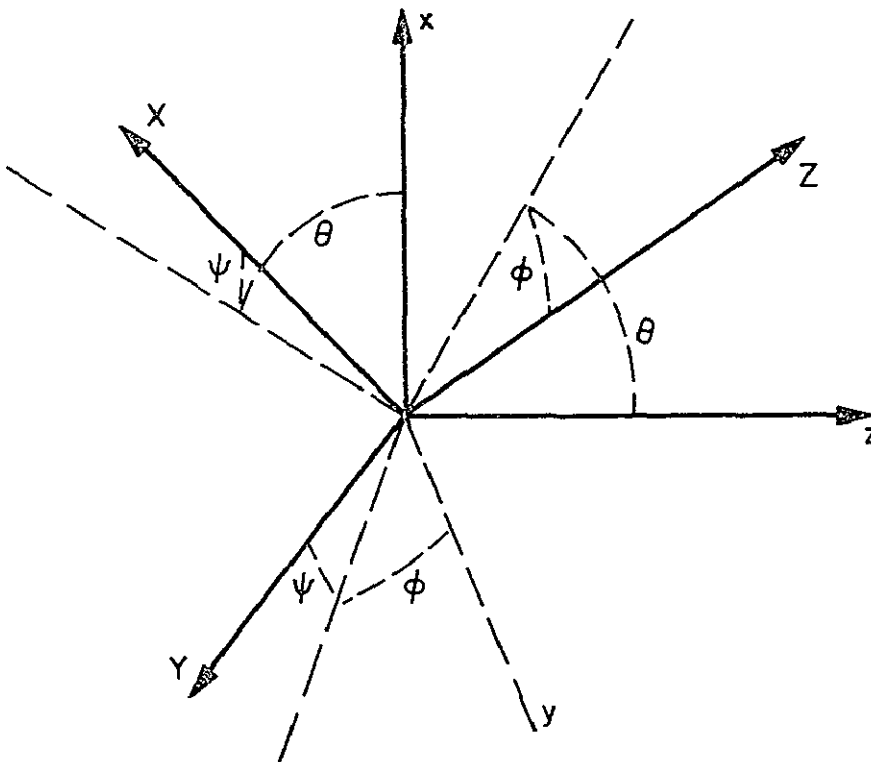


Figure 8. Attitude Error Representation Roll  $\phi$  Pitch  $\theta$ , Yaw  $\psi$



### Presentation of Special Cases

Several cases are presented in this report to illustrate the conditions in which the earth sensor or the sun sensor or both fails to provide the sufficient information. If any of these cases does occur, it is convenient to use a graphical representation of the curves related  $a_{13}$ ,  $a_{23}$  and  $a_{33}$  on a unity sphere which is defined as a sphere in  $a_{13}$ ,  $a_{23}$ ,  $a_{33}$  space of unit radius and fixed to the spacecraft axis system with center at the origin. Since

$$a_{13}^2 + a_{23}^2 + a_{33}^2 = 1$$

or

$$a_{33}^2 = 1 - a_{13}^2 - a_{23}^2 \quad (D1)$$

the local vertical certainly lies on the surface of this sphere at any moment and provides a convenient means for visualizing the motion of the local vertical with respect to body axes. It is useful to indicate the geometric region for displaying the local vertical on such a unity sphere and project it on the  $a_{13}$ ,  $a_{23}$  plane for positive and negative  $a_{33}$ .

#### (A) Case 1

First, a worst case is considered if the spacecraft is in the earth's shadow and if the earth sensor doesn't cut the earth. Evidently, the local vertical and the sun-line cannot be determined. Since the earth sensor does not see the earth, it can be concluded that the local vertical must lie within one of two possible regions in which body axes vary as the gimbal angle  $\gamma$  changes. If the equation (E8)

and the equation (D1) are combined to eliminate  $a_{33}$ , after some algebraic rearrangement, the result gives:

$$\left( \frac{a_{23} - C_{23} \cos \gamma}{\sqrt{1 - C_{23}^2} \sin \gamma} \right)^2 + \left( \frac{a_{13}}{\sqrt{1 - C_{23}^2}} \right)^2 = 1 \quad (D2)$$

for

$$\gamma \neq 0$$

Equation (D2) represents an ellipse within the  $a_{13}$  versus  $a_{23}$  plane unity circles with center at  $E(C_{23} \cos \gamma, 0)$ , with major axis  $1/2 \sqrt{1 - C_{23}^2}$  and minor axis  $1/2 \sqrt{1 - C_{23}^2} \sin \gamma$ . There are two possible sets of curves:

(A<sub>a</sub>) When the scan cone does not cut earth,  $\rho = 0$  and  $C_{23} < \cos(\alpha_e + \sigma)$

then

$$\left( \frac{a_{23} - C_{23} \cos \gamma}{\sin(\alpha_e + \sigma) \sin \gamma} \right)^2 + \left( \frac{a_{13}}{\sin(\alpha_e + \sigma)} \right)^2 < 1 \quad (D3)$$

(A<sub>b</sub>) When the entire scan cone intersects the earth, i.e.,  $\rho = 180$  and

$C_{23} > \cos(\alpha_e - \sigma)$ , then:

$$\left( \frac{a_{23} - C_{23} \cos \gamma}{\sin(\alpha_e - \sigma) \sin \gamma} \right)^2 + \left( \frac{a_{13}}{\sin(\alpha_e - \sigma)} \right)^2 > 1 \quad (D4)$$

The shaded areas in Figure 9 and 10 represent respectively the required projections of region on the  $a_{23} - a_{13}$  plane indicated by equation (D3) and equation (D4).

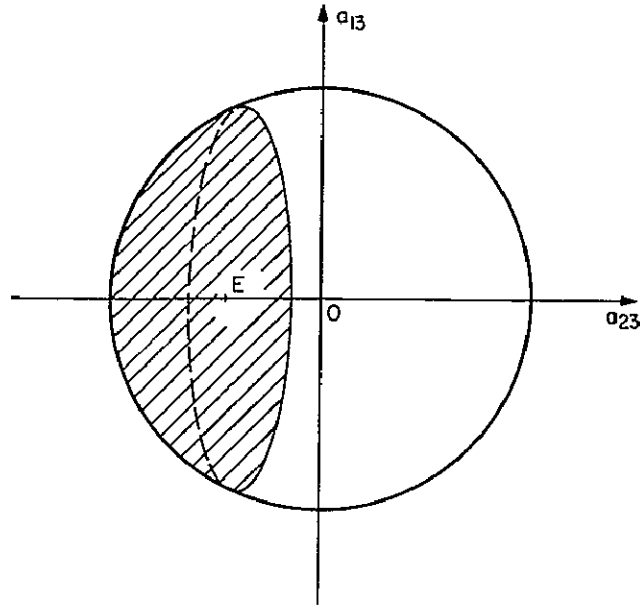


Figure 9. Projection of Error Region in  $a_{23} - a_{13}$  plane when  $\rho = 0, \gamma \neq 0$

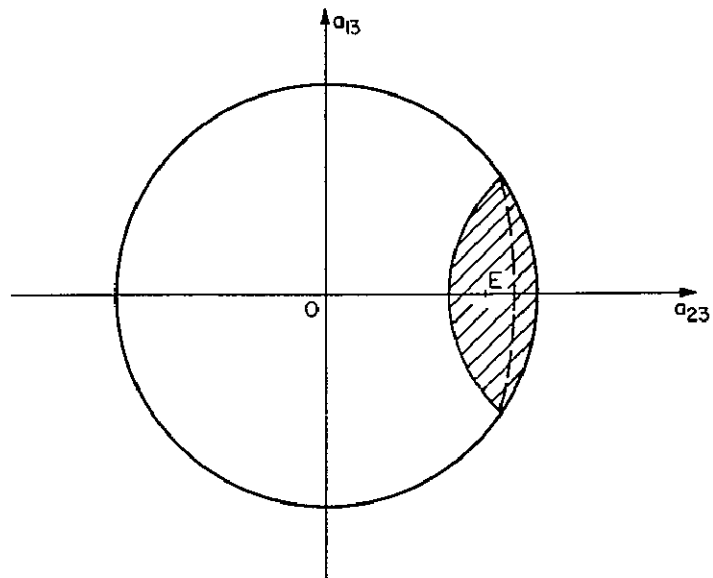


Figure 10. Projection of Error Region in  $a_{23} - a_{13}$  plane when  $\rho = 180, \gamma \neq 0$

If  $\gamma = 0$ , equation (E8) becomes:

$$a_{23} = c_{23} \quad (D5)$$

Therefore, the ellipse represented by equation (D2) degenerates into a straight line, the location of the line depends on the value of  $\rho = 0$ , or  $\rho = 180$ , when

$$\rho = 0$$

then

$$a_{23} < \cos(\alpha_e + \sigma) \quad (D6)$$

when

$$\rho = 180$$

then

$$a_{23} > \cos(\alpha_e - \sigma) \quad (D7)$$

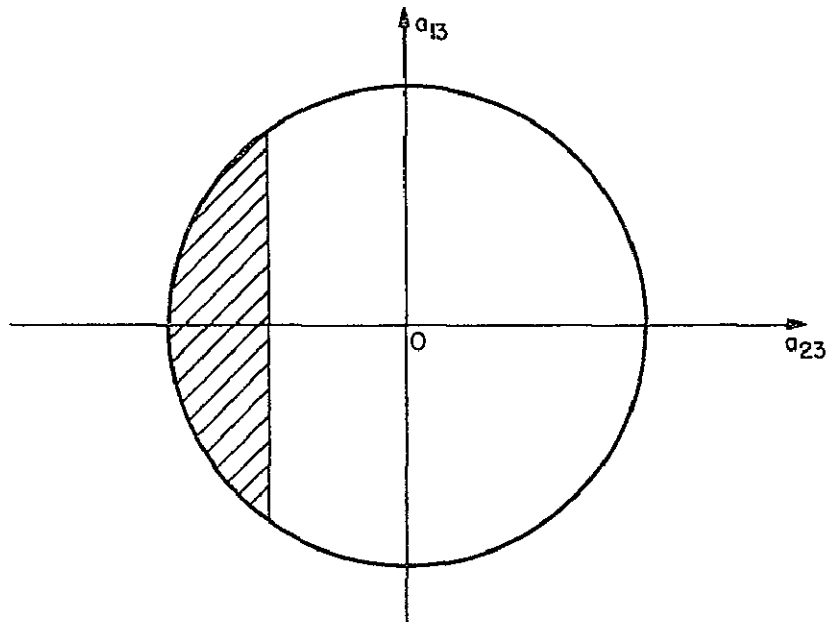


Figure 11 Error Region in  $a_{23} - a_{13}$  plane when  $\rho = 0, \gamma = 0$

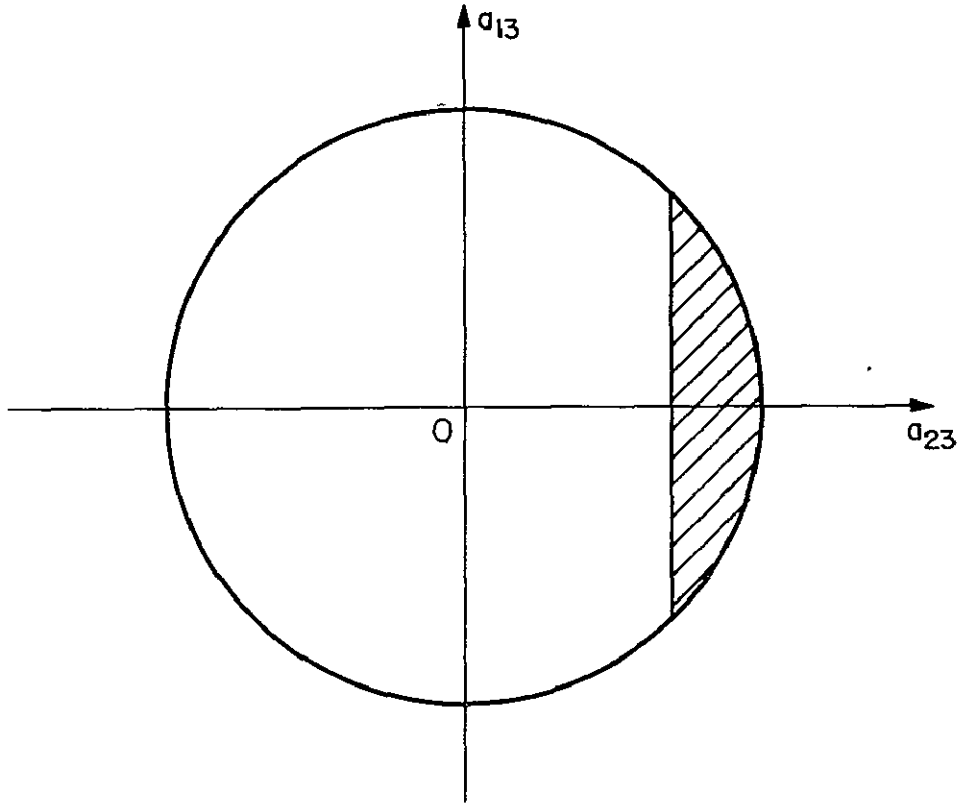


Figure 12. Error Region in  $a_{23} - a_{13}$  plane when  $\rho=180, \gamma=0$

The above two cases are illustrated in Figure 11 and Figure 12, respectively.

(B) Case 2.

Second, consider the case when the sun sensors do receive the sunlight, but the half earth pulse provided by the earth sensor shows its value of 0 or 180 degrees. In addition to the knowledge that the local vertical lies within specified regions as noted previously, it is also known that it must lie somewhere on a specified curve. From the information given by the sun sensor, equation (D1) and equation (E16) are useful for specifying this curve. Rewrite equation (E16) gives:

$$S_z a_{33} = S_3 - S_x a_{13} - S_y a_{23} \quad (D7)$$

If  $S_z = 0$ , equation (D7) represents a straight line in  $a_{23} - a_{13}$  plane.

If  $S_z \neq 0$ , combining equation (D1) and equation (D7) to remove the  $a_{33}$  term, yields:

$$(S_x^2 + S_y^2) a_{13}^2 + 2 S_x S_y a_{13} a_{23} + (S_y^2 + S_z^2) a_{23}^2 - 2 S_3 (S_x a_{13} + S_y a_{23}) = S_z^2 - S_3^2 \quad (D8)$$

This is the general form of an ellipse with the center at  $C (S_3 S_y, S_3 S_x)$ . If the origin is translated to the center of ellipse, the two first degree terms of equation (D8) can be removed. Then rotate the axes through an angle  $\lambda$  ( $\lambda \leq 90$ ) where

$$\tan 2\lambda = \frac{2 S_x S_y}{S_y^2 - S_x^2}$$

Equation (D8) is thus transformed into a standard equation of an ellipse having the foci on the new  $a_{13}$  axis with semi-axes  $\sqrt{1 - S_3^2}$ , and  $S_z \sqrt{1 - S_3^2}$ .

When either  $S_x$  or  $S_y$  is zero, it is obvious that no rotation of the axes is necessary.

When both  $S_x$  and  $S_y$  are zero, Equation (D8) represents a circle of radius  $\sqrt{1 - S_3^2}$ , a point, if  $S_3 = \pm 1$ .

Within the  $a_{13} - a_{23}$  plane, the projection of the curves defined by equation (D7) and (D1) to represent the direction cosines of the body axes with respect to the local vertical also must satisfy the condition mentioned in section (A) Case 1.

Hence the portion of the projection presented graphically on the  $a_{23} - a_{13}$  plane within the region described in the previous section (A) according to  $\rho = 0$  or  $\rho = 180$  is all that can be obtained to meet the required conditions.

It is helpful to generate another view of the sphere and plot the locus of the orbital Y axis ( $A_{12}, A_{22}, A_{32}$ ). Obviously, it is a curve which lies on the surface of the  $A_{12}$  ( $i = 1, 2, 3$ ) direction unity sphere since

$$A_{12}^2 + A_{22}^2 + A_{32}^2 = 1 \quad (D9)$$

The projection of this curve on the  $A_{32} - A_{12}$  plane is given for both  $A_{22} > 0$  and  $A_{22} < 0$ .

Using equation (T3) to form the scalar product  $\vec{S} \cdot \vec{Y}$ , yields:

$$\vec{S} \cdot \vec{Y} = \sin \beta \quad (D10)$$

For convenience, the above equation is therefore written as:

$$S_x A_{12} + S_y A_{22} + S_z A_{32} = \sin \beta \quad (D11)$$

If  $S_y = 0$ , then equation (D11) defines a straight line on the  $A_{32} - A_{12}$  plane assuming  $S_y \neq 0$  and solving the equation (D11) for  $A_{22}$ , gives:

$$A_{22} = \frac{\sin \beta - S_x A_{12} - S_z A_{32}}{S_y} \quad (D12)$$

From equation (D9) and equation (D12) we have

$$\begin{aligned} (S_y^2 + S_z^2) A_{32}^2 + 2 S_x S_z A_{32} A_{12} + (S_x^2 + S_y^2) A_{12}^2 - 2 S_z \sin \beta A_{32} \\ - 2 S_x \sin \beta A_{12} = S_y^2 - \sin^2 \beta \end{aligned} \quad (D13)$$

It can be seen easily that the above equation represents a circle of  $S_y = 1$ . However, if  $S_y$  is neither equal to 0 nor equal to 1, (D13) is the general equation of an ellipse of which the axes of symmetry are:

$$a_{32} = S_z \sin \beta$$

$$a_{12} = S_x \sin \beta$$

and the center of symmetry is  $(S_z \sin \beta, S_x \sin \beta)$ . Like equation (D8) by a proper translation of the axes to remove the  $a_{32}$  term and the  $a_{12}$  term and then by a rotation of the axes with a chosen angle  $\lambda'$  for the removal of the  $a_{32}$   $a_{12}$  term where

$$\tan 2 \lambda' = \frac{2 S_x S_z}{S_z^2 - S_x^2}$$

The equation (D13) is reduced to a standard form with the foci on the new  $a_{12}$  axis and with semi axes  $\cos \beta$  and  $S_y \cos \beta$ .

From the above discussion, it is believed that the graphical techniques are useful to solve the algebraic system of relationships. This approach certainly limits the region from which the error of the attitudes of Delta PAC can be pictured. Such graphical representation can be plotted easily if a computer is used. A FORTRAN program subroutine plot is written for this purpose (see Appendix).



(C) Case 3

Third, for the case when the earth sensor does intersect the earth but the spacecraft enters into the earth's shadow, the informations from the earth sensor will indicate one of the following possibilities

(A)  $0 < \rho < 180^\circ$  and e is not saturated

This case appears as if there were no sun sensors aboard in the spacecraft at that moment. The given informations can only indicate the possible locations of the local vertical which are represented by four sets of equations as shown in equations (E15). No selection can be made from these four sets of equations to represent the true local vertical because the information about the sun line is lacking. However if the locations of the local vertical determined at previous sequences of time during the same orbit are examined, it is always possible to make the proper selection.

(B)  $0 < \rho < 180^\circ$  and e is saturated

This one is similar to case 1 discussed previously. As before the information given by the earth sensor can only be used to obtain equation (D2) for  $\gamma \neq 0$  and equation (D5) for  $\gamma = 0$ . The only difference is that the variable  $C_{23}$  in this case has a definite value instead of having limited range. Therefore if the curve which relates  $a_{13}$ ,  $a_{23}$  and  $a_{33}$  on the surface of the unit sphere is projected on the  $a_{23} - a_{13}$  plane, the projection expresses exactly the relationship between  $a_{23}$  and  $a_{13}$ . Clearly the projection is either an ellipse represented by equation (D2) for  $\gamma \neq 0$  or a straight line given by equation (D5) for  $\gamma = 0$ .

## Error Propagation

It is a known fact that measurements are always accompanied by errors. However small errors are more frequent than large ones since the goal of the sensor design is to keep errors as small as possible. Therefore it is reasonable to assume that the measured quantities which indirectly determine the PAC attitude contain errors but small ones. Because of these errors, the predicted attitudes might not be the true values but the difference between them is expected to be small also.

For Delta PAC there are five quantities being measured by using the sun sensors and the earth sensor, namely: the angles  $\alpha_1$  and  $\beta_1$  which determine the direction cosines of the sun line with respect to the  $i$  th sensor system, the gimbal angle  $\gamma$ , the half earth pulse  $\rho$ , and the telemetered pitch error  $e$ . The corresponding errors are defined by the symbols  $\Delta\alpha_1, \Delta\beta_1, \Delta\gamma, \Delta\rho$  and  $\Delta e$  which are assumed to have the standard deviation  $\sigma_{\alpha_1}, \sigma_{\beta_1}, \sigma_\gamma, \sigma_\rho$  and  $\sigma_e$  respectively. All measurement errors due to unknown cause are of random type, and are independent to each other. They are assumed to follow the laws of probability as given by the normal distribution.

It should be noted that the two angles  $\alpha$  and  $\beta$  defined in the sun sensor section are not measured quantities. At any time their values will be determined by orbit calculation and are assumed to contain no errors.

In this report the methods of probability are applied to analyze the effect of measurement errors. It would be of interest to present the results in terms of error ellipses for a given attitude, the error ellipses are ellipses of equal probability which are similarly placed and are centered at the predicted position (h, k). They are determined by the formula

$$\frac{1}{2(1-r^2)} \left\{ \left( \frac{u-h}{\sigma_u} \right)^2 - 2r \left( \frac{u-h}{\sigma_u} \right) \left( \frac{w-k}{\sigma_w} \right) + \left( \frac{w-k}{\sigma_w} \right)^2 \right\} = L = \text{const} \quad (\text{R1})$$

for  $0 < r < 1$  where  $\sigma_u$ ,  $\sigma_w$  and  $r$  are respectively standard deviation of two variables  $u$  and  $w$  and their correlation coefficients. If  $u$  and  $w$  are not correlated variables then the value of  $r$  is equal to 0 and equation (R1) becomes

$$\left( \frac{u-h}{\sigma_u} \right)^2 + \left( \frac{w-k}{\sigma_w} \right)^2 = 2L \quad (\text{R2})$$

For each ellipse there is a definite probability that a true position ( $u, w$ ) will lie within the ellipse. The probability for ( $u, w$ ) to belong to the ellipse  $L$  given by either equation (R1) or (R2) is

$$P = 1 - e^{-L} \quad (\text{R3})$$

and with  $L = 1/2, 4/2,$  and  $9/2$  the probabilities are 0.4, 0.8, and 0.99 respectively. Of course, formula (R3) can be inverted to find the value of  $L$  corresponding to a given value of  $P$ .

It is essential that the determination of error ellipses for two variables  $u$  and  $w$  depend upon their predicted values of  $h$  and  $k$  and their standard deviations of  $\sigma_u$  and  $\sigma_w$  if  $r = 0$ . Therefore in considering the error ellipses of  $a_{23}$  and  $a_{13}$  and also of  $a_{32}$  and  $a_{12}$  for delta PAC, the former can be defined either at  $0 < \rho < 180^\circ$  and  $|e| < e_s$  without sun sensor information, or at  $0 < \rho < 180^\circ$  and  $|e| = e_s$  in conjunction with the sun's direction; the later can only be presented when the local vertical and the sun line are known but not parallel. The calculation of prediction errors for PAC attitude is thus limited to the case when  $0 < \rho < 180$ .

If the attitude control system of Delta PAC works well, the values of variables  $a_{23}$  and  $a_{13}$  are expected to be much less than 1. That is, the projection of the true local vertical in  $a_{23} - a_{13}$  plane should stay well within  $a_{23} - a_{13}$  unity circle. For this reason, the correlation coefficient of  $a_{23}$  and  $a_{13}$  can be assumed to have the value of 0, and equation (R2) is used to obtain error ellipse for  $a_{23}$  and  $a_{13}$ .

Similarly, equation (R2) can be applied to obtain the error ellipse for  $a_{32}$  and  $a_{12}$ . For Delta PAC each of these ellipses are centered at the prediction point with the probability of 0.99 that a true point falls within it. The lengths of the semi-axes are equal to  $3(\sigma a_{i3})$ ,  $i = 2, 1$  for the  $a_{23} - a_{13}$  ellipse and  $3(\sigma a_{j2})$ , ( $j = 3, 1$ ) for the  $a_{32} - a_{12}$  ellipse where  $\sigma a_{13}$ ,  $\sigma a_{23}$ ,  $\sigma a_{32}$ , and  $\sigma a_{12}$  are the standard deviations of  $a_{13}$ ,  $a_{23}$ ,  $a_{32}$  and  $a_{12}$ , respectively.

If equations (S9) are differentiated with respect to  $\alpha_1$  and  $\beta_1$  the following equations are obtained:

$$\left. \begin{aligned}
\frac{\partial s_{11}}{\partial \alpha_1} &= -s_{11}^2 s_{31} \sec^2 \alpha_1 \\
\frac{\partial s_{11}}{\partial \beta_1} &= -s_{11}^2 s_{21} \sec^2 \beta_1 \\
\frac{\partial s_{21}}{\partial \alpha_1} &= -s_{11} s_{21} s_{31} \sec^2 \alpha_1 \\
\frac{\partial s_{21}}{\partial \beta_1} &= (1 - s_{21}^2) s_{11} \sec^2 \beta_1 \\
\frac{\partial s_{31}}{\partial \alpha_1} &= (1 - s_{31}^2) s_{11} \sec^2 \alpha_1 \\
\frac{\partial s_{31}}{\partial \beta_1} &= -s_{11} s_{21} s_{31} \sec^2 \beta_1
\end{aligned} \right\} \quad (R4)$$

By differentiating equations (S10) with respect to  $s_{11}$ ,  $s_{21}$  and  $s_{31}$  respectively, yield

$$\left. \begin{aligned}
\frac{\partial S_x}{\partial s_{11}} &= -\cos \eta_1 \sin \xi_1 \\
\frac{\partial S_x}{\partial s_{21}} &= \sin \eta_1 \sin \xi_1 \\
\frac{\partial S_x}{\partial s_{31}} &= -\cos \xi_1 \\
\frac{\partial S_y}{\partial s_{11}} &= \cos \eta_1 \cos \xi_1 \\
\frac{\partial S_y}{\partial s_{21}} &= -\sin \eta_1 \cos \xi_1 \\
\frac{\partial S_y}{\partial s_{31}} &= -\sin \xi_1 \\
\frac{\partial S_z}{\partial s_{11}} &= -\sin \eta_1 \\
\frac{\partial S_z}{\partial s_{21}} &= -\cos \eta_1 \\
\frac{\partial S_z}{\partial s_{31}} &= 0
\end{aligned} \right\} \quad (R5)$$

Since  $S_x$ ,  $S_y$  and  $S_z$  are functions of variables  $s_{11}$ ,  $s_{21}$  and  $s_{31}$  where each variable is a function of the independent variables  $\alpha_1$  and  $\beta_1$  therefore:

$$\left. \begin{aligned} \frac{\partial S_x}{\partial \alpha_1} &= \frac{\partial S_x}{\partial s_{11}} \frac{\partial s_{11}}{\partial \alpha_1} + \frac{\partial S_x}{\partial s_{21}} \frac{\partial s_{21}}{\partial \alpha_1} + \frac{\partial S_x}{\partial s_{31}} \frac{\partial s_{31}}{\partial \alpha_1} \\ &= - \{S_x s_{31} + \cos \xi_1\} s_{11} \sec^2 \alpha_1 \\ \frac{\partial S_x}{\partial \beta_1} &= \frac{\partial S_x}{\partial s_{11}} \frac{\partial s_{11}}{\partial \beta_1} + \frac{\partial S_x}{\partial s_{21}} \frac{\partial s_{21}}{\partial \beta_1} + \frac{\partial S_x}{\partial s_{31}} \frac{\partial s_{31}}{\partial \beta_1} \\ &= - \{S_x s_{21} - \sin \eta_1 \sin \xi_1\} s_{11} \sec^2 \beta_1 \end{aligned} \right\} \quad (R6)$$

A similar method is applied to calculate the derivatives of

$$\frac{\partial S_y}{\partial \alpha_1}, \frac{\partial S_y}{\partial \beta_1}, \frac{\partial S_z}{\partial \alpha_1} \text{ and } \frac{\partial S_z}{\partial \beta_1}$$

with the aid of equations (R4) and equations of (R5) the results are:

$$\frac{\partial S_y}{\partial \alpha_1} = - \{S_y s_{31} + \sin \xi_1\} s_{11} \sec^2 \alpha_1 \quad (R7)$$

$$\frac{\partial S_y}{\partial \beta_1} = - \{S_y s_{21} + \sin \eta_1 \cos \xi_1\} s_{11} \sec^2 \beta_1 \quad (R8)$$

$$\frac{\partial S_z}{\partial \alpha_1} = - S_z s_{11} s_{31} \sec^2 \alpha_1 \quad (R9)$$

$$\frac{\partial S_z}{\partial \beta_1} = - \{S_z s_{21} + \cos \eta_1\} s_{11} \sec^2 \beta_1 \quad (R10)$$

Consider the equation (E6) where  $C_{23}$  is a function of  $\rho$  only. The error of  $C_{23}$  can be expressed in terms of measurement error of half earth pulse as

$$\Delta C_{23} = \frac{d C_{23}}{d \rho} \Delta \rho \quad (R11)$$

The standard deviation of  $C_{23}$  is given by

$$\sigma C_{23} = \left| \frac{d C_{23}}{d \rho} \right| \sigma \rho \quad (\text{R12})$$

where  $\sigma \rho$  is the standard deviation of  $\rho$  and

$$\frac{d C_{23}}{d \rho} = - \frac{(1 - C_{23}^2) \sin \sigma \sin \rho}{C_{23} \sin \sigma \cos \rho - \cos \sigma \sqrt{1 - C_{23}^2}} \quad (\text{R13})$$

For Delta PAC,  $\sigma = 45^\circ$ , thus, it can be shown that  $C_{23} \cos \rho$  is less than  $\sqrt{1 - C_{23}^2}$  whenever  $0 < \rho < 180$ . The differentiation of equation (E6) with respect to  $\rho$  in order to obtain equation (R12) is always possible.

As mentioned before, there are two different cases which satisfy the condition  $0 < \rho < 180$ , but require separate approaches. They are:

$$(A) \quad 0 < \rho < 180^\circ \text{ and } |e| < e_s$$

In this case, the local vertical is determined by the earth sensor equation and is not affected by the measurement errors from the sun sensors. Therefore,

$$\left. \begin{aligned} \frac{\partial a_{13}}{\partial \alpha_1} = \frac{\partial a_{13}}{\partial \beta_1} = 0 \\ \frac{\partial a_{23}}{\partial \alpha_1} = \frac{\partial a_{23}}{\partial \beta_1} = 0 \\ \frac{\partial a_{23}}{\partial \alpha_1} = \frac{\partial a_{33}}{\partial \beta_1} = 0 \end{aligned} \right\} \quad (\text{R14})$$

Evidently, the components of the local vertical depend only on the values of  $\gamma$ ,  $\rho$ , and  $e$  as shown in equation (E15). The standard deviation of  $a_{13}$ ,  $a_{23}$ , and  $a_{33}$  can be expressed in terms of  $\sigma_\gamma$ ,  $\sigma_\rho$  and  $\sigma_e$  as follows

$$\left. \begin{aligned} \sigma_{a_{13}}^2 &= \left( \frac{\partial a_{13}}{\partial \gamma} \right)^2 \sigma_\gamma^2 + \left( \frac{\partial a_{13}}{\partial \rho} \right)^2 \sigma_\rho^2 + \left( \frac{\partial a_{13}}{\partial e} \right)^2 \sigma_e^2 \\ \sigma_{a_{23}}^2 &= \left( \frac{\partial a_{23}}{\partial \gamma} \right)^2 \sigma_\gamma^2 + \left( \frac{\partial a_{23}}{\partial \rho} \right)^2 \sigma_\rho^2 + \left( \frac{\partial a_{23}}{\partial e} \right)^2 \sigma_e^2 \\ \sigma_{a_{33}}^2 &= \left( \frac{\partial a_{33}}{\partial \gamma} \right)^2 \sigma_\gamma^2 + \left( \frac{\partial a_{33}}{\partial \rho} \right)^2 \sigma_\rho^2 + \left( \frac{\partial a_{33}}{\partial e} \right)^2 \sigma_e^2 \end{aligned} \right\} \quad (R15)$$

As noted at the end of the earth sensor equation section, there are four sets of equations, (see equation E13), representing the position of the local vertical. Only one set of these equations give the correct components of the local vertical and it is selected. The technique applied for the selection was mentioned before.

The partial derivatives

$$\frac{\partial a_{13}}{\partial \gamma}, \frac{\partial a_{13}}{\partial \rho}, \text{ and } \frac{\partial a_{13}}{\partial e}, \quad (i = 1, 2, 3)$$

which appear in this case should be calculated from the equations of that selected set.

Since the components of the sun line are independent from the errors of the earth sensor, this implies



$$\left. \begin{aligned}
 \frac{\partial S_x}{\partial \gamma} = \frac{\partial S_x}{\partial \rho} = \frac{\partial S_x}{\partial e} = 0 \\
 \frac{\partial S_y}{\partial \gamma} = \frac{\partial S_y}{\partial \rho} = \frac{\partial S_y}{\partial e} = 0 \\
 \frac{\partial S_z}{\partial \gamma} = \frac{\partial S_z}{\partial \rho} = \frac{\partial S_z}{\partial e} = 0
 \end{aligned} \right\} \quad (R16)$$

Moreover, since  $S_1$ ,  $S_2$ , and  $S_3$  depend only on the value of  $\alpha$  and  $\beta$  which are assumed to contain no errors. It is obvious

$$\left. \begin{aligned}
 \frac{\partial S_1}{\partial \alpha_1} = \frac{\partial S_2}{\partial \alpha_1} = \frac{\partial S_3}{\partial \alpha_1} = 0 \\
 \frac{\partial S_1}{\partial \beta_1} = \frac{\partial S_2}{\partial \beta_1} = \frac{\partial S_3}{\partial \beta_1} = 0 \\
 \frac{\partial S_1}{\partial \gamma} = \frac{\partial S_2}{\partial \gamma} = \frac{\partial S_3}{\partial \gamma} = 0 \\
 \frac{\partial S_1}{\partial \rho} = \frac{\partial S_2}{\partial \rho} = \frac{\partial S_3}{\partial \rho} = 0 \\
 \frac{\partial S_1}{\partial e} = \frac{\partial S_2}{\partial e} = \frac{\partial S_3}{\partial e} = 0
 \end{aligned} \right\} \quad (R17)$$

By the use of above equations the partial derivatives of  $A_{12}$  and  $A_{32}$  with respect to  $\alpha_1, \beta_1, \gamma, \rho$  and  $e$  are:

$$\left. \begin{aligned}
\frac{\partial A_{12}}{\partial \alpha_1} &= \frac{1}{1 - S_3^2} \left\{ S_2 \left( \frac{\partial S_x}{\partial \alpha_1} \right) + S_1 \left( a_{23} \frac{\partial S_z}{\partial \alpha_1} - a_{33} \frac{\partial S_y}{\partial \alpha_1} \right) \right\} \\
\frac{\partial A_{12}}{\partial \beta_1} &= \frac{1}{1 - S_3^2} \left\{ S_2 \left( \frac{\partial S_x}{\partial \beta_1} \right) + S_1 \left( a_{23} \frac{\partial S_z}{\partial \beta_1} - a_{33} \frac{\partial S_y}{\partial \beta_1} \right) \right\} \\
\frac{\partial A_{12}}{\partial \gamma} &= \frac{1}{1 - S_3^2} \left\{ - (S_2) (S_3) \left( \frac{\partial a_{13}}{\partial \gamma} \right) + S_1 \left( S_z \frac{\partial a_{23}}{\partial \gamma} - S_y \frac{\partial a_{33}}{\partial \gamma} \right) \right\} \\
\frac{\partial A_{12}}{\partial \rho} &= \frac{1}{1 - S_3^2} \left\{ - (S_2) (S_3) \left( \frac{\partial a_{13}}{\partial \rho} \right) + S_1 \left( S_z \frac{\partial a_{23}}{\partial \rho} - S_y \frac{\partial a_{33}}{\partial \rho} \right) \right\} \\
\frac{\partial A_{12}}{\partial e} &= \frac{1}{1 - S_3^2} \left\{ - (S_2) (S_3) \left( \frac{\partial a_{13}}{\partial e} \right) + S_1 \left( S_z \frac{\partial a_{23}}{\partial e} - S_y \frac{\partial a_{33}}{\partial e} \right) \right\}
\end{aligned} \right\} \quad (R18)$$

and

$$\begin{aligned}
\frac{\partial A_{32}}{\partial \alpha_1} &= \frac{1}{1 - S_3^2} \left\{ S_2 \left( \frac{\partial S_z}{\partial \alpha_1} \right) + S_1 \left( a_{13} \frac{\partial S_y}{\partial \alpha_1} - a_{23} \frac{\partial S_x}{\partial \alpha_1} \right) \right\} \\
\frac{\partial A_{32}}{\partial \beta_1} &= \frac{1}{1 - S_3^2} \left\{ S_2 \left( \frac{\partial S_z}{\partial \beta_1} \right) + S_1 \left( a_{13} \frac{\partial S_y}{\partial \beta_1} - a_{23} \frac{\partial S_x}{\partial \beta_1} \right) \right\} \\
\frac{\partial A_{32}}{\partial \gamma} &= \frac{1}{1 - S_3^2} \left\{ - (S_2) \left( S_3 \frac{\partial a_{33}}{\partial \gamma} \right) + S_1 \left( S_y \frac{\partial a_{13}}{\partial \gamma} - S_x \frac{\partial a_{23}}{\partial \gamma} \right) \right\} \\
\frac{\partial A_{32}}{\partial \rho} &= \frac{1}{1 - S_3^2} \left\{ - (S_2) \left( S_3 \frac{\partial a_{33}}{\partial \rho} \right) + S_1 \left( S_y \frac{\partial a_{13}}{\partial \rho} - S_x \frac{\partial a_{23}}{\partial \rho} \right) \right\} \\
\frac{\partial A_{32}}{\partial e} &= \frac{1}{1 - S_3^2} \left\{ - (S_2) \left( S_3 \frac{\partial a_{33}}{\partial e} \right) + S_1 \left( S_y \frac{\partial a_{13}}{\partial e} - S_x \frac{\partial a_{23}}{\partial e} \right) \right\}
\end{aligned} \quad (R19)$$

In terms of known partial derivatives the standard deviation of  $A_{12}$  and  $A_{32}$  can be obtained from the following

$$\begin{aligned}
\sigma_{A_{12}}^2 &= \left( \frac{\partial A_{12}}{\partial \alpha_1} \right)^2 \sigma_{\alpha_1}^2 + \left( \frac{\partial A_{12}}{\partial \beta_1} \right)^2 \sigma_{\beta_1}^2 + \left( \frac{\partial A_{12}}{\partial \gamma} \right)^2 \sigma_{\gamma}^2 + \left( \frac{\partial A_{12}}{\partial \rho} \right)^2 \sigma_{\rho}^2 + \left( \frac{\partial A_{12}}{\partial e} \right)^2 \sigma_e^2 \\
\sigma_{A_{32}}^2 &= \left( \frac{\partial A_{12}}{\partial \beta_1} \right)^2 \sigma_{\alpha_1}^2 + \left( \frac{\partial A_{32}}{\partial \beta_1} \right)^2 \sigma_{\beta_1}^2 + \left( \frac{\partial A_{32}}{\partial \gamma} \right)^2 \sigma_{\gamma}^2 + \left( \frac{\partial A_{32}}{\partial \rho} \right)^2 \sigma_{\rho}^2 + \left( \frac{\partial A_{32}}{\partial e} \right)^2 \sigma_e^2
\end{aligned} \quad (R20)$$

$$(B) \quad \begin{matrix} \text{if} \\ \text{if} \end{matrix} \quad 0 < \rho < 180 \quad \text{and} \quad |e| = e_s$$

As stated in the earth sensor equation section the components of the local vertical are determined by employing the following three equations for case in which  $0 < \rho < 180$  and  $|e| = e_s$

$$\left. \begin{aligned} S_x a_{13} + S_y a_{23} + S_z a_{33} &= -\cos \alpha \cos \beta \\ \cos \gamma a_{23} + \sin \gamma a_{33} &= C_{23} \\ a_{13}^2 + a_{23}^2 + a_{33}^2 &= 1 \end{aligned} \right\} \quad (R21)$$

In view of the above equations, it is apparent that the variables  $a_{13}$ ,  $a_{23}$ , and  $a_{33}$  are now dependent on the values of  $S_x$ ,  $S_y$ , and  $S_z$  as well as  $\gamma$  and  $\rho$ .

Differentiating the equations (R21) with respect to  $S_x$ ,  $S_y$ ,  $S_z$ ,  $\gamma$ , and  $\rho$  respectively and making use of the equations (R16) in addition to

$$\left. \begin{aligned} \frac{\partial \gamma}{\partial S_x} = \frac{\partial \gamma}{\partial S_y} = \frac{\partial \gamma}{\partial S_z} &= 0 \\ \frac{\partial C_{23}}{\partial S_x} = \frac{\partial C_{23}}{\partial S_y} = \frac{\partial C_{23}}{\partial S_z} &= 0 \end{aligned} \right\} \quad (R22)$$

give the following five sets of equations

$$\left. \begin{aligned} S_x \frac{\partial a_{13}}{\partial S_x} + S_y \frac{\partial a_{23}}{\partial S_x} + S_z \frac{\partial a_{33}}{\partial S_x} &= -a_{13} \\ \cos \gamma \frac{\partial a_{23}}{\partial S_x} + \sin \gamma \frac{\partial a_{33}}{\partial S_x} &= 0 \\ a_{13} \frac{\partial a_{13}}{\partial S_x} + a_{23} \frac{\partial a_{23}}{\partial S_x} + a_{33} \frac{\partial a_{33}}{\partial S_x} &= 0 \end{aligned} \right\} \quad (R23)$$

$$\left. \begin{aligned}
 S_x \frac{\partial a_{13}}{\partial S_y} + S_y \frac{\partial a_{23}}{\partial S_y} + S_z \frac{\partial a_{33}}{\partial S_y} &= -a_{23} \\
 \cos \gamma \frac{\partial a_{23}}{\partial S_y} + \sin \gamma \frac{\partial a_{33}}{\partial S_y} &= 0 \\
 a_{13} \frac{\partial a_{13}}{\partial S_y} + a_{23} \frac{\partial a_{23}}{\partial S_y} + a_{33} \frac{\partial a_{33}}{\partial S_y} &= 0
 \end{aligned} \right\} \quad (R24)$$

$$\left. \begin{aligned}
 S_x \frac{\partial a_{13}}{\partial S_z} + S_y \frac{\partial a_{23}}{\partial S_z} + S_z \frac{\partial a_{33}}{\partial S_z} &= -a_{33} \\
 \cos \gamma \frac{\partial a_{23}}{\partial S_z} + \sin \gamma \frac{\partial a_{33}}{\partial S_z} &= 0 \\
 a_{13} \frac{\partial a_{13}}{\partial S_z} + a_{23} \frac{\partial a_{23}}{\partial S_z} + a_{33} \frac{\partial a_{33}}{\partial S_z} &= 0
 \end{aligned} \right\} \quad (R25)$$

$$\left. \begin{aligned}
 S_x \frac{\partial a_{13}}{\partial \gamma} + S_y \frac{\partial a_{23}}{\partial \gamma} + S_z \frac{\partial a_{33}}{\partial \gamma} &= 0 \\
 \cos \gamma \frac{\partial a_{23}}{\partial \gamma} + \sin \gamma \frac{\partial a_{33}}{\partial \gamma} &= a_{23} \sin \gamma - a_{33} \cos \gamma \\
 a_{13} \frac{\partial a_{13}}{\partial \gamma} + a_{23} \frac{\partial a_{23}}{\partial \gamma} + a_{33} \frac{\partial a_{33}}{\partial \gamma} &= 0
 \end{aligned} \right\} \quad (R26)$$

and

$$\left. \begin{aligned}
 S_x \frac{\partial a_{13}}{\partial \rho} + S_y \frac{\partial a_{23}}{\partial \rho} + S_z \frac{\partial a_{33}}{\partial \rho} &= 0 \\
 \cos \gamma \frac{\partial a_{23}}{\partial \rho} + \sin \gamma \frac{\partial a_{33}}{\partial \rho} &= \frac{d C_{23}}{d \rho} \\
 a_{13} \frac{\partial a_{13}}{\partial \rho} + a_{23} \frac{\partial a_{23}}{\partial \rho} + a_{33} \frac{\partial a_{33}}{\partial \rho} &= 0
 \end{aligned} \right\} \quad (R27)$$

where  $dC_{23}/d\rho$  is given in equation (R13).

If

$$\Delta = \begin{vmatrix} S_x & S_y & S_z \\ 0 & \cos \gamma & \sin \gamma \\ a_{13} & a_{13} & a_{33} \end{vmatrix} \neq 0 \quad (\text{R28})$$

then each set of equations (R23), (R24), (R25), (R26), and (R27) can be solved to obtain the derivatives

$$\frac{\partial a_{13}}{\partial x}, \frac{\partial a_{13}}{\partial y}, \frac{\partial a_{13}}{\partial z}, \frac{\partial a_{13}}{\partial \gamma} \quad \text{and} \quad \frac{\partial a_{13}}{\partial \rho}, \quad (i = 1, 2, 3.)$$

Let

$$P = a_{23} \sin \gamma - a_{33} \cos \gamma$$

Then

$$\left. \begin{aligned} \frac{\partial a_{13}}{\partial S_x} &= \frac{a_{13} P}{\Delta} \\ \frac{\partial a_{23}}{\partial S_x} &= \frac{-a_{13}^2 \sin \gamma}{\Delta} \\ \frac{\partial a_{33}}{\partial S_x} &= \frac{a_{13}^2 \cos \gamma}{\Delta} \end{aligned} \right\} \quad (\text{R29})$$

$$\left. \begin{aligned} \frac{\partial a_{13}}{\partial S_y} &= \frac{a_{23} P}{\Delta} \\ \frac{\partial a_{23}}{\partial S_y} &= \frac{-a_{13} a_{23} \sin \gamma}{\Delta} \\ \frac{\partial a_{33}}{\partial S_y} &= \frac{a_{13} a_{23} \cos \gamma}{\Delta} \end{aligned} \right\} \quad (\text{R30})$$

$$\left. \begin{aligned} \frac{\partial a_{13}}{\partial S_z} &= \frac{a_{33} P}{\Delta} \\ \frac{\partial a_{23}}{\partial S_z} &= \frac{-a_{13} a_{33} \sin \gamma}{\Delta} \\ \frac{\partial a_{33}}{\partial S_z} &= \frac{a_{13} a_{33} \cos \gamma}{\Delta} \end{aligned} \right\} \quad (R31)$$

$$\left. \begin{aligned} \frac{\partial a_{13}}{\partial \gamma} &= (a_{23} S_z - a_{33} S_y) P/\Delta \\ \frac{\partial a_{23}}{\partial \gamma} &= (S_x a_{33} - S_z a_{13}) P/\Delta \\ \frac{\partial a_{33}}{\partial \gamma} &= (S_y a_{13} - S_x a_{23}) P/\Delta \end{aligned} \right\} \quad (R32)$$

and

$$\left. \begin{aligned} \frac{\partial a_{13}}{\partial \rho} &= (S_z a_{23} - S_y a_{33}) \frac{d C_{23}}{d \rho} / \Delta \\ \frac{\partial a_{23}}{\partial \rho} &= (S_x a_{33} - S_z a_{13}) \frac{d C_{23}}{d \rho} / \Delta \\ \frac{\partial a_{33}}{\partial \rho} &= (S_y a_{13} - S_x a_{23}) \frac{d C_{23}}{d \rho} / \Delta \end{aligned} \right\} \quad (R33)$$

With the use of the equations (R29) to (R33) and the equations (R6) to (R10), the following partial derivatives can be evaluated.

$$\begin{aligned} \frac{\partial a_{13}}{\partial \alpha_1} &= \frac{\partial a_{13}}{\partial S_x} \frac{\partial S_x}{\partial \alpha_1} + \frac{\partial a_{13}}{\partial S_y} \frac{\partial S_y}{\partial \alpha_1} + \frac{\partial a_{13}}{\partial S_z} \frac{\partial S_z}{\partial \alpha_1} \\ &= - (S_3 s_{31} + a_{13} \cos \xi_1 + a_{23} \sin \xi_1) (s_{11} \sec^2 \alpha_1) P/\Delta \end{aligned} \quad (R34)$$

$$\frac{\partial a_{13}}{\partial \beta_1} = \frac{\partial a_{13}}{\partial S_x} \frac{\partial S_x}{\partial \beta_1} + \frac{\partial a_{13}}{\partial S_y} \frac{\partial S_y}{\partial \beta_1} + \frac{\partial a_{13}}{\partial S_z} \frac{\partial S_z}{\partial \beta_1} \quad (\text{R35})$$

$$= (-S_3 s_{21} + a_{13} \sin \eta_1 \sin \xi_1 - a_{23} \sin \eta_1 \cos \xi_1 - a_{33} \cos \eta_1) (s_{21} \sec^2 \beta_1 P) / \Delta$$

$$\frac{\partial a_{23}}{\partial \alpha_1} = \frac{\partial a_{23}}{\partial S_x} \frac{\partial S_x}{\partial \alpha_1} + \frac{\partial a_{23}}{\partial S_y} \frac{\partial S_y}{\partial \alpha_1} + \frac{\partial a_{23}}{\partial S_z} \frac{\partial S_z}{\partial \alpha_1} \quad (\text{R36})$$

$$= (S_3 s_{31} + a_{13} \cos \xi_1 + a_{23} \sin \xi_1) (s_{11} \sec^2 \alpha_1) a_{13} \sin \gamma / \Delta$$

$$\frac{\partial a_{23}}{\partial \beta_1} = \frac{\partial a_{23}}{\partial S_x} \frac{\partial S_x}{\partial \beta_1} + \frac{\partial a_{23}}{\partial S_y} \frac{\partial S_y}{\partial \beta_1} + \frac{\partial a_{23}}{\partial S_z} \frac{\partial S_z}{\partial \beta_1} \quad (\text{R37})$$

$$= (S_3 s_{21} - a_{13} \sin \eta_1 \sin \xi_1 + a_{23} \sin \eta_1 \cos \xi_1 + a_{33} \cos \eta_1) (s_{11} \sec^2 \beta_1 a_{13} \sin \gamma) / \Delta$$

Hence

$$\left. \begin{aligned} \sigma_{a_{13}}^2 &= \left( \frac{\partial a_{13}}{\partial \alpha_1} \right)^2 \sigma_{\alpha_1}^2 + \left( \frac{\partial a_{13}}{\partial \beta_1} \right)^2 \sigma_{\beta_1}^2 + \left( \frac{\partial a_{13}}{\partial \gamma} \right)^2 \sigma_{\gamma}^2 + \left( \frac{\partial a_{13}}{\partial \rho} \right)^2 \sigma_{\rho}^2 \\ \sigma_{a_{23}}^2 &= \left( \frac{\partial a_{23}}{\partial \alpha_1} \right)^2 \sigma_{\alpha_1}^2 + \left( \frac{\partial a_{23}}{\partial \beta_1} \right)^2 \sigma_{\beta_1}^2 + \left( \frac{\partial a_{23}}{\partial \gamma} \right)^2 \sigma_{\gamma}^2 + \left( \frac{\partial a_{23}}{\partial \rho} \right)^2 \sigma_{\rho}^2 \end{aligned} \right\} \quad (\text{R38})$$

Also the partial derivatives of  $A_{12}$  and  $A_{32}$  with respect to  $\alpha_1$ ,  $\beta_1$ ,  $\gamma$ , and  $\rho$  can be derived from the equations (A19) and (A25) and then are simplified as follows

$$\begin{aligned}
\frac{\partial A_{12}}{\partial \alpha_1} &= \frac{1}{1-S_3^2} \left( S_2 \left( \frac{\partial S_x}{\partial \alpha_1} - S_3 \frac{\partial a_{13}}{\partial \alpha_1} \right) + S_1 \left( a_{23} \frac{\partial S_z}{\partial \alpha_1} + S_z \frac{\partial a_{23}}{\partial \alpha_1} - a_{33} \frac{\partial S_y}{\partial \alpha_1} - S_y \frac{\partial a_{33}}{\partial \alpha_1} \right) \right) \\
\frac{\partial A_{12}}{\partial \beta_1} &= \frac{1}{1-S_3^2} \left( S_2 \left( \frac{\partial S_x}{\partial \beta_1} - S_3 \frac{\partial a_{13}}{\partial \beta_1} \right) + S_1 \left( a_{23} \frac{\partial S_z}{\partial \beta_1} + S_z \frac{\partial a_{23}}{\partial \beta_1} - a_{33} \frac{\partial S_y}{\partial \beta_1} - S_y \frac{\partial a_{33}}{\partial \beta_1} \right) \right) \\
\frac{\partial A_{12}}{\partial \gamma} &= \frac{1}{1-S_3^2} \left( -S_2 S_3 \frac{\partial a_{13}}{\partial \gamma} + S_1 \left( S_z \frac{\partial a_{23}}{\partial \gamma} - S_y \frac{\partial a_{33}}{\partial \gamma} \right) \right) \\
\frac{\partial A_{12}}{\partial \rho} &= \frac{1}{1-S_3^2} \left( -S_2 S_3 \frac{\partial a_{13}}{\partial \rho} + S_1 \left( S_z \frac{\partial a_{23}}{\partial \rho} - S_y \frac{\partial a_{33}}{\partial \rho} \right) \right) \\
\text{and} & \\
\frac{\partial A_{32}}{\partial \alpha_1} &= \frac{1}{1-S_3^2} \left( S_2 \left( \frac{\partial S_z}{\partial \alpha_1} - S_3 \frac{\partial a_{33}}{\partial \alpha_1} \right) + S_1 \left( a_{13} \frac{\partial S_y}{\partial \alpha_1} + S_y \frac{\partial a_{13}}{\partial \alpha_1} - a_{23} \frac{\partial S_x}{\partial \alpha_1} - S_x \frac{\partial a_{23}}{\partial \alpha_1} \right) \right) \\
\frac{\partial A_{32}}{\partial \beta_1} &= \frac{1}{1-S_3^2} \left( S_2 \frac{\partial S_z}{\partial \beta_1} - S_3 \frac{\partial a_{33}}{\partial \beta_1} + S_1 \left( a_{13} \frac{\partial S_y}{\partial \beta_1} + S_y \frac{\partial a_{13}}{\partial \beta_1} - a_{23} \frac{\partial S_x}{\partial \beta_1} - S_x \frac{\partial a_{23}}{\partial \beta_1} \right) \right) \\
\frac{\partial A_{32}}{\partial \gamma} &= \frac{1}{1-S_3^2} \left( -S_2 S_3 \frac{\partial a_{33}}{\partial \gamma} + S_1 \left( S_y \frac{\partial a_{13}}{\partial \gamma} - S_x \frac{\partial a_{23}}{\partial \gamma} \right) \right) \\
\frac{\partial A_{32}}{\partial \rho} &= \frac{1}{1-S_3^2} \left( -S_2 S_3 \frac{\partial a_{33}}{\partial \rho} + S_1 \left( S_y \frac{\partial a_{13}}{\partial \rho} - S_x \frac{\partial a_{23}}{\partial \rho} \right) \right)
\end{aligned} \tag{R39}$$

In terms of the above partial derivatives, the standard deviation of  $a_{12}$  and  $a_{32}$  can thus be calculated. They are

$$\begin{aligned}
\sigma_{A_{12}}^2 &= \left( \frac{\partial A_{12}}{\partial \alpha_1} \right)^2 \sigma_{\alpha_1}^2 + \left( \frac{\partial A_{12}}{\partial \beta_1} \right)^2 \sigma_{\beta_1}^2 + \left( \frac{\partial A_{12}}{\partial \gamma} \right)^2 \sigma_{\gamma}^2 + \left( \frac{\partial A_{12}}{\partial \rho} \right)^2 \sigma_{\rho}^2 \\
\sigma_{A_{32}}^2 &= \left( \frac{\partial A_{32}}{\partial \alpha_1} \right)^2 \sigma_{\alpha_1}^2 + \left( \frac{\partial A_{32}}{\partial \beta_1} \right)^2 \sigma_{\beta_1}^2 + \left( \frac{\partial A_{32}}{\partial \gamma} \right)^2 \sigma_{\gamma}^2 + \left( \frac{\partial A_{32}}{\partial \rho} \right)^2 \sigma_{\rho}^2
\end{aligned} \tag{R40}$$



## Conclusion

This report contains the formulation and the derivation of the attitude determination equations, the calculation of the required elements of error ellipses, and the discussion of various possible special cases for the spacecraft Delta PAC. Based on the analysis developed in this report, A FORTRAN IV program was written to compute the spacecraft attitude with a subroutine to output the plots of the error ellipses and to indicate the region of possible attitude solution for incomplete sensor data. The program was written for the SDS-9300 digital computer and was utilized with actual flight data obtained from the Delta PAC which was launched on Aug. 9, 1969. In the course of examining the flight attitude data, the program was exercised for all modes of data processing including no sun information, with sun information, etc. The results yielded by the use of this analysis have determined the attitude behavior of Delta PAC and evaluated the spacecraft attitude control system. - The program is essential to prove that the flight performance of Delta PAC has successfully met its mission objectives.

## REFERENCES

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2. Shapiro, I. I., "The Prediction of Ballistic Missile Trajectories from Radar Observations," New York: McGraw-Hill Book Company, Inc., April 1958
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## APPENDIX A

### INSTRUCTION FOR DIGITAL PROGRAM

The first input card for the digital program contains ten input values in which the first one, ISUN, indicates the sun sensor information. All initial conditions for the earth sensor equations are specified on this card. If ISUN is zero, the spacecraft is in the earth shadow and a run is initiated immediately. If ISUN is not zero, a second input card is needed to read in the values of  $\alpha_1$ ,  $\beta_1$ ,  $\alpha$ , and  $\beta$ . The value of ISUN determines whether one or two input cards is required for each run. The run is terminated when a blank card is read in. A summary of the total runs is then listed in the output.

APPENDIX B

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C*****LISTING OF FORTRAN PROGRAM*****
C ATTITUDE DETERMINATION FOR DELTA PAC
C BY A. C. FANG
C
C DIMENSION ETA(3) ,SETA(3) ,CETA(3) ,SE(3) ,SSE(3) ,CSE(3)
C DIMENSION AA11(2) ,AA21(2) ,AA31(2) ,AA12(2) ,AA22(2) ,AA32(2)
C DIMENSION AA13(4) ,AA23(4) ,AA33(4) ,TH(4) ,TE(4) ,PH(4)
C DIMENSION NT(2) ,DVS(4) ,DAA12(2) ,DAA32(2)
C DIMENSION DAE13(2) ,DAG13(2) ,DAR13(2) ,DAE23(2) ,DAG23(2) ,JAR23(2)
C DIMENSION DAE33(2) ,DAG33(2) ,DAR33(2) ,JAA13(4) ,DAA23(4) ,DAA33(4)
C DIMENSION DAI13(2) ,DAI23(2) ,DAI33(2) ,DBI13(2) ,DBI23(2) ,DBI33(2)
C DIMENSION AW(15)
C DIMENSION IA(90,55) ,IB(90,55)
C
C COMMON C23,IA,IB,R0MD,IDAY,ITIM
C
C *****DEFINE THE VARIABLES
C
C *****
C VAR 1 VAR 2 UNIT
C *****
C ,RADIUS OF EARTH ORBITAL ATTITUDE N.M.
C RE H
C *****
C GIMBAL ANGLE SCAN CONE HALF ANGLE DEGREE
C GAMD SIGMA
C *****
C HALF EARTH PULSE ELECTRONIC SATURATION DEGREE
C R0MD ED
C *****
C TELEMETERED PITCH ERROR GEOMETRICAL PITCH COMPUTED ERROR DEGREE
C ETD THETA
C *****
C ORBITAL POSITION ANGLE EARTH SUBTENDED ANGLE (HALF) DEGREE
C DALFA ALFA
C *****
C ANGLE BETWEEN ORBIT PLANE AND EARTH SUN LINE DEGREE
C DBETA
C *****
C AZIMUTH ELEVATION DEGREE
C ALID BTID
C *****
C
C CONTINUE
C
C *****
C *****
C BODY AXES :
C X Y Z
C *****
C RIGHT HANDED SYSTEM FOR SUN SENSOR HEAD I, I=1,2,3.
C EI FI GI
C *****

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```

C      ANGLE BETWEEN Y-Z PLANE AND PLANE CONTAINING E1 AND Z
C      SE1          SE2          SE3          DEGREE
C*****
C      ANGLE BETWEEN E1 AND X-Y PLANE
C      ETAD1       ETAD2       ETAD3       DEGREE
C*****
C      COSINE OF ANGLE BETWEEN SCANNER CONE AXIS AND LOCAL VERTICAL
C      C23
C*****
C      DIRECTION COSINES OF THE LOCAL VERTICAL W.R.T. THE BODY AXES
C      AA13        AA23        AA33
C*****
C      DIRECTION COSINES OF SUN-LINE W.R.T. THE SUN SENSOR SYSTEM
C      ST          SU          SV
C*****
C      DIRECTION COSINES OF SUN-LINE W.R.T. THE LOCAL VERTICAL SYSTEM
C      S1          S2          S3
C*****
C      DIRECTION COSINES OF SUN-LINE W.R.T. BODY AXES
C      SX          SY          SZ
C*****
C      DIRECTION COSINES OF ORBIT NORMAL W.R.T. THE BODY AXES
C      A12        A22        A32
C*****
C      DIRECTION COSINES OF VELOCITY VECTOR W.R.T. THE BODY AXES
C      A11        A21        A31
C*****
C      ROLL ANGLE      PITCH ANGLE      YAW ANGLE
C      TH            TE            PH
C*****
C      ERROR ESTIMATE IN ETB, GAMAD, ROMD, ALID, BTID ARE
C      DET          DGA          DRM
C      DALI        DBFI
C*****
C*****DERIVATIVES
C*****
C      VAR 1          VAR 2          VAR 3          NOTE
C*****
C      PARTIAL DERIVATIVE OF SX, SY AND SZ W.R.T. ALID
C      DSXA          DSYA          DSZA
C*****
C      PARTIAL DERIVATIVE OF SX, SY AND SZ W.R.T. BTID
C      DSXB          DSYB          DSZB
C*****
C      PARTIAL DERIVATIVE OF A13(1), A23(1), A33(1) W.R.T. SX ARE
C      AX131        AX231        AX331
C*****
C      PARTIAL DERIVATIVE OF A13(2), A23(2), A33(2) W.R.T. SX ARE
C      AX132        AX232        AX332
C*****
C      PARTIAL DERIVATIVE OF A13(1), A23(1), A33(1) W.R.T. SY ARE
C      AY131        AY231        AY331

```

```

C*****
C PARTIAL DERIVATIVE OF A13(2),A23(2),A33(2) W.R.T. SY ARE
C AY132 AY232 AY332
C*****
C PARTIAL DERIVATIVE OF A13(1),A23(1),A33(1) W.R.T. SZ ARE
C AZ131 AZ231 AZ331
C*****
C PARTIAL DERIVATIVE OF A13(2),A23(2),A33(2) W.R.T. SZ ARE
C AZ132 AZ232 AZ332
C*****
C PARTIAL DERIVATIVE OF A13(1),A23(1),A33(1) W.R.T. AL1D
C DA113(1) DA123(1) DUA133(1)
C*****
C PARTIAL DERIVATIVE OF A13(1),A23(1),A33(1) W.R.T. B11D
C DB113(1) DB123(1) DB133(1)
C*****
C PARTIAL DERIVATIVE OF A12 W.R.T. AL1D,BT1D,ETD,GAMAD,R0MD. ARE:
C DA12 DB12 DE12
C DG12 DR12
C*****
C PARTIAL DERIVATIVE OF A32 W.R.T. AL1C,BT1D,ETD,GAMAD,R0MD. ARE:
C DA32 DB32 DE32
C DG32 DR32
C*****
C PARTIAL DERIVATIVE OF A13(1),A23(1),A33(1) W.R.T. E1D
C DAE13(1) DAE23(1) DAE33(1)
C*****
C PARTIAL DERIVATIVE OF A13(1),A23(1),A33(1) W.R.T. GAMAD
C DAG13(1) DAG23(1) DAG33(1)
C PARTIAL DERIVATIVE OF A13(1),A23(1),A33(1) W.R.T. R0MD
C DAR13(1) DAR23(1) DAR33(1)
C*****
C*****
C*****STORE INPUT DATA AND START THE CALCULATION.
C*****
DATA RE/3448.0/
DATA RAU,PI/0.01745329,3.14159265/
DATA SIGMAD,ED/45.0,45.0/
DATA ETAD1,ETAD2,ETAD3/26.08,26.06,26.25/
DATA SE1,SE2,SE3/0.07,120.06,239.87/
DATA DEI,DGA,DRM,DAL1,DBTI/.017453,.00873,.017453,.017453,.017453/
MRUN=0
TDIF=0.0001
E=ED*RAD
SIGMA=SIGMAD*RAD
SE(1)=SE1*RAD
SE(2)=SE2*RAD
SE(3)=SE3*RAD
ETA(1)=ETAD1*RAD
ETA(2)=ETAD2*RAD
ETA(3)=ETAD3*RAD
SWS = SIN (SIGMA)
CSS=COS(SIGMA)
DO 10 I=1,3

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```

SSE(I)=SIN(SL(I))
CSE(I)=COS(SL(I))
SETA(I)=SIN(LTA(I))
CETA(I)=COS(LTA(I))
1J CONTINUE
C*****
2L READ 3J, ISUN, GAMAU, ETD, ROMU, IDAY, IHR, MIN, ISEC, IO, H
C*****
3L FORMAT(12, F12.2, F14.0, F14.0, I4, I4, I4, I4, I6, F0.3)
IF (IO.EQ.0) GO TO 99U
IF (MRUN.EQ.0) GO TO 4U
IF (ISR.NE.13) GO TO 99U
4C ITIM=IHR*1000+MIN*100+ISEC
PRINT 6U
PRINT 5U, IDAY, IHR, MIN, ISEC, IO, H
5U FORMAT(1X, $ DAY = $, I4, $, HR=$, I4, $, MIN=$, I4, $, SEC=$, I4,
1 $, SR=I=$, I4, 1X, $, r = $, F10.4, /)
AW(1)=IDAY
AW(2)=IHR
AW(3)=MIN
AW(4)=ISEC
AW(5)=IO
AW(6)=GAMAU
AW(7)=ROMU
AW(8)=J.J
AW(9)=L.L
AW(10)=U.U
AW(11)=U.U
AW(12)=U.U
AW(13)=U.U
ALFAE = ASIN (RE/(RE+H))
ALFAL=ALFAE/RAD
SNA=SIN(ALFAE)
CSA = COS (ALFAE)
CAPS=COS(ALFAE+SIGMA)
CAMS=COS(ALFAE-SIGMA)
PRINT 6U
6C FORMAT(1X, $*****$ /)
PRINT 7U
7L FORMAT(1X, $EARTH SENSOR INFORMATION GIVES:$ /)
PRINT 8J, GAMAU, ETD, ROMU
8J FORMAT(1X, $ GAMA (IN DEGREE)=$, F10.6, $, EI (IN DEGREE)=$,
IF10.6, $, ROM (IN DEGREE)=$, F11.0, /)
GAMA=GAMAU*RAD
ET=ETD*RAD
ROM=ROMU*RAD
CSG = COS (GAMA)
SNG=SIN(GAMA)
SNR=SIN(ROM)
CSR=COS(ROM)
TNG=SNG/CSG
DO 9J I=1,2
DAE13(I)=0.0
IAE23(I)=0.0

```

```

DAE33(I)=0.0
DAG13(I)=0.0
DAG23(I)=0.0
DAG33(I)=0.0
DAR13(I)=0.0
DAR23(I)=0.0
DAR33(I)=0.0
DAI13(I)=0.0
DAI23(I)=0.0
DAI33(I)=0.0
DBI13(I)=0.0
DBI23(I)=0.0
DBI33(I)=0.0
DAA13(I)=0.0
DAA23(I)=0.0
DAA12(I)=0.0
DAA32(I)=0.0
95 CONTINUE
CT1=C.0
CT2=J.0
FAL=0.0
FAL=0.0
FBT=0.0
FBT=0.0
LE=0
C*****
C*PART 1.....APPLICATION OF EARTH SENSOR
C*****
AEI=ABSF(EI)
IF (R0MD.EQ.180.0) GO TO 160
IF (R0MD.EQ.0.0) GO TO 170
100 CB=CSS*LSA
CJ=SNS*SNS*SNR*SNR
CA=1.0-CD
CC=SNS+CSR*SQRTF(SNA*SNA-CD)
C23=(CB-CC)/CA
RRC23=1-C23*C23
SRC23=SQRTF(RRC23)
DNC23=-RRC23*SNS*SNR*DRM
DDC23=CSR*SNS*C23-CSS*SRC23
LE=0
IF (DDC23 .NE. 0.0) GO TO 110
LE=1
GO TO 120
110 DC23=DNC23/DDC23
120 ES=AMIN (R0M,PI-R0M,L)
IF (AET.L(.ES) GO TO 130
C*****WHEN ABSF(COS(R0M))<1 BUT AEI=ES PUT LP=0
LP=0
MLP=0
GO TO 180
C*****WHEN ABSF(COS(R0M))<1 AND AEI<ES PUT LP=3
130 LP=3
MLP=3

```



```

C*****CALCULATION OF A13,A23,A33 WHEN (ET) IS NOT EQUAL TO 0
      THETAS=ET
C      THETAS=GEOMETRICAL PITCH COMPUTED ERROR(-180<THETAS<180)
      SNT=SIN(THETAS)
      CNT=COS(THETAS)
      AA13(1)=SNT*SRC23
      AA13(2)=-AA13(1)
      AA23(1)=C23*CSG+SRC23*SNG*CNT
      AA23(2)=C23*CSG-SRC23*SNG*CNT
      AA33(1)=C23*SNG+SRC23*CSG*CNT
      AA33(2)=C23*SNG-SRC23*CSG*CNT
      AA13(3)=AA13(1)
      AA13(4)=AA13(2)
      AA23(3)=AA23(2)
      AA23(4)=AA23(1)
      AA33(3)=AA33(2)
      AA33(4)=AA33(1)
      DO 140 I=1,4
      TH(I)=ASIN(AA23(I))/RAD
      TE(I)=ATAN(-AA13(I),AA33(I))/RAD
14L CONTINUE
      IF(LE .EQ. 1) GO TO 180
C*****CALCULATE STANDARD DEVIATION OF A13,A23,A33
      DAE13(1)=SRC23*CNT*DET
      DAE13(2)=-DAE13(1)
      DAG13(1)=0.0
      DAG13(2)=0.0
      DAR13(1)=-C23*SNT*DC23/SRC23
      DAR13(2)=-DAR13(1)
      DAE23(1)=-SRC23*SNG*SNT*DET
      DAE23(2)=-DAE23(1)
      DAG23(1)=- (AA33(1))*DGA
      DAG23(2)=- (AA33(2))*DGA
      DAR23(1)=(CSG-C23*SNG*CNT/SRC23)*DC23
      DAR23(2)=(CSG+C23*SNG*CNT/SRC23)*DC23
      DAE33(1)=SRC23*CSG*SNT*DET
      DAE33(2)=-DAE33(1)
      DAG33(1)=(AA23(1))*DGA
      DAG33(2)=(AA23(2))*DGA
      DAR33(1)=(SNG+C23*CSG*CNT/SRC23)*DC23
      DAR33(2)=(SNG-C23*CSG*CNT/SRC23)*DC23
      LE 150 I=1,2
      DAA13(I)=SQRT(DAE13(I)**2+DAR13(I)**2+DAG13(I)**2)
      DAA23(I)=SQRT(DAE23(I)**2+DAR23(I)**2+DAG23(I)**2)
      DAA33(I)=SQRT(DAE33(I)**2+DAR33(I)**2+DAG33(I)**2)
      DAI13(I)=0.0
      DBI13(I)=0.0
      DAI23(I)=0.0
      DBI23(I)=0.0
      DAI33(I)=0.0
      DBI33(I)=0.0
15J CONTINUE
      DAA23(3)=DAA23(2)
      DAA23(4)=DAA23(1)

```

```

DAA33(3)=DAA33(2)
DAA33(4)=DAA33(1)
GO TO 180
C*****
C***** *** WHEN COS(R0M)=-1, PUT LP=2 TO PLOT THE REGION
160 LP=2
MLP=2
C23=CAMS
GO TO 180
C***** *** WHEN COS(R0M)=1 PUT LP=1 TO PLOT THE REGION
170 LP=1
MLP=1
C23=CAPS
GO TO 180
C*****
C*PART 2.....SUN SENSOR APPLICATION
C*****
180 IF (ISUN.E0.) GO TO 590
C*****
READ 190,ALID,BTID,DALFA,DBETA
C*****
190 FORMAT(6F14.8)
ALI=ALID*RAD
BTI=BTID*RAD
TAI=(SIN(ALI))/COS(ALI)
TBI=(SIN(BTI))/COS(BTI)
TAB=SQRT(1.0+(TAI**2)+(TBI**2))
SI=1.0/TAB
SU=TBI/TAB
SV=TAI/TAB
PRINT 60
PRINT 200,ISUN
200 FORMAT(1X,$ISUN=$,I2,$, SUN SENSOR INFORMATION GIVES:$/)
PRINT 210,ALID,BTID
210 FORMAT(3X,$AZIMUTH=$,F10.5,$ DEG, ELEVATION=$,F10.5,$ DEG,$/)
PRINT 220,SI,SU,SV,DALFA,DBETA
220 FORMAT(3X,$S1I=$,F9.6,$, S2I=$,F9.6,$, S3I=$,F9.0,
1$, ALFA(IN DEGREE)=$,F11.6,$, BETA(IN DEGREE)=$,F11.6,/)
ALFA=DALFA*RAD
BETA=DBETA*RAD
C*****CALCULATE THE DIRECTION COSINES OF SUN LINE W.R.T. BODY AXIS
J=ISUN
SX=-ST*CETA(J)*SSE(J)+SU*SETA(J)*SSE(J)-SV*CSE(J)
SY=ST*CETA(J)*CSE(J)-SU*SETA(J)*CSE(J)-SV*SSE(J)
SZ=-ST*SETA(J)-SU*CETA(J)
PRINT 230,SX,SY,SZ
230 FORMAT(2X,$ THE DIRECTION COSINES OF SUN LINE W.R.T. SPACECRAFT
1AXES ARE $,F12.0,$, $,F12.8,$, $,F12.8,/)
C*****CALCULATE PARTIAL DERIVATIVE OF SX,SY,AND SZ W.R.T.ALID AND BTID
SEAI=1.0+TAI*TAI
SEBI=1.0+TBI*TBI
STSAI=SI*SEAI*DALI
STSBII=ST*SEBI*DBTI
DSXA=(-SX*SV-CSE(J))*STSAI

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DSXB=(-SX*SU+SETA(J)*SSE(J))*S1SBI
DSYA=(-SY*SV-SSE(J))*S1SAI
DSYB=(-SY*SU-SETA(J)*CSE(J))*S1SBI
DSZA=-SZ*SV*S1SAI
DSZB=(-SZ*SU-CEFA(J))*S1SBI
*****CALCULATE THE DIRECTION COSINES OF SUN LINE W.R.T. GEOCENTRIC AXIS
S1=-COS(BETA)*(SIN(ALFA))
S2=SIN(BETA)
S3=-COS(BETA)*(COS(ALFA))
PRINT 24J,S1,S2,S3
PRINT 0J
24J FORMAT(3X,$THE DIRECTION COSINES OF SUN LINE W.R.T.GEOCENTRIC AXIS
1=$,F12.8,$,$,F12.8,$,$,F12.8,/)
IF(LP.EQ.0)GO TO 250
GO TO (610,610,450),LP
*****CALCULATE THE LOCAL VERTICAL WHEN EI=ES.
250 SYSG=SY*SNG
SYCG=SY*CSG
SZSG=SZ*SNG
SZCG=SZ*CSG
FCC=SYSG-SZCG
IF (FCC.EQ.0.0)GO TO 280
FCR=FCC**2
FA=SX*SX+FCR
FB=SX*(-S3+C23*(SZSG+SYCG))
FC=FLR+2.0*C23*S3*(SZSG+SYCG)-S3*S3-C23*C23*(1.0-SX*SX)
ROT=FB*FB+FA*FC
IF(ROT .LT. 0.) GO TO 260
AA13(1)=(-FB+SQRTE(ROT))/FA
AA13(2)=(-FB-SQRTE(ROT))/FA
AA23(1)=(SNG*(S3-SX*AA13(1))-C23*S2)/FCC
AA23(2)=(SNG*(S3-SX*AA13(2))-C23*S2)/FCC
AA33(1)=(CSG*(-S3+SX*AA13(1))+C23*SY)/FCC
AA33(2)=(CSG*(-S3+SX*AA13(2))+C23*SY)/FCC
GO TO 320
260 PRINT 27J
27J FORMAT(3X,$A13,A23,A33 ARE NOT REAL NUMBER. THEREFORE$)
GO TO 420
280 PRINT 29J
290 FORMAT(3X,$SY*SIN(GAMA)-SZ*COS(GAMA)=0$,/)
IF (SX.EQ.0.)GO TO 420
C TRY TO USE OTHER METHOD FOR SEEKING THE SOLUTION OF A13,A23,A33
AA13(1)=(S3*CSG-SY*C23)/SX*CSG
AA13(2)=AA13(1)
ROBT=1.0-AA13(1)*AA13(1)-C23*C23
IF(ROBT.LT.0.0)GO TO 260
SQRR=SQRTE(ROBT)
AA23(1)=C23*CSG+SNG*SQRR
AA23(2)=C23*CSG-SNG*SQRR
AA33(1)=C23*SNG-CSG*SQRR
AA33(2)=C23*SNG+CSG*SQRR
PRINT 30J, AA13(1),AA23(1),AA33(1)
30C FORMAT(3X,$A13(1)=$,F11.8,$, A23(1)=$,F11.8,$, A33(1)=$,F11.8,/)
PRINT 31J,AA13(2),AA23(2),AA33(2)

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310 FORMAT(3X, #A13(2)=#, F11.8, #, A23(2)=#, F11.8, #, A33(2)=#, F11.8, /)
C*****
C TEST IF A13 SATISFY THE CONDITION FOR THE REQUIRED REGION
C TL, TM ARE THE LOWER LIMITS FOR THE TWO POSSIBLE REGIONS
C TU, TN ARE THE UPPER LIMITS FOR THE TWO POSSIBLE REGIONS
320 TL=SRC23*SIN(ES)
TU=SRC23
TM=-TU
TN=-TL
DO 340 I=1,2
CK=AA13(I)
IF((CK.GE.TL.AND.CK.LE.TU).OR.(CK.GE.TM.AND.CK.LE.TN))GO TO 330
NT(I)=0
GO TO 340
330 NT(I)=1
340 CONTINUE
MNT=AMIN (NT(1),NT(2))
MAI=AMAX (NT(1),NT(2))
PRINT 300, AA13(1),AA23(1),AA33(1)
PRINT 310,AA13(2),AA23(2),AA33(2)
PRINT 60
PRINT 350, TL, TU, TM, TN
350 FORMAT(3X, #NOTE : A13 HAS TO SATISFY THE REQUIRED CONDITION FOR
1THE LIMITED REGION, #, F8.5, # < A13 < #, F8.5, #, OR, #, F8.5,
2# < A13 < #, F8.5, /)
IF(MAT.NE.MNT)GO TO 370
PRINT 360
360 FORMAT(3X, #THE TWO SOLUTIONS OF A13 DO NOT SATISFY THE CONDITION
1FOR THE REQUIRED REGION #, /)
GO TO 420
370 TH(1)=ASIN(AA23(1))/RAD
TH(2)=ASIN(AA23(2))/RAD
TE(1)=ATAN (-AA13(1),AA33(1))/RAD
TE(2)=ATAN (-AA13(2),AA33(2))/RAD
PRINT 380, TH(1),TE(1)
380 FORMAT(3X, #ROLL ANGLE TH(1)=#, F10.5, # DEG., PITCH ANGLE TE(1)=#,
1F10.5, # DEG. #, /)
PRINT 390, TH(2),TE(2)
390 FORMAT(3X, #ROLL ANGLE TH(2)=#, F10.5, # DEG., PITCH ANGLE TE(2)=#,
1F10.5, # DEG. #, /)
IF(LE .EQ. 1) GO TO 450
C****CALCULATE PARTIAL DIFFERENTIAL OF A13, A23, A33 w.r.t. SX, SY, SZ, ETC
C****AND CALCULATE STANDARD DEVIATION OF A13, A23, A33
SA131=AA13(1)*AA13(1)
SA132=AA13(2)*AA13(2)
CFR1=SN6*AA23(1)-CS6*AA33(1)
CFR2=SN6*AA23(2)-CS6*AA33(2)
CDN1=SX*(-CFR1)+FCC*AA13(1)
CDN2=SX*(-CFR2)+FCC*AA13(2)
IF(CDN1 .EQ. 0.0) GO TO 400
AX131=CFR1*AA13(1)/CDN1
AX231=-SN6*SA131/CDN1
AX331= CS6*SA131/CDN1
AY131=CFR1*AA23(1)/CDN1

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AY231=-SNG*AA13(1)*AA23(1)/CDN1
AY331= CSG*AA13(1)*AA23(1)/CDN1
AZ131=CFR1*AA33(1)/CDN1
AZ231=-SNG*AA13(1)*AA33(1)/CDN1
AZ331= CSG*AA13(1)*AA33(1)/CDN1
RG131=(SZ*AA23(1)-SY*AA33(1))/CDN1
RG231=(SX*AA33(1)-SZ*AA13(1))/CDN1
DAG13(1)=CFR1*RG131*DGA
DAG23(1)=CFR1*RG231*DGA
DAR13(1)=RG131*DC23
DAR23(1)=RG231*DC23
DAI13(1)= AX131*DSXA+AY131*DSYA+AZ131*DSZA
PAI23(1)= AX231*DSXA+AY231*DSYA+AZ231*DSZA
DAI33(1)= AX331*DSXA+AY331*DSYA+AZ331*DSZA
DBI13(1)= AX131*DSXB+AY131*DSYB+AZ131*DSZB
DBI23(1)= AX231*DSXB+AY231*DSYB+AZ231*DSZB
DBI33(1)= AX331*DSXB+AY331*DSYB+AZ331*DSZB
DAA13(1)=SQRTF(DAI13(1)**2+DBI13(1)**2+DAG13(1)**2+DAR13(1)**2)
DAA23(1)=SQRTF(DAI23(1)**2+DBI23(1)**2+DAG23(1)**2+DAR23(1)**2)
DAE13(1)=0.0
DAE23(1)=0.0
DAE33(1)=0.0
GO TO 410

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C\*\*\*\*\*

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400 LE=2
410 IF(CDN2 .EQ.0.0) GO TO 440
RG132=(SZ*AA23(2)-SY*AA33(2))/CDN2
RG232=(SX*AA33(2)-SZ*AA13(2))/CDN2
DAG13(2)=CFR2*RG132*DGA
DAG23(2)=CFR2*RG232*DGA
DAR13(2)=RG132*DC23
DAR23(2)=RG232*DC23
AX132= CFR2*AA13(2)/CDN2
AX232=-SNG*SA132/CDN2
AX332= CSG*SA132/CDN2
AY132=CFR2*AA23(2)/CDN2
AY232=-SNG*AA13(2)*AA23(2)/CDN2
AY332= CSG*AA13(2)*AA23(2)/CDN2
AZ132=CFR2*AA33(2)/CDN2
AZ232=-SNG*AA13(2)*AA33(2)/CDN2
AZ332= CSG*AA13(2)*AA33(2)/CDN2
DAI13(2)= AX132*DSXA+AY132*DSYA+AZ132*DSZA
PAI23(2)= AX232*DSXA+AZ232*DSYA+AZ232*DSZA
DAI33(2)= AX332*DSXA+AY332*DSYA+AZ332*DSZA
DBI13(2)= AX132*DSXB+AY132*DSYB+AZ132*DSZB
DBI23(2)= AX232*DSXB+AY232*DSYB+AZ232*DSZB
DBI33(2)= AX332*DSXB+AY332*DSYB+AZ332*DSZB
DAA13(2)=SQRTF(DAI13(2)**2+DBI13(2)**2+DAG13(2)**2+DAR13(2)**2)
DAA23(2)=SQRTF(DAI23(2)**2+DBI23(2)**2+DAG23(2)**2+DAR23(2)**2)
DAE13(2)=0.0
DAE23(2)=0.0
DAE33(2)=0.0
PRINT 480,DAA13(1),DAA13(2)
PRINT 490,DAA23(1),DAA23(2)

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PRINT 60
GO TO 450
420 PRINT 430
430 FORMAT(3X, 'NO SOLUTION FOR A13, A23, A33. TRY TO USE PLOT$', /)
MLP=1
GO TO 610
440 LE=LE+3
C*****
C*****CHECKING WHICH ONE OF THE LOCAL VERTICAL SHOULD BE SELECTED
C K=1 OR 2 MEANS ONE ANSWER FOR LOCAL VERTICAL. K=3 MEANS TWO.
450 ID0=2+(LP+1)/2
DO 460 I=1, ID0
CHECK1=(AA13(I))*2+(AA23(I))*2+(AA33(I))*2-1.0
CHECK2=(AA23(I))*CSG+(AA33(I))*SNG-C23
CHECK3=SX*AA13(I)+SY*AA23(I)+SZ*AA33(I)-S3
DVS(I)=AMAX(ABSF(CHECK1), ABSF(CHECK2), ABSF(CHECK3))
460 CONTINUE
PRINT 60
ASV=AMIN (DVS(1), DVS(2))
IF(LP.EQ.J) GO TO 510
ASW=AMIN (DVS(3), DVS(4))
IF(ASV.LE.ASW) GO TO 470
AA23(1)=AA23(3)
AA23(2)=AA23(4)
AA33(1)=AA33(3)
AA33(2)=AA33(4)
TH(1)=TH(3)
TH(2)=TH(4)
TE(1)=TE(3)
TE(2)=TE(4)
ASV=ASW
DVS(1)=DVS(3)
DVS(2)=DVS(4)
470 PRINT 300, AA13(1), AA23(1), AA33(1)
PRINT 310, AA13(2), AA23(2), AA33(2)
PRINT 380, TH(1), TE(1)
PRINT 390, TH(2), TE(2)
PRINT 60
IF(LE.EQ.1) GO TO 510
PRINT 60
PRINT 480, DAA13(1), DAA13(2)
480 FORMAT(3X, 'DEVIATION IN A13, DA13(1)=F8.5, DA13(2)=F8.5)
PRINT 490, DAA23(1), DAA23(2)
490 FORMAT(3X, 'DEVIATION IN A23, DA23(1)=F8.5, DA23(2)=F8.5)
PRINT 500, DAA33(1), DAA33(2)
500 FORMAT(3X, 'DEVIATION IN A33, DA33(1)=F8.5, DA33(2)=F8.5)
PRINT 60
510 DIF=ABSF(DVS(1)-DVS(2))
IF(DIF.LE.TDIF) GO TO 550
IF(ASV.EQ.DVS(1)) GO TO 520
K=2
GO TO 530
520 K=1
530 IF(LP.EQ.3) GO TO 570

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      IF(NT(K).EQ.K) GO TO 570
      AW(8)=TH(1)
      AW(9)=TE(1)
      AW(11)=TH(2)
      AW(12)=TE(2)
      PRINT 54U
40  FORMAT(3X,*,THE POSITION OF A13 DOES NOT SATISFY THE REQUIRED LIMIT
      1REGION*,/)
      GO TO 42J
50  K=3
      IF(LP .EQ. 3) GO TO 66U
      IF(MNT.NE.0)GO TO 61J
      IF(NT(1).EQ.3) GO TO 56U
      K=1
      GO TO 57U
60  K=2
670 A13=AA13(K)
      A23=AA23(K)
      A33=AA33(K)
      VAE13=DAE13(K)
      VAE23=DAE23(K)
      VAE33=DAE33(K)
      VAG13=DAG13(K)
      VAR13=DAR13(K)
      VAG23=DAG23(K)
      VAR23=DAR23(K)
      VAG33=DAG33(K)
      VAR33=DAR33(K)
      VAI13=DAI13(K)
      VAI23=DAI23(K)
      VAI33=DAI33(K)
      VBI13=DBI13(K)
      VBI23=DBI23(K)
      VBI33=DBI33(K)
      VAA13=DAA13(K)
      VAA23=DAA23(K)
      PRINT 60
      PRINT 58U,A13,A23,A33
58U  FORMAT(3X,*,DIRECTION COSINES OF LOCAL VERTICAL ARE:*,F12.8,F12.8,
      1F12.8)
      IF(LP .EQ. 0) GO TO 610
      GO TO 66U
590 PRINT 60U,ISUN
600  FORMAT(1X,*,ISUN=*,I2,*, THE SPACECRAFT IS IN EARTH SHADOW, THE
      1DIRECTION COSINES OF SUN LINE CAN NOT BE DETERMINED*,/)
      IF(LP .EQ.U) GO TO 65U
      GO TO (630,640,86U),LP
610 PRINT 62U
620  FORMAT(1H1)
      WHEN LP IS LESS THAN 3, CT1, CT2, EAL,EBT, , ARE INPUT DATA TO
      SUBROUTINE FOR PRINTING PURPOSE.
      CT1=DALFA
      CT2=DBETA
      EAL=ALID

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EBT=BTID
IF (LP.EQ.0) GO TO 650
GO TO (630,640),LP
630 CALL PLOT(LP,GAMA,ISUN,SX,SY,SZ,S2,S3,CT1,CT2,EAL,EBT)
GO TO 970
640 CALL PLOT(LP,GAMA,ISUN,SX,SY,SZ,S2,S3,CT1,CT2,EAL,EBT)
GO TO 970
650 CALL PLOT(LP,GAMA,ISUN,SX,SY,SZ,S2,S3,CT1,CT2,EAL,EBT)
IF(ISUN .EQ. 0) GO TO 970
IF(MLP .EQ. 1) GO TO 970
C*****
C*PART 3.....DETERMINATION OF BODY AXIS
C*****
660 IF(K .EQ. 3) GO TO 670
I=K
GO TO 680
670 DO 830 I=1,2
A13=AA13(I)
A23=AA23(I)
A33=AA33(I)
VAE13=DAE13(I)
VAE23=DAE23(I)
VAE33=DAE33(I)
VAG13=DAG13(I)
VAG23=DAG23(I)
VAR13=DAR13(I)
VAR23=DAR23(I)
VAG33=DAG33(I)
VAR33=DAR33(I)
VAI13=DAI13(I)
VAI23=DAI23(I)
VAI33=DAI33(I)
VBI13=DBI13(I)
VBI23=DBI23(I)
VBI33=DBI33(I)
VAA13=DAA13(I)
VAA23=DAA23(I)
680 Q=S3
D=Q*Q-1.0
IF(ABSF(D).LT.(.001**2 )) GO TO 840
A11=(S1*(A13*Q-SX)+S2*(SZ*A23-SY*A33))/D
A21=(S1*(A23*Q-SY)+S2*(SX*A33-SZ*A13))/D
A31=(S1*(A33*Q-SZ)+S2*(SY*A13-SX*A23))/D
A12=(S2*(A13*Q-SX)-S1*(SZ*A23-SY*A33))/D
A22=(S2*(A23*Q-SY)-S1*(SX*A33-SZ*A13))/D
A32=(S2*(A33*Q-SZ)-S1*(SY*A13-SX*A23))/D
THI=ASIN(A23)/RAD
THE=ATAN ( -A13,A33)/RAD
IF(ABSF(A23).GE.1.0) GO TO 700
QSA23=SQRTF(1.0-A23*A23)
PHE=ASIN(-A21/QSA23)/RAD
GO TO 720
700 PRINT 710
710 FORMAT(3X,'$YAW ANGLE CAN NOT BE DETERMINED,SET IT TO 50 DEG.$',/)

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PHE=50.0
720 TH(I)=THI
TE(I)=THE
PH(I)=PHE
LU=10+(I-1)*3
AW(LU)=PHE
LV=9+(I-1)*3
AW(LV)=THE
LW=8+(I-1)*3
AW(LW)=THI
AA11(I)=A11
AA21(I)=A21
AA31(I)=A31
AA12(I)=A12
AA22(I)=A22
AA32(I)=A32
AA13(I)=A13
AA23(I)=A23
AA33(I)=A33
IF(LE.EQ.0) GO TO 730
IF((LE.EQ.1).OR.(LE.EQ.5)) GO TO 870
IF(((LE-1).EQ.2).OR.((LE-1).EQ.0)) GO TO 730
GO TO 87J

C*****
C      PLOT THE ERROR ELLIPSE OF A23,A13
730 LP=3
CT1=-A23
CT2=-A13
EAL=3.0*VAA23
EBT=3.0*VAA13
CALL PLOT(LP,GAMA,ISUN,SX,SY,SZ,S2,S3,CT1,CT2,EAL,EBT)

C*****
C*****CALCULATE THE PARTIAL DERIVATIVES OF Q W.R.T.ALID,B11D,GAMA,ETC.
C*****CALCULATE THE STANDARD DEVIATION OF A12,A22,A32
IF(MLP.EQ.0)GO TO 740
DA12=(S2*(-DSXA)-S1*(A23*DSZA-A33*DSYA))/D
DB12=(S2*(-DSXB)-S1*(A23*DSZB-A33*DSYB))/D
DE12=(S2*Q*VAE13-S1*(SZ*VAE23-SY*VAE33))/D
DG12=(S2*Q*VAG13-S1*(SZ*VAG23-SY*VAG33))/D
DR12=(S2*Q*VAR13-S1*(SZ*VAR23-SY*VAR33))/D
DDA12=SQRTE(DA12**2+DB12**2+DE12**2+DG12**2+DR12**2)
DA32=(S2*(-DSZA)-S1*(A13*DSYA-A23*DSXA))/D
DB32=(S2*(-DSZB)-S1*(A13*DSYB-A23*DSXB))/D
DE32=(S2*Q*VAE33-S1*(SY*VAE13-SX*VAE23))/D
DG32=(S2*Q*VAG33-S1*(SY*VAG13-SX*VAG23))/D
DR32=(S2*Q*VAR33-S1*(SY*VAR13-SX*VAR23))/D
DDA32=SQRTE(DA32**2+DB32**2+DE32**2+DG32**2+DR32**2)
GO TO 75J
740 DK12= S2*(Q*VAI13-DSXA)-S1*(SZ*VAI23+A23*DSZA-SY*VAI33-A33*DSYA)
DK32= S2*(Q*VAI33-DSZA)-S1*(SY*VAI13+A13*DSYA-SX*VAI23-A23*DSXA)
DL12= S2*(Q*VBI13-DSXB)-S1*(SZ*VBI23+A23*DSZB-SY*VBI33-A33*DSYB)
DL32= S2*(Q*VBI33-DSZB)-S1*(SY*VBI13+A13*DSYB-SX*VBI23-A23*DSYB)
DA12=DK12/D
DA32=DK32/D

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DB12=DL12/D
DB32=DL32/D
DG12=(S2*G*VAG13-S1*(SZ*VAG23-SY*VAG33))
DG32=(S2*Q*VAG33-S1*(SY*VAG13-SX*VAG23))/D
DR12=(S2*Q*VAR13-S1*(SZ*VAR23-SY*VAR33))
DR32=(S2*Q*VAG33-S1*(SY*VAG13-SX*VAG23))/D
DDA12=SQRTF(DA12**2+DB12**2+DR12**2+DG12**2)
DDA32=SQRTF(DA32**2+DB32**2+DG32**2+DR32**2)
750 DAA12(I)=DDA12
DAA32(I)=DDA32
PRINT 60
PRINT 76J,I,AA13(I),I,AA23(I),I,AA33(I)
760 FORMAT(3X,$A13($,I1,$)=$,F11.8,2X,$A23($,I1,$)=$,F11.8,
12X,$A33($,I1,$)=$,F11.8,/)
PRINT 77J,I,AA11(I),I,AA21(I),I,AA31(I)
770 FORMAT(3X,$A11($,I1,$)=$,F11.8,2X,$A21($,I1,$)=$,F11.8,
12X,$A31($,I1,$)=$,F11.8,/)
PRINT 78J,I,AA12(I),I,AA22(I),I,AA32(I)
780 FORMAT(3X,$A12($,I1,$)=$,F11.8,2X,$A22($,I1,$)=$,F11.8,
12X,$A32($,I1,$)=$,F11.8,/)
PRINT 79U,I,TH(I),I,TE(I),I,PH(I)
790 FORMAT(3X,$ROLL ANGLE TH($,I1,$)=$,F10.5,$ DEG.,PITCH ANGLE TE($,
I1,$)=$,F10.5,$ DEG., YAW ANGLE PH($,I1,$)=$,F10.5,$ DEG.$,/)
PRINT 60
PRINT 60
PRINT 82U,I,DAA12(I),I,DAA32(I)
820 FORMAT(3X,$STANDARD DEVIATION OF A12,A32 ARE,DA12($,I1,
1$)=$,F8.5,$, DA32($,I1,$)=$,F9.6,/)
C*****
LP=4
DT1=-A32
DT2=-A12
FAL=3.0*DDA32
FBT=3.0*DDA12
CALL PLOT(LP,GAMA,ISUN,SX,SY,SZ,S2,S3,DT1,DT2,FAL,FBT)
830 CONTINUE
GO TO 97U
C*****
840 PRINT 850
850 FORMAT(1X,$SUNLINE AND LOCAL VERTICAL ARE IN THE SAME LINE,NO
1FURTHER CALCULATION FOR A12,A22,A32,ETC.$,/)
IF(LE.NE.0)GO TO 870
IF(K.EQ.3)GO TO 940
KA1=K
KA2=K
GO TO 95U
860 PRINT 30U,AA13(1),AA23(1),AA33(1)
PRINT 31U,AA13(2),AA23(2),AA33(2)
PRINT 38U,TH(1),TE(1)
PRINT 39U,TH(2),TE(2)
PRINT 60
AW(8)=TH(1)
AW(9)=TE(1)
AW(11)=TH(2)

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      AW(12)=TE(2)
      IF(LE.EQ.0) GO TO 940
870  GO TO (880,900,920,920,920),LE
880  PRINT 890
890  FORMAT(3X,DEVIATION OF C23 IS TOO LARGE,HENCE THERE IS NO ERROR
      1CALCULATION FOR A13,A23,A33,/)
      GO TO 970
900  PRINT 910
910  FORMAT(3X,DAG13(1),DAG23(1) CAN NOT BE DETERMINED. NO FURTHER
      1CALCULATION FOR ERROR ELLIPSE,/)
      GO TO 970
920  PRINT 930
930  FORMAT(3X,DAG23(1),DAG23(2) CAN NOT BE DETERMINED. NO FURTHER
      1CALCULATION FOR ERROR ELLIPSE,/)
      GO TO 970
940  PRINT 480,DAA13(1),DAA13(2)
      PRINT 490,DAA23(1),DAA23(2)
      PRINT 60
      KA1=1
      KA2=2
950  DO 960 I=KA1,KA2
      CT1=-AA23(I)
      CT2=-AA13(I)
      EAL=3.0*DAA23(I)
      EBT=3.0*DAA13(I)
      CALL PLOT(LP,GAMA,ISUN,SX,SY,SZ,S2,S3,CT1,CT2,EAL,EBT)
960  CONTINUE
C*****
C*PART 4.....SUMMARY OF RUNS
C*****
970  MRUN=MRUN+1
      WRITE OUTPUT TAPE 2, 980,(AW(I),I=1,13)
980  FORMAT (1X,13,13(1X,F8.2),/)
      IOR=IO
      GO TO 20
990  DO 1000 J=1, MRUN
1000  BACKSPACE 2
      PRINT 1010,IOR
1010  FORMAT(4JX,.....SUMMARY OF RUNS FOR ORBIT NO.,F8.2,2X,.....)
      PRINT 60
      PRINT 1020
1020  FORMAT(2X,NO,3X,DATE,7X,HR,5X,MIN,7X,SEC,
      14X,ORBIT,5X,GAMA,4X,ROMC,7X,THI,5X,THETA,
      25X,PHI,5X,THI,5X,THETA,6X,PHI,/)
      DO 1030 JI=1, MRUN
      READ INPUT TAPE 2, 980,(AW(I),I=1,13)
      PRINT 980, JI,(AW(I),I=1,13)
1030  CONTINUE
      PRINT 60
      MRUN=0
      IOR=IO
      PRINT 620
      IF(IO.NE.0) GO TO 40
      END

```

```

SUBROUTINE PLOT (LP, GA, ISUN, SX, SY, SZ, S2, S3, CT1, CT2, EAL, EBT)
COMMON C23, IA(90,55), IB(90,55), R0MD, IDAY, ITIM
C
C THIS PROGRAM IS USING POLAR COORDINATES FOR PLOT
C AE AND BE ARE THE MAJOR AND MINOR AXIS OF ELLIPSE
C IA REPRESENTS THE PLOT FOR A13 AND A23.
C IB REPRESENTS THE PLOT FOR A12 AND A32.
C WHEN LP=0, CC=C23, WHERE C23 IS CALCULATE FROM MAIN PROGRAM
C WHEN LP=1, C23 IS LESS THAN -0.3733(PUT CC=-0.3733
C WHEN LP=2, C23 IS GREATER THAN 0.9277(PUT CC=0.9277)
C
SNG=SIN(GA)
CNG=COS(GA)
GAD=GA/0.01745329
CC=C23
RC=1.0-CC*CC
CCN=CC*CNG
SRC=SQRTF(RC)
AE=SRC*SNG
BE=SRC
CE=-CCN
IXC=(CC+1)*42+1.5
IYM=SRC*25+26.5
DO 10 I=1,90
DO 10 J=1,55
IA(I,J)=1H
IB(I,J)=1H
10 CONTINUE
C PLOT Y AXIS
DO 20 J=1,53
IA(43,J)=1H*
IB(43,J)=1H*
20 CONTINUE
C PLOT X-AXIS
DO 30 J=1,87,2
IA(J,26)=1H*
IB(J,26)=1H*
30 CONTINUE
C LABEL FOR A23 AXIS
IA(90,26)=1HA
IA(89,26)=1H2
IA(88,26)=1H3
IA(87,26)=1H+
C LABEL FOR A32 AXIS
IB(90,26)=1HA
IB(89,26)=1H3
IB(88,26)=1H2
IB(87,26)=1H+
DO 40 J=1,85,21
IA(J,26)=1H1
IB(J,26)=1H1
40 CONTINUE
C*****
C LABEL FOR THE CENTER O OF CELESTIAL SPHERE

```

```

      IA(43,26)=1H0
      IB(43,26)=1H0
C     AL REPRESENTS POLAR ANGLE
      AL=0
      RP=1.0
      CALL POINT(AL,RP,X1,Y1,IX1,IY1)
C     DAL IS THE INCREMENT OF ANGLE FOR POLAR COORDINATES METHOD
      DAL=5.0*U.01745329
      IF(LP.L1.3)GO TO 100
C*****
C     PLOT ERROR ELLIPSE WHEN LP IS GREATER THAN 3
      SNU=0.0
      CSU=1.0
50  CALL ELIPS(SNU,CSU,AL,EAL,EAT,CT1,CT2,XER,YER,IXR,IYR)
      IA(IX1,IY1)=1H*
      IB(IX1,IY1)=1H*
      IF(ABS(XER).GT.1.0)GO TO 60
      YEM=SQRTF(1.0-XER*XER)
      IF(YER.GT.YEM)GO TO 60
      IF((IXR.LT.0. OR.IXR.GT.85).OR.(IYR.LT.0. OR.IYR.GT.52))GO TO 60
      IA(IXR,IYR)=1HR
      IB(IXR,IYR)=1HR
60  AL=AL+DAL
      CALL POINT(AL,RP,X1,Y1,IX1,IY1)
      IF(AL.LE.0.3) GO TO 50
      CT1=-CT1
      CT2=-CT2
      IXER=(CT1+1)*42+1.5
      IYER=CT2*25+26.5
      IA(IXER,IYER)=1HR
      IB(IXER,IYER)=1HR
      PRINT 510
      IF(LP.EQ.4)GO TO 80
      PRINT 70,CT1,CT2,EAL,EAT
70  FORMAT(1X,ERROR ELLIPSE OF A23,A13: CENTER AT A23=#,F8.5,#, A13
1-#,F8.5,#, AXIS PARALLEL TO A23=#,F8.5,#, AXIS PARALLEL TO A13=#,
2F8.5)
      GO TO 60J
80  PRINT 90,CT1,CT2,EAL,EAT
90  FORMAT(1X,ERROR ELLIPSE OF A32,A12: CENTER AT A32=#,F8.5,#, A12
1-#,F8.5,#, AXIS PARALLEL TO A32=#,F8.5,#, AXIS PARALLEL TO A12=#,
2F8.5)
      GO TO 530
C*****
100 PRINT 110
110 FORMAT(20X,ERROR ELLIPSE OF A23,A13: CENTER AT A23=#,F8.5,#, A13
1-#,F8.5,#, AXIS PARALLEL TO A23=#,F8.5,#, AXIS PARALLEL TO A13=#,
2F8.5)
C     ISUN=0 MEANS NO INFORMATION OF SUN LINE DIRECTION
C     ISUN=1,2,3, MEANS INFORMATION OF SUN LINE IS AVAILABLE
      IF(ISUN.EQ.0) GO TO 360
C     INFORMATION FROM SUN SENSOR AVAILABL.PLOT THE REGION
      IXA=(S3+1)*42+1.5
      IXB=(S2+1)*42+1.5
      SR=SQRTF(ABS(1.0-S2*S2))

```

```

C      ST=SQRTF(ABSF(1.0-S3*S3))
C      AP AND BP ARE THE MAJOR AND MINOR AXIS OF ELLIPSE IN A13,A23,PL01.
      AP=ABSF(SZ*ST)
      BP=ST
C      AQ AND BQ ARE THE MAJOR AND MINOR AXIS OF ELLIPSE IN A12,A32 PL01.
      AQ=ABSF(SY*SR)
      BQ=SR
C      (EX,EY),(FX,FY) ARE CENTER OF EACH ELLIPSE.
      EX=-S3*SY
      EY=-S3*SX
      FX=-SZ*S2
      FY=-SX*S2
      RXY=SQRTF(SX*SX+SY*SY)
      RXZ=SQRTF(SX*SX+SZ*SZ)
C*****
      IF(SX .EQ. 0.0) GO TO 120
      SM=-SY/SX
      SN=-SZ/SX
      SK=S3/SX
      SL=S2/SX
      120 CALL TEST(SX,SY,SZ,IS)
      CALL TEST(SX,SZ,SY,LS)
      IF(LS .NE. 1) GO TO 130
      SNF=SX/RXZ
      CSF=SZ/RXZ
      FL=ASIN(SNF)
      130 GO TO (160,180,140,150),IS
C*****
      140 IY=IY1
      IX=IXA
      X3=S3
      Y3=Y1
      GO TO 190
C*****
      150 Y3=SM*X1+SK
      IY =Y3*25+26.5
      IX=IX1
      X3=X1
      GO TO 190
C*****
      160 SNE=SX/RXY
      CSE=SY/RXY
      EL=ASIN(SNE)
      170 ANGLE=AL+EL
      CALL ELIPS(SNE,CSE, ANGLE, AP,BP,EX,EY,X3,Y3,IX,IY)
      GO TO 190
C*****
C      CALCULATE THE COORDINATES OF CIRCLE
      180 CALL POINT(AL,ST,X3,Y3,IX,IY)
      190 IF((IX.LT.0. OR IX.GT.85).OR.(IY.LT.0. OR IY.GT.52)) GO TO 250
      IA(IX,IY)=IH=
      GO TO (200,200,230,230),IS
      200 A3=(S3-X3*SY-Y3*SX )/SZ
      PRINT 210,X3,Y3,A3

```

```

210 FORMAT(1X,#A23=#,F9.6,10X,#A13=#,F9.6,10X,#A33=#,F9.6)
    IF( A3 .LT. 0.) GO TO 220
    IA(IX,IY)=1HP
    GO TO 250
220 IA(IX,IY)=1HN
    GO TO 250
230 PRINT 240,X3,Y3
240 FORMAT(11X,#A23=#,F9.6, 8X,#A13=#,F9.6, 8X,#SIGN OF A33 UNKNOWN#)
250 GO TO (260,270,280,290),LS
C*****CALCULATE A12,A32 REGION BY SUN SENSOR INFORMATION
260 ANGLF=AL+FL
    CALL ELIPS(SNF,CSF, ANGLF, A0,B0,FX,FY,X2,Y2,IX,IY)
    GO TO 300
270 CALL POINT(AL,SR,X2,Y2,IX,IY)
    GO TO 300
280 IX=IXB
    IY=IY1
    Y2=Y1
    X2=S2
    GO TO 300
290 Y2=SN*X1+SL
    IY=Y2*25+26.5
    X2=X1
    IX=IX1
300 IF((IX.LT.0. OR.IX.GT.85).OR.(IY.LT.0. OR.IY.GT.52)) GO TO 360
    IB(IX,IY)=1H=
C*****
    GO TO(310,310,340,340),LS
310 A2=(S2-X2*S2-Y2*SX)/SY
    PRINT 320,X2,Y2,A2
320 FORMAT(1X,#A32=#,F9.6, 8X,#A12=#,F9.6,8X,#A22=#,F9.6,/)
    IF( A2 .LT. 0.) GO TO 330
    IB(IX,IY)=1HP
    GO TO 360
330 IB(IX,IY)=1HN
    GO TO 360
340 PRINT 350,X2,Y2
350 FORMAT(1X,#A32=#,F9.6,10X,#A12=#,F9.6,9X,#SIGN OF A22 UNKNOWN#)
C*****
C    INFORMATION FROM EARTH SENSOR ONLY
360 IF(GA .NE. 0.0) GO TO 370
    NS=1
    GO TO 380
370 NS=2
C    PLOT THE CELESTIAL SPHERE.
380 IA(IX1,IY1)=1H*
    IB(IX1,IY1)=1H*
    GO TO ( 390,410),NS
C    IF GAMA IS ZERO, THE REGION IS A STRAIGHT LINE, A23=CC
C    PLOT THIS LINE
390 IX=IXC
    IF(IY1.GT.1YM)GO TO 470
    IY=IY1
    X=CC

```

```

      Y=Y1
      PRINT 400,X,Y
400  FORMAT(5X, #A23=#,F9.6, 8X, #A13=#,F9.6,6X, #SIGN OF A33 UNKNOWN#)
      GO TO 430
C*****
C      IF IT IS ANOTHER ELLIPSE (BY EARTH SENSOR INFORMATION)
C      CALCULATE THE COORDINATES OF THIS ELLIPSE
410  CALL ELIPS(0.0,1.0,AL,AE,BE,CE,0.0,X,Y,IX,IY)
      AZ3=(CC-X*CN3)/SNG
      PRINT 420, X,Y,AZ3
420  FORMAT(65X, #A23=#,F9.6, 8X, #A13=#,F9.6,6X, #A33=#,F9.6)
430  IF((IX.LT.0. OR.IX.GT.85).OR.(IY.LT.0. OR.IY.GT.52))GO TO 470
      IA(IX,IY)=1H1
      IF(GA.EQ.0.0) GO TO 470
      IF(LP.EQ.0) GO TO 450
      IF(AZ3.GE.0) GO TO 440
      GO TO 470
440  IA(IX,IY)=1H.
      GO TO 470
450  IF(AZ3.GT.0)GO TO 460
      IA(IX,IY)=1H-
      GO TO 470
460  IA(IX,IY)=1H+
C*****
470  AL=AL+DAL
      RP=1
      CALL POINT(AL,RP,X1,Y1,IX1,IY1)
      IF(AL.GT.6.3) GO TO 480
      IF(ISUN.EQ.0) GO TO 380
      GO TO (170,180,140,150),IS
480  IF(ISUN.EQ.0) GO TO 570
      IF(IS.NE.1) GO TO 490
      IX=(-EX+1)*42+1.5
      IY=-EY*25+26.5
      IA(IX,IY)=1HE
490  IF(LS.NE.1) GO TO 500
      IX=(-FX+1)*42+1.5
      IY=-FY*25+26.5
      IB(IX,IY)=1HF
500  PRINT 510
510  FORMAT(JH1)
      PRINT 520,ISUN,C11,C12,EAL,EB1
520  FORMAT(1X, #ISUN=#,I2, #, ALFA=#,F8.3, # DEG, B11A=#,F8.3, # DEG, AZ=#,
1,F8.3, # DEG, CL=#,F8.3, # DEG, P FOR+A22(OR+A33),N FOR-A22(OR-A33)#)
530  DO 550 J=1,52
      PRINT 540,(IB(I,J),I=90,1,-1)
540  FORMAT(2JA,1J0A1)
550  CONTINUE
      PRINT 560, IDAY, I1IM
560  FORMAT(64X, #A12+#,35X, I4, #, #, I6, /)
      IF(LP .GE. 3) GO TO 690
      GO TO 590
570  PRINT 510
      PRINT 580

```



```

580 FORMAT(1X,$N0 SUN SENSOR INFORMATION,N0 PLOT FOR A12,A32$)
590 PRINT 510
    IF(LP.GT.0) GO TO 610
    PRINT 600,GAD,CC,R0MD
600 FORMAT(1X,$GAMA=$,F6.2,$ DEG, C23=$,F7.4,$ R0M=$,F7.3,$ DEG, IF
1GAMA NOT 0, + FOR +A33,- FOR -A33; IF GAMA=0, I FOR BOTH$)
    GO TO 660
010 GO TO (020,040),LP
620 PRINT 630,GAD,CC
630 FORMAT(1X,$GAMA=$,F6.2,$ DEG, R0M=0, C23<=$,F6.3,$, . IS FOR +A33
1I FOR - A33 IF GAMA NOT 0; WHEN GAMA=0, I IS FOR EITHER$)
    GO TO 660
040 PRINT 650,GAD,CC
650 FORMAT(1X,$GAMA=$,F6.2,$ DEG, R0M=180 DEG,C23>=$,F6.3,$, . FOR +A33
1I FOR -A33 IF GAMA NOT 0; WHEN GAMA=0, I FOR EITHER$)
660 DO 670 J=1,52
    PRINT 540,(IA(I,J),I=90,1,-1)
670 CONTINUE
    PRINT 680,I0AY,ITIM
680 FORMAT(64X,$A13+$,35X,14,$, $,16,/)
690 RETURN
    END
    SUBROUTINE TEST(TX,TY,TZ,JS)

```

```

C
C   CONDITION FOR ELLIPSE, CIRCLE, OR STRAIGHT LINE IN PLOT
C       (1) WHEN JS =1, IT IS AN ELLIPSE
C       (2) WHEN JS=2, IT IS A CIRCLE
C       (3) WHEN JS =3, IT IS A STRAIGHT LINE PARALLEL TO Y AXIS
C       (4) WHEN JS=4, IT IS A STRAIGHT LINE WITH SLOPE = TY/IX.
C

```

```

    ATZ=ABS (TZ)
    IF(ATZ.NE.0.0) GO TO 20
    IF(TX.NE.0.0) GO TO 10
    JS=3
    GO TO 40
10 JS=4
    GO TO 40
20 IF (ATZ.EQ.1) GO TO 30
    JS=1
    GO TO 40
30 JS=2
40 RETURN
    END

```

```

SUBROUTINE ELIPS(SNU,CSU,ANGL,A,B,CX,CY,XE,YE,IX,IY)
C
C   CALCULATE THE COORDINATES OF ELLIPSE.
C   A AND B ARE THE HALF LENGTH OF MAJOR AND MINOR AXIS
C   ANGL REPRESENTS CURRENT POLAR ANGLE
C   (CX,CY) REPRESENTS THE CENTER OF ELLIPSE
C   (XE,YE) ARE CURRENT POINT REPRESENTED AS (IX,IY) FOR PLOT
C
    SAN=SIN(ANGL)
    CAN=COS(ANGL)
    XEX=A*CAN

```

```

YEY=B*SAN
XP=XEX*CSU-YEY*SNU
YP=XEX*SNU+YEY*CSU
XE=XP-CX
YE=YP-CY
IX=(XE+1)*42+1.5
IY=(YE+1)*25+26.5
RETURN
END
SUBROUTINE POINT(AL,RP,X,Y,IX,IY)

```

```

C
C   AL REPRESENTS POLAR ANGLE
C   (X,Y) REPRESENTS CURRENT POINT OF A CIRCLE.
C   RR REPRESENTS RADIUS
C

```

```

SNL=SIN(AL)
CSL=COS(AL)
X=RP*CSL
Y=RP*SNL
IX=(X+1)*42+1.5
IY=(Y+1)*25+26.5
RETURN
END

```

APPENDIX C SAMPLE OF OUTPUT

DAY = 257, HR = 23, MIN = 51, SEC = 48, ORBIT = 556, H = 267.1000

\*\*\*\*\*

EARTH SENSOR INFORMATION GIVES:

GAMA(IN DEGREE)=-23.570000, ET(IN DEGREE)= -5.050000, ROM(IN DEGREE)= 56.000000

\*\*\*\*\*

ISUN= 2, SUN SENSOR INFORMATION GIVES:

AZIMUTH= -43.00000 DEG, ELEVATION= -39.90000 DEG.

S11= .623941, S21= -.521696, S31= -.581834, ALFA(IN DEGREE)= 95.080000, BETA(IN DEGREE)= 1.185000

THE DIRECTION COSINES OF SUN LINE W.R.T. SPACECRAFT AXES ARE -.97492669, .10801531, .19455232

THE DIRECTION COSINES OF SUN LINE W.R.T. GEOCENTRIC AXIS: -.99585902, .02068067, .08852740

\*\*\*\*\*

\*\*\*\*\*

A13(1)= -.08797894, A23(1)= -.42776804, A33(1)= -.89959668

A13(2)= .08797894, A23(2)= .36844860, A33(2)= .92547572

ROLL ANGLE TH(1)= -25.32600 DEG., PITCH ANGLE TE(1)= 174.41437DEG.

ROLL ANGLE TH(2)= 21.61997 DEG., PITCH ANGLE TE(2)= -5.43042DEG.

\*\*\*\*\*

\*\*\*\*\*

DEVIATION IN A13, DA13(1)= .01738, DA13(2)= .01738

DEVIATION IN A23, DA23(1)= .01506, DA23(2)= .01549

DEVIATION IN A33, DA33(1)= .00729, DA33(2)= .00632

\*\*\*\*\*

\*\*\*\*\*

DIRECTION COSINES OF LOCAL VERTICAL ARE: .08797894 .36844860 .92547572

A13(2) = .08797894 A23(2) = .36844860 A33(2) = .925475,2

A11(2) = .98696566 A21(2) = -.05651465 A31(2) = -.120727,9

A12(2) = .00790463 A22(2) = .92438405 A32(2) = -.367739,1

ROLL ANGLE TH(2) = 21.61997 DEG., PITCH ANGLE TE(2) = -5.43042 DEG., YAW ANGLE PH(2) = 3.48524 DEG.

\*\*\*\*\*

\*\*\*\*\*

STANDARD DEVIATION OF A12, A32 ARE, DA12(2) = .01936, DA32(2) = .015313

ERROR ELLIPSE OF A23,A13: CENTER AT A23= .36895, A13= .08798, AXIS PARALLEL TO A23= .04647, AXIS PARALLEL TO A13= .05213

A23+ \* \* \* \* \* I \* \* \* \* \* D \* \* \* \* \* 1 \* \* \* \* \*

RRRR  
R RR  
RRRR  
R

ERROR ELLIPSE OF A32,A12: CENTER AT A32= 36774, A12= 00790, AXIS PARALLEL TO A32= 04594, AXIS PARALLEL TO A12= 05807

A32+

RHRH

RHRH

RR

.....SUMMARY OF RUNS FOR ORBIT NO. 556.00 . . .

\*\*\*\*\*

NO.	DATE	HR.	MIN	SEC	ORBIT	GAMA	RQMD	THI	THETA	PHI	THI	THETA	PHI
1	257.00	23.00	49.00	10.00	556.00	-22.82	54.40	.00	.00	.00	-25.96	-176.34	-2.37
2	257.00	23.00	51.00	48.00	556.00	-23.57	56.00	.00	.00	.00	21.62	-5.43	2.47
3	257.00	23.00	54.00	6.00	556.00	-22.82	57.60	.00	.00	.00	22.11	-7.22	3.32
4	257.00	23.00	56.00	17.00	556.00	-20.59	56.00	-22.29	174.54	.00	18.73	-5.33	.00
5	257.00	23.00	58.00	12.00	556.00	-18.36	57.60	-18.72	175.53	.00	17.90	-4.45	.00

\*\*\*\*\*